High Energy Experiments: A Smashing good time



Meet the Group: CMS and D-Zero Projects

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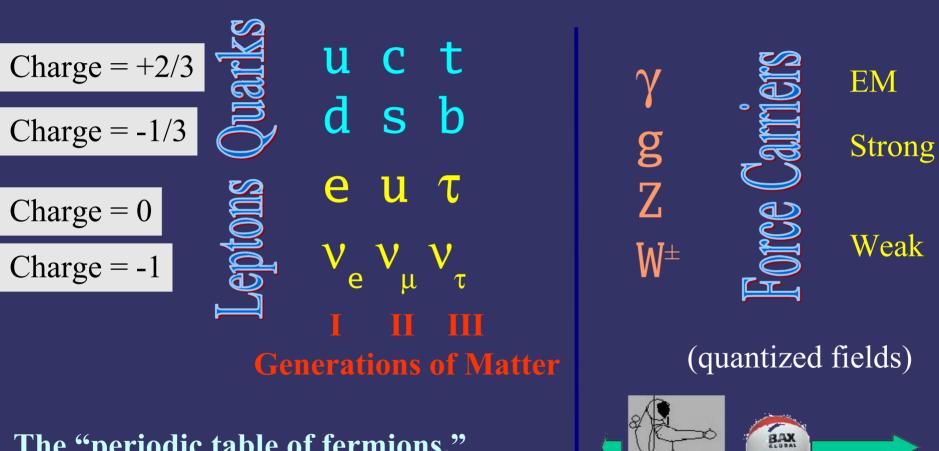
The basic questions

What are the basic building blocks of everything?

What are the forces of nature?

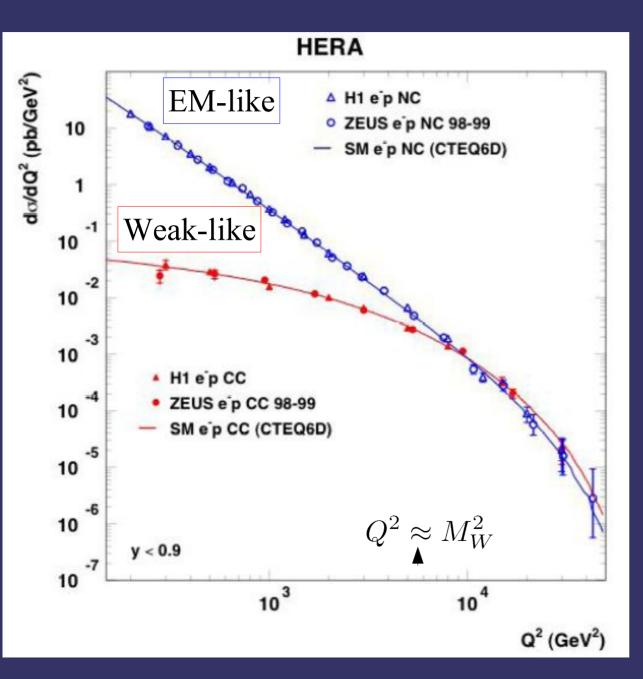
Some things we know about matter:

Elementary Particles



The "periodic table of fermions."

Some things we know about matter:



Electroweak Unification

EM/Weak interactions unified at large energy/momentum transfer

Some things we know about matter:

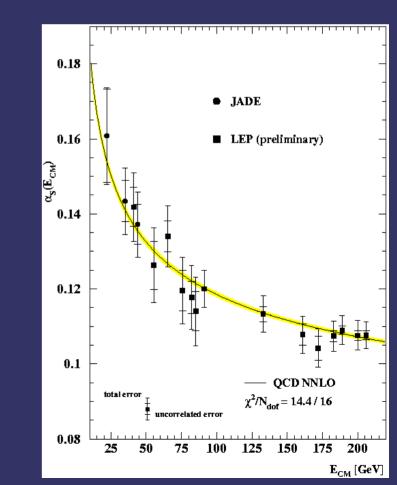
Quantum Chromodynamics

Gauge theory (like electromagnetism) describes fermions (quarks) which carry an SU(3) charge (color) and interact through the exchange of vector bosons (gluons)

Interesting features:

- gluons themselves have color
- interactions are strong
- coupling constant runs rapidly becomes weak at momentum transfers above a few GeV

In a more general theory (GUT), expect unification w/ electroweak force

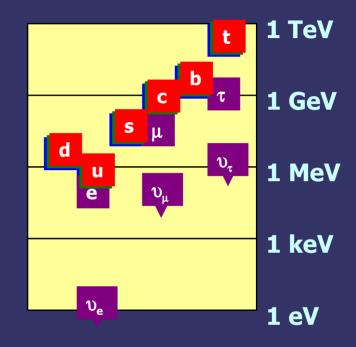


Some things we don't know about matter:

Masses

Second and third generations of quarks and leptons are much more massive than first

Origin of mass difference for bosons in EW force?



mass = 80.4 GeV photon mass = 0 Electroweak asymmetry

>12 orders of magnitude in mass!

Some ideas about matter:

The Higgs Mechanism

In the Standard Model

- (Glashow, Weinberg, Salam, 't Hooft, Veltmann)
 - "electroweak symmetry breaking" through introduction of a scalar field ϕ \rightarrow masses of W and Z
- Higgs field permeates space with a finite vacuum expectation value
 - cosmological implications! (inflation)
- If φ also couples to fermions \rightarrow generates fermion masses

An appealing picture: is it correct?

- One clear and testable prediction: there exists a neutral scalar particle which is an excitation of the Higgs field
- All its properties (production and decay rates, couplings) are fixed except its own mass

Highest priority of worldwide high energy physics program: find it!

W's, Z's, Top's, and Higgs's (oh my!):

- Fundamental parameters of Standard Model (SM)
- Affect predictions of SM via radiative corrections:
 - BB mixing

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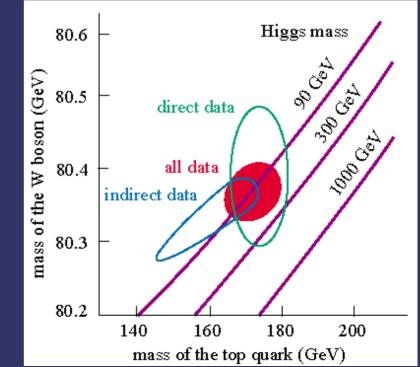
W

– W and Z mass

$$\frac{W}{-} - \frac{W}{-} - \frac{W}{-} - \frac{W}{-}$$

– measurements of M_W , m_t constrain M_H

- Large mass of top quark
 - may provide clues about electroweak symmetry breaking



Theoretical problems
Mass scale for M_{EW}?
how to achieve grand unification?
how to include gravity?
what explains proliferation of quark and lepton types and determines their mixings?

Experimental problems SM fit to electro-weak data has probability of 4.4% what is dark matter? new type of matter? – can be produced at the LHC! "dark" because of undiscovered properties of space-time? – can be probed at LHC!

More general theories make predictions that can be tested at the Tevatron & LHC

Beyond the Higgs:

Standard Model works well for observed phonomena and would be completed by the discovery of the Higgs, but the Higgs may be the first window on to a new domain of physics at the electroweak scale

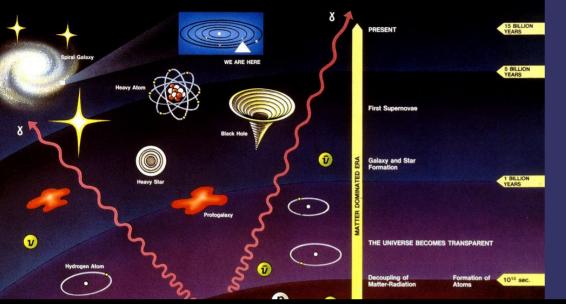
This Higgs is unlike any other particle in the SM (no other elementary scalars)

SM Higgs would have a mass unstable to radiative corrections (quantum effects): m_H would become very large $m_H \sim 10^{15}$ GeV, unless parameters fine tuned at the level of 1 part in 10^{26}

the patterns of the fundamental particles suggest a deeper structure the SM is a low energy approximation to some more general theory

Theoretically the most attractive option is supersymmetry

History of the Universe

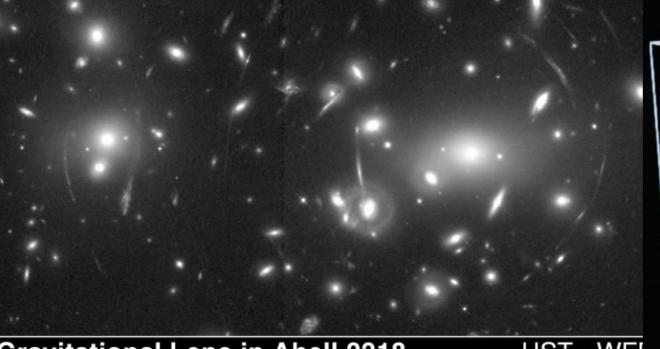


(Now...)

Now (15 billion years)

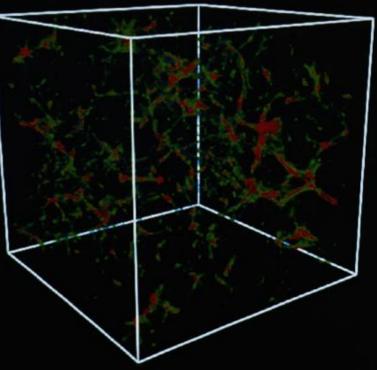
Large scale galactic clustering, mass deficit in universe, lensing

where/what is dark matter?



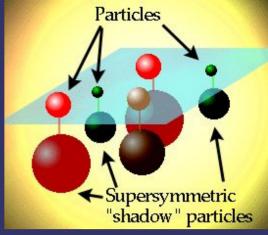
Gravitational Lens in Abell 2218 PF95-14 · ST Scl OPO · April 5, 1995 · W. Couch (UNSW), NASA

HST • WFI

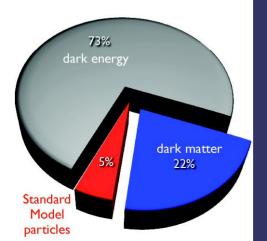


Supersymmetry

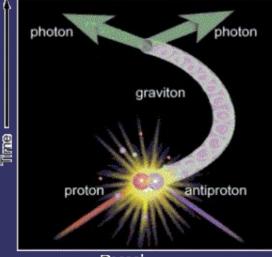
- Introduce a symmetry between bosons and fermions:
 - more massive super-partners for all particles
 - Allows a fundamental scalar (the Higgs) at low mass
 - cancels the divergences in $m_{\rm H}$



- closely approximates the standard model at low energies
- allows unification of forces with common couplings at higher energies
- provides a path to the incorporation of gravity and string theory: Local Supersymmetry = Supergravity
- lightest neutral superpartner (neutralino, etc) is massive, weakly interacting + stable → cosmic dark matter candidate!



- Connections w/ gravity: allows TEV scale affects for processes probing extra spacial dimensions
 - Flexible framework & calculable!



But, SUSY or other extensions to the Standard Model aren't obvious

example of various analysis channels								CDF Run II preliminary (927 pb ⁻¹)			
Final State	Data	Backgrou	und	Pinal State	Data	Backgro		Final State	Data	2011 C 100 C	round
$3)\tau +$	71			2e+j	13	9.8 ±		$o + \gamma p$	141	144.2	
5j	1661			$2e + e_{-}$	12	4.8 ±		4:+ 12-70	5-4	42.6	
$2j\tau +$	233			2e+	23	$36.1 \pm$		$a + \mu + p$	13	10.9	
hetj	2207			$2b \Sigma p_T > 400 \text{ GeV}$	327	335.8 ±		0+µ-	153	127.6	
$3J \Sigma p_T < 400 GeV$		37294.6 ± 5 1751.6 ± 4		$2b \Sigma p_T < 400 \text{ GeV}$	28	173.1 ± 33.5 ±		ei+1	386880		\pm 5031.8 \pm 2.9
e+3jp be+2j	1954	695.3 ± 1		263j $\Sigma p_T < 400 \text{ GeV}$ 262j $\Sigma p_T > 400 \text{ GeV}$	355	$33.5 \pm 326.3 \pm$		$e+j2\gamma$ $e+j\tau+$	79		± 2.9
$3jp \Sigma p_T > 400 \text{ GeV}$		967.5 ± 3		$2b2j \Sigma p_T > 400 \text{ GeV}$ $2b2j \Sigma p_T < 400 \text{ GeV}$	56	80.2 ±		e+j7-	162	148.8	
e+a+	26			2b2j 2pp < 400 Gev	16	15.4 +		e+i#		57391.7	
e+->	636	851.9 ± 1		Sha	37	31.7 ±		+j~p	52	76.9	
e+3j	100 C	27281.0 ± 4		$261.25p_{T} > 400 GeV$	410	393.8 ±		-+14-p	22	13.1	
b5j	131	95 ± 4		$2bj \Sigma_{PT} < 400 GeV$	161	$195.8 \pm$		e+ja-	28	26.8	
127+	50			$2bj \neq \Sigma p_T > 400 \text{ GeV}$	28	23.2 ±		e+e-41	103	113.5	
j++	7.4			2biz	25	24.7 +		e+e-3j	456		+ 14.6
$h \neq \Sigma_{PT} > 400 GeV$	10			21	1.5	12.8 +		0 + 0. 2jp	30		+ + 4
0+17	280	309.4 ± 2	1	2be+2]	30	30.5 ±	2.5	e+e-21	21.49	\$152	1 40.1
e+jp7-	29	14.2 ± 1	.8	2be+j	28	$29.1 \pm$	2.8	e+e-++	1.4	11.1	± 2
$2j \Sigma p_T < 400 GeV$	96502	92437.3 ± 1	354.5	2bc+	48	45.2 ±		e+e-p	491	487.9	
be+3)	356	298.6 ± 7	7.7	T+T-	498	428.5 ±	22.7	0+0-3	127	132.3	± 4.2
81	11	6.1 ± 2	5		177	204.4 +	5.4	e+e-i	10726	10669.3	+ 123.5
73	57	95.0 ± 4		76	1052	1945.8 ±	77.1	n+n-5¢	157	1-6-6	+ 11.2
6)	335	298.4 ± 1	4.7	$\mu + \tau +$	18	19.8 ±	2.3	0+0-3Y	26	45.6	± 4.7
$4j \Sigma p_T > 400 GeV$	39665	40898.8 ± 6	49.2	µ++	151	$179.1 \pm$	4.7	e+e-	58344	58575.6	± 603.9
$4j \Sigma p_T < 400 GeV$	8241	8403.7 ± 1	44.7	$\mu + p$	321351	320500 ±	3475.5	ъбј	24	15.5	± 2.3
4127	3.8	57.5 ± 1	1	11+107-	22	$25.8 \pm$	2.7	b4j $\Sigma_{PT} > 400 \text{ GeV}$	1.3	9.2	+ 1.8
45+ +	20	86.0 ± 2	24	1++->	260	285 5 ±	5.0	hdj $\Sigma_{PT} < 400 \mathrm{GeV}$	46.4	400.2	+ 19.4
$4j \neq 2ip \gamma > 400 \text{ GeV}$	010	525.2 ± 3	5-6.0	$\mu + \gamma p$	269	282.2 ±	0.0	b3j $\Sigma p_T > 400 \mathrm{GeV}$	0304	0280	± 72.4
4370	28	53.8 ± 1	1.1	4+ + 1+ - P	49	61.4 ±	3.5	b3j $\Sigma p_{T} < 400~{\rm GeV}$	1639	1558.9	± 24.1
4j-y	3693	3827.2 ± 1	12.1	$\mu + \mu - \gamma$	32	$29.9 \pm$	2.6	$b3j \neq \Sigma p_T > 400 GeV$	111	116.8	± 11.2
4jµ+	576	568.2 ± 2	26.1	44 + 44-	10648	$10845.6 \pm$	96	b3j ->	182	194.1	+ 8.8
4.5 pt + pt	232	224.7 ± 8	4 . Eb	324	2196	$2200.3 \pm$	35.2	$b 3j \mu + p$	37	3.4.1	+ 2
- ×4 + ×4 6.4	1.7	20.1 ± 2	0.0	j2-12	38	27.3 ±	3.2	b3jp+	47	52.2	± 3
37	13	24.2 ± 3	\$	++t	563	$585.7 \pm$	10.2	b2γ	1.5	14.6	± 2.1
$3j \Sigma p_T > 400 GeV$	75894	75939.2 ± 1	043.9	$j \not \equiv \Sigma p_T > 400 \mathrm{GeV}$	4183	$4209.1 \pm$	56.1	b2j $\Sigma p_T > 400 \mathrm{GeV}$	8812	8576.2	± 97.9
3j27	145	178.1 ± 7	r. a	37	49052	$48743~\pm$	546.3	b2) $\Sigma p_T < 400 \mathrm{GeV}$	4691	4646.2	± 57.7
$3j \neq \Sigma_{T'T} < 400 \mathrm{GeV}$	20			3~++	106	$104 \pm$	4.1	b2ig $\Sigma p_T > 400 { m GeV}$	198	209.2	\pm s s
3377+	1.9	11 ± 3		371 10	913	965 9 ±	41.5	52J-7	420	425.1	1.13.1
3) 7 P	83			344+		$34026.7 \pm$		$p \leq (\mu + b)$	46	40.1	
Bjr	11424	11506.4 ± 1		jµ+≠-	29	$37.5 \pm$		b2jμ+	56		± 3.4
$3)\mu + p$	1114			jµ++g++	10	9.6 ±		$b\tau +$	19		± 2.2
$35\mu + \mu -$	61	84.5 ± 9		$3 \mu + i \epsilon$		$46316.4 \pm$		ba	976	1034.8	
3.j	2132			31 + #	7.8	69.8 ±		To syste	18	16.7	
$3b$ $\Sigma p_T > 400 GeV$				344.4.7	70	98.4 ±		bp+	303	263.5	
27+	316			344 + 44-	1977	$2093.3 \pm$		$b\mu + p$	204	218.1	
27\$	161	176 ± 9		e+4j	7144	6661.9 ±		bj $\Sigma p_T > 400 \mathrm{GeV}$	9060	9275.7	
27	8482			e+4j#	403	$363 \pm$		bj $\Sigma p_T < 400 \mathrm{GeV}$	7236	7030.8	
Visitive the		92789.5 ± 1		+3j7-	11	7.6 ±		P15-2	1.3	17.6	
2527	645			e+3j-	27	21.7 ±		bj++	1.9	12.9	
2)7+7-	1.5			$e+2\gamma$	47	74-5 ±		$bjp \Sigma p_T > 400 \text{ GeV}$	5.3		± 19.9
$2j\not \in \Sigma_{PT} > 400 \text{ GeV}$				e+21		122457 ±			937	989.4	
$2j \neq \Sigma p_T < 400 \text{ GeV}$				e+2j+-	53			biy≱	34	30.5	
237		33259.9 ± 3		e+2j7+	20			pin+w	104		
Sior+	48			+2jpi + 2jpi		12130.1 ± 65.9 ±		bj#+	173	52.2	
2)7¢	-4-0-3			0+2j7	101			be+3j# be+3i#	68		
2)µ+p	7287			e+τ-	609	555.9 ±		be+2jø	87		± 3.3
$2j\mu + \gamma p$	13			e+++	225	211.2 ±		be+p	330	347.2 176.6	
2jµ+7	41			e+#	476424	479572 ±			211	34.6	
254 1 4 -	9513			o i pir-	20	35 ± 18.7 ±		he i e j	02		± 3.1
91 j ps +	9013	2005 0 ± 1	00.0	$\mathbf{c} + \mathbf{p} \mathbf{r} + \mathbf{r}$	20	10-X T		pot-o-	0.3	-00	- O.Y.

Numerous specific models proposed

Many tunable parameters affect phenomenology w/o breaking SM as observed.

Majority of proposed models must be wrong, but this is a good thing...

Whatever is out there (and it's ~5x more plentiful than normal matter) will be a surprise!

How do we see any of this?

- Analyze states produced in proton-(anti)proton collisions at high energy
- $E=Mc^2$, so Big E = Large reach in creating new matter states
- Study dynamics of collisions How do the forces work...
- Only by understanding the Standard Model precisely can we hope to find new physics in the dynamics of our collisions
- Study massive states What kinds of things can exist, what are there properties...
- Precise knowledge of W, top properties are central to understanding where the Higgs may be...
 - Collider physics = precision studies + direct access to new physics Literally 100's of thesis topics

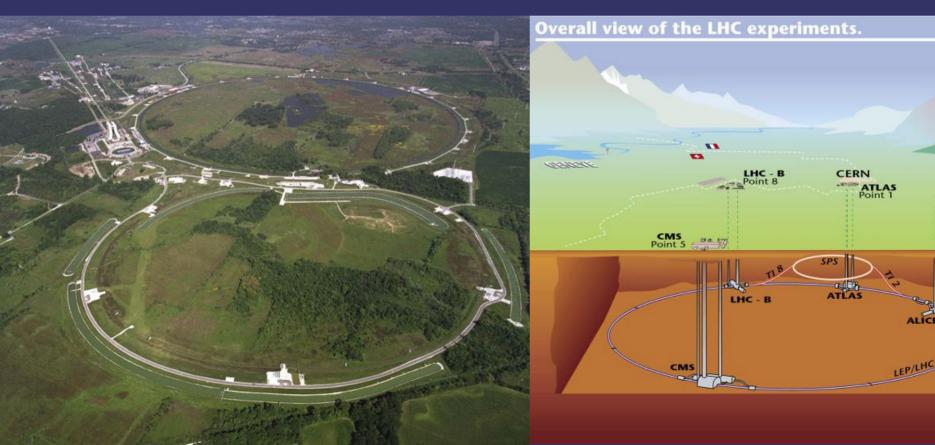
The Labs

Fermilab (Near Chicago) The Tevatron proton-antiproton collider at c.o.m. Energy = ~2TeV

6.3 km circumference

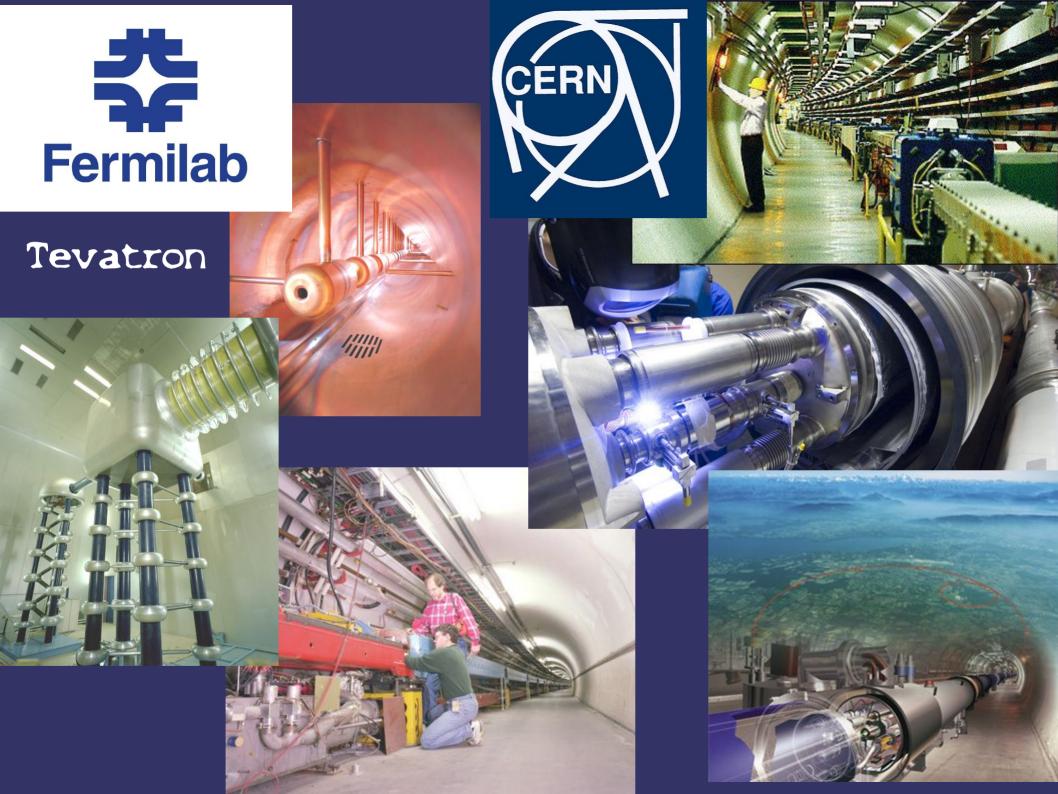
CERN (Geneva, Switzerland) The LHC proton-proton collider at c.o.m. Energy = 14TeV

27 km circumference



ALICE

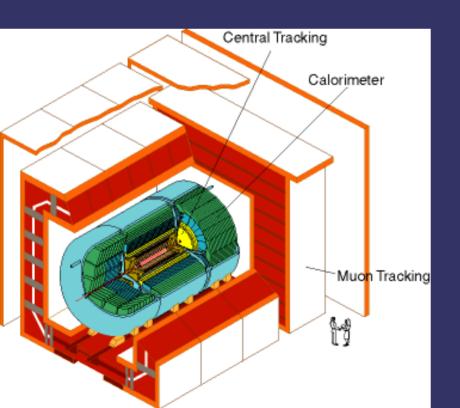
Point 2



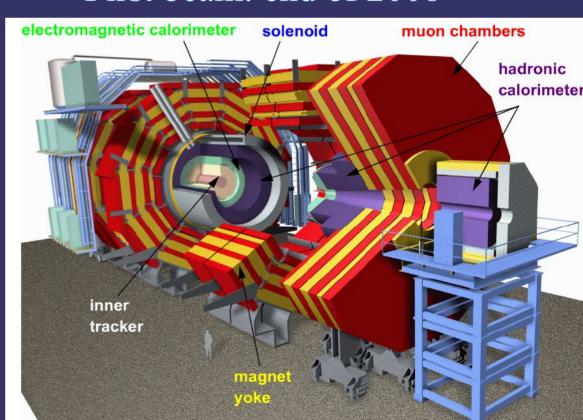
The Experiments

DØ - Fermilab Weighs 5000 tons ~10⁶ channels of information Inspects ~3-30 x10⁶ collisions/sec.

Running now



CMS - CERN Weighs 12,500 tons ~10⁷ channels of information Inspects ~40-1000 x10⁶ collisions/sec. First beam: end of 2008



DØ

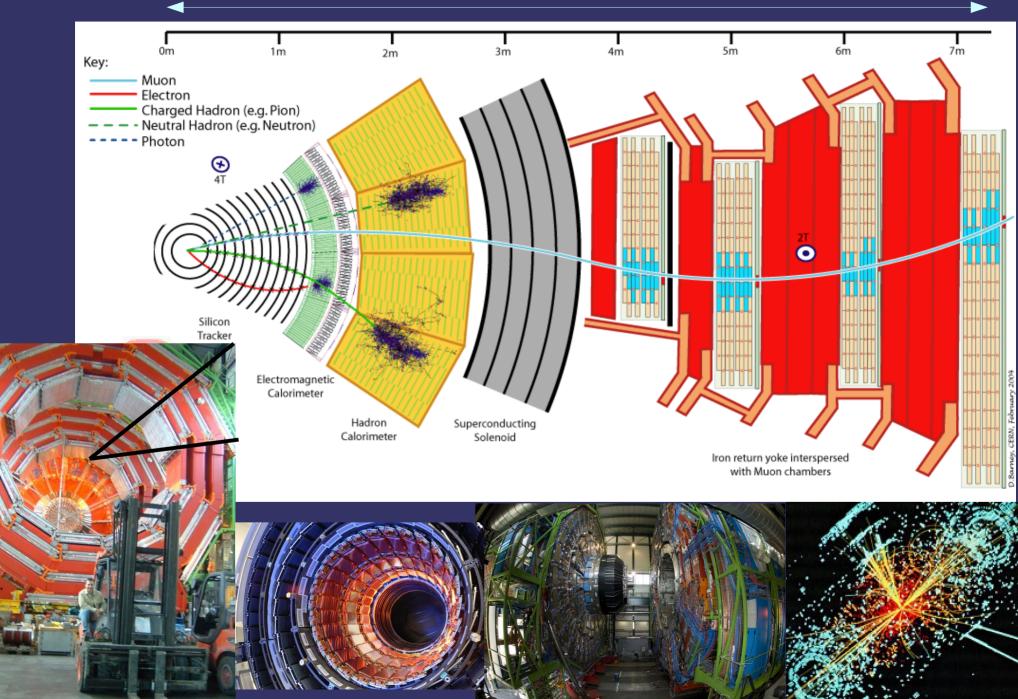
The Calorimeter

The full cube, w/ muon chambers



Slice of CMS

7m





Accumulate world's largest data sets at highest energies before LHC era

massive physics menu includes:

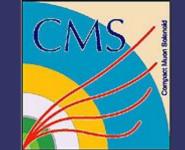
* Precision top physics (first ever)
* Precision EW physics (order of magnitude increase W statistics)*

★ Searches: Higgs* and for new physics (explore much of SUSY phase space, extradimensions, exotic matter states)

* QCD and proton hadronic structure (new levels of precision, smallest distance scales yet)*

* Heavy flavor physics (Precision measurements in heavy quark sector, relationship between generations, matter antimatter asymmetry,...

The best HEP program happening!Dukes, Hirosky* general group interests at present



LHC: The New Frontier

massive physics menu includes:

★ Copious top production

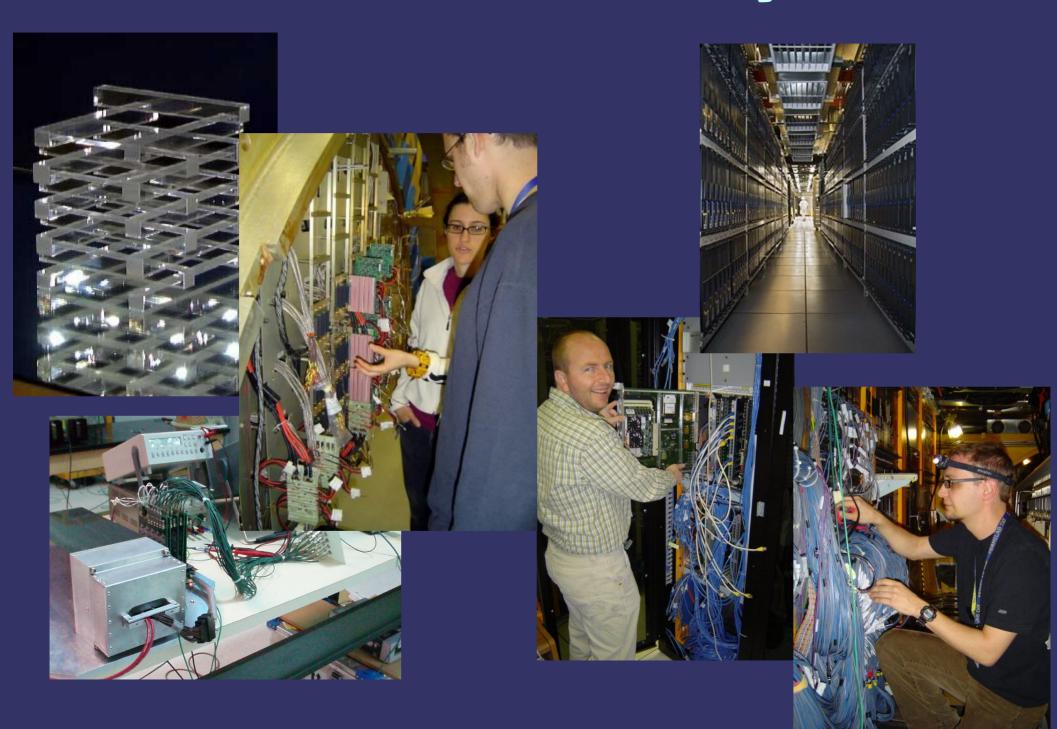
★ Test of EW physics to "unitarity limit"

- Copious Higgs production, test SM vs.
 SUSY Higgs
- Direct observations of SUSY states* or elimination of the lifetime work of many theorists :)
- * Order of magnitude increase in reach to physics at smallest distance scales*
- * Open to weirdness: black hole production, extremely massive exotic states, new types of strong interactions, extra-dimensions*....

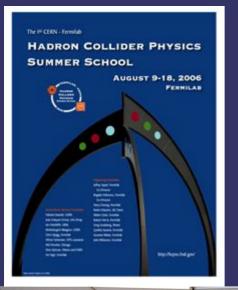
A new era for HEP is only a couple 1 year away!

Conetti, Cox, Hirosky

Various Technical Projects







Workshops and Physics Schools









Around CERN





CERN Surroundings



Where do HEP students go?

HEP students and postdocs go many places after completing their research, some (very few really) examples:

Industry

- Industrial research and instrumentation design
- Wireless technology and network infrastructure
- IT Consulting
- Financial analysis, modeling
- Design of medical treatment devices
- Non-Industry
 - Private and public "think tanks"
 - National research laboratories (not only HEP)
 - Law, Media

Academia

- 4-year undergraduate colleges
- Tier-1 research universities

But, you should only choose this or any other research area because you are interested in the physics!

no passion, no progress!

What can you learn along w/ physics?

Lots:

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 detector technologies, typical HEP experiments can easily employ many varieties of particle detection – enormous amount of practical physics in development of detector systems

 high performance data readout systems, electronics (HEP detectors must typically process 10's of TB of data each second)

• high performance computing:

• need to cull above data rate to manageable levels in real time

• handle data sets at many PetaByte level

sophisticated data analysis techniques, statistical reasoning, multivariate approaches for problems -> extracting maximal information from data
working with engineers and detector/accelerator physics experts to bring experiments on-line

experience w/ detailed simulations of detector systems and physics processes
working in a large expert physics community, amazing access to expertise in world wide community + opportunities to contribute

Interested in Experimental HEP?

Expected opportunities for graduate students this year:

<u>experiment</u>	<u>~advisor</u>	<u>colleagues</u>
CMS	Hirosky	(w/ Cox and Conetti)
DØ/NOvA	Dukes	(w/ Hirosky)

To learn more (start by doing some homework):

- seminars, colloquia, ...
- Symmetry Magazine: http://www.symmetrymagazine.org
- CERN Courier: http://cerncourier.com
- Femilab Today: http://www.fnal.gov/pub/today/

Contact me or another group member if you have questions about the group and our projects.

Interested in signing up? I'll be happy to interview interested students after the Fall semester.