





## Introduction



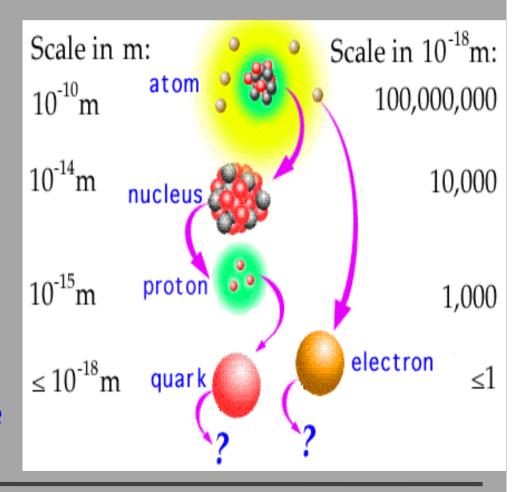
### Setting the Scale for Particle Physics



Particle physics is the study of the most basic building blocks of matter and their interactions

#### Why high energy?

- Small distances and high energies: λ = h/p
- Optical resolution proportional to λ
- So, we need high energy/momentum to probe the fundamental building blocks of nature

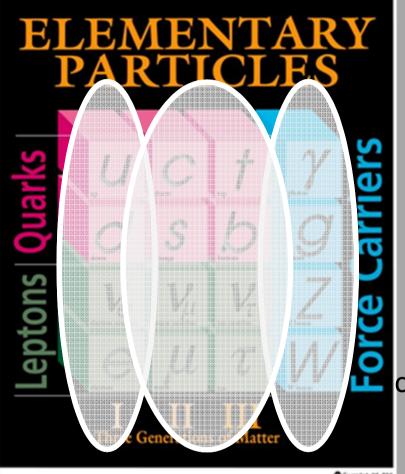




#### The "Standard Model" of Particle Physics



Make up all "regular" matter
In the Universe



Force Carriers

Unstable matter created in high-energy collisions

Standard Model (SM) of particle physics includes these experimentally observed particles and their interactions



## What is the origin of Electroweak Symmetry Breaking?



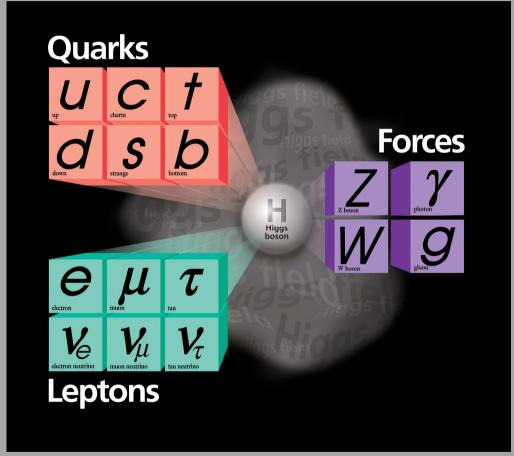
- Consider the Electromagnetic and the Weak Forces
- Coupling at low energy: EM:  $\alpha$ , Weak:  $\alpha/(M_{W,Z})^2$ 
  - Coupling strength governed by the same dimensionless constant
  - Difference due to the mass of the W and Z bosons
    - Electroweak symmetry: M<sub>Y</sub>=M<sub>Z</sub>=M<sub>W</sub>
    - But photons massless and W and Z are massive?
- SM postulates a mechanism of electroweak symmetry breaking via the Higgs mechanism
  - massive vector bosons and mass terms for the fermions
  - → Theory predicts a massive new particle called the Higgs boson!
  - → Higgs mass not predicted by theory



# Are there undiscovered fundamental particles?



The standard model really looks more like this!



Discovery (or exclusion) of the Higgs boson, will shine light on the question of the origin of EWK symmetry breaking



## We have a discovery!



#### Observation of a New Particle in the Search for the Standard Model Higgs Boson with the ATLAS Detector at the LHC

The ATLAS Collaboration

Phys. Lett. B 716, 1 (2012)

Observation of a new boson at a mass of 125 GeV with the CMS experiment at the LHC

The CMS Collaboration\*

Phys. Lett. B 716, 30 (2012)

Congratulations to ATLAS and CMS on this fantastic discovery!



## 3 am Higgs Announcement Party







## 3 am Higgs Announcement Party







## We have a discovery!



Observation of a New Particle in the Search for the Standard Model Higgs Boson with the ATLAS Detector at the LHC

If the Standard Model Higgs boson exists with a mass of 125 GeV/c<sup>2</sup> we should be able to produce it at the Fermilab Tevatron

The CMS Collaboration\*

Phys. Lett. B 716, 30 (2012)

## But, what exactly has been discovered?



### The Tevatron at Fermilab



The Tevatron provided the World's highest energy proton-antiproton collisions:

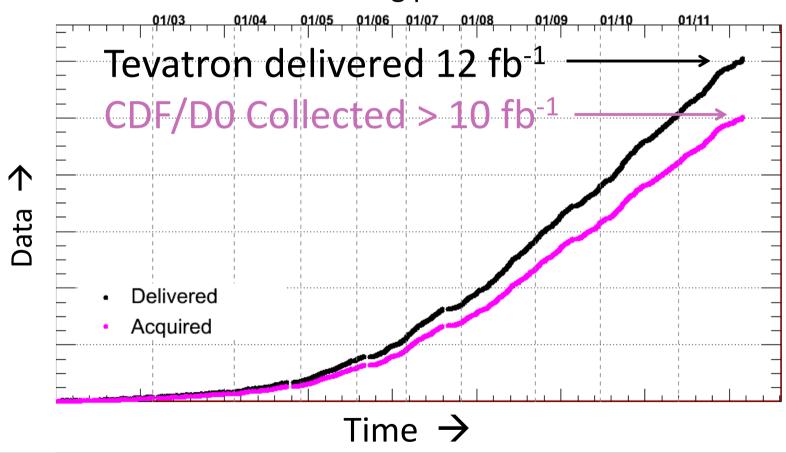




#### **Tevatron Performance**



An enormous thanks to the Fermilab Accelerator Division for an outstanding performance!

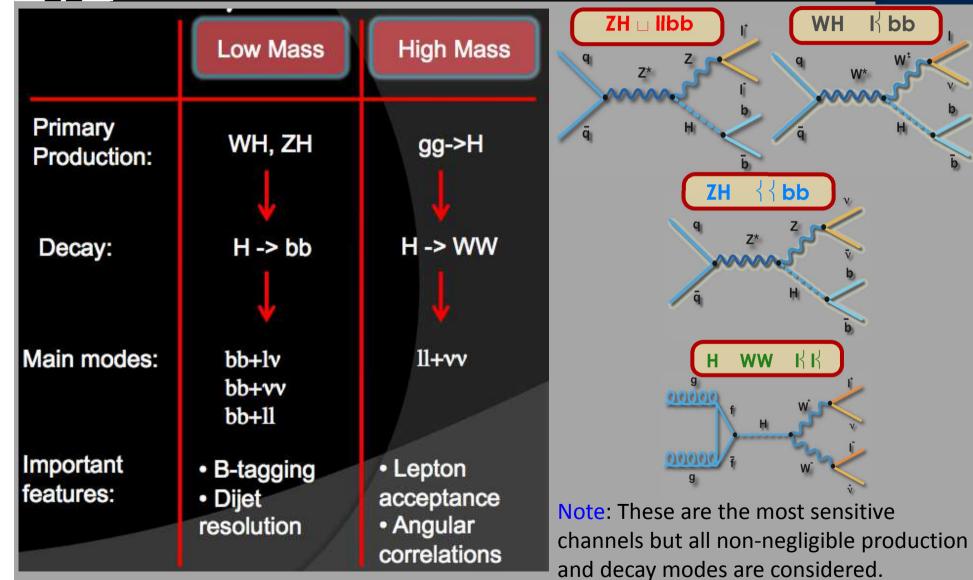




### Higgs Searches: The basics



K bb





### Higgs Searches: The basics



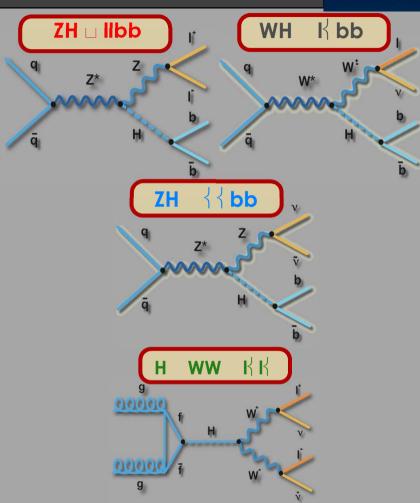
## Complementary to the LHC: ATLAS and CMS

Strong evidence in diphoton channel  $(H-\gamma\gamma)$ 

Evidence of H – ZZ

Weak Evidence of H- WW

No evidence so far of fermionic decays (Tau channels are data deficit)

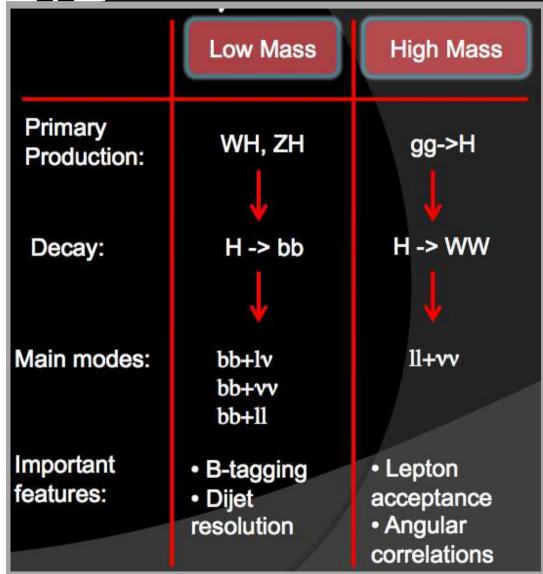


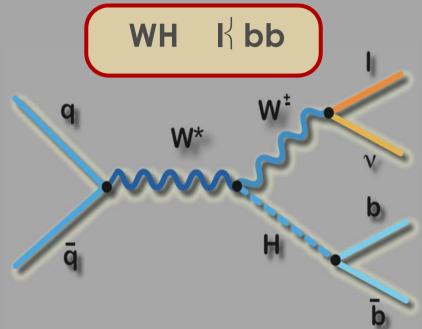
Note: These are the most sensitive channels but all non-negligible production and decay modes are considered.



## Higgs Searches: The basics







Our UVA effort focused mostly on this channel so I will use it for examples.

WH-lvbb at CDF is the most sensitive Tevatron channel for 125 GeV Higgs.

For WH-lvbb we need to identify leptons, neutrinos, and b-quarks.



#### Particle Identification

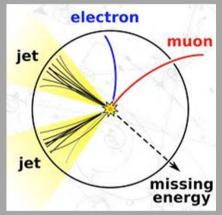


Photons

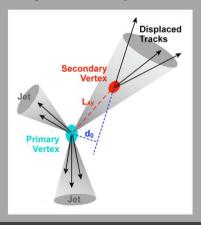
Electrons
Positrons
Muons

Charged
hadrons
Neutral
hadrons

Quarks and gluons "hadronize" to form "jets" of clusters of particles in our detector.



#### b-jets are special:



Neutrinos

Innermost layer

Outermost layer



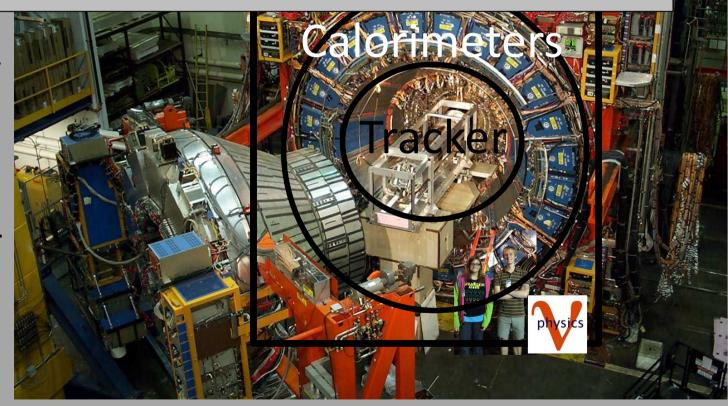
## The CDF Experiment at FNAL



# Our UVa effort joined the CDF experiment in October 2010!

Collaboration

15 Countries 56 Institutions 475 Authors ~50 pubs/year





## CDF @ UVa

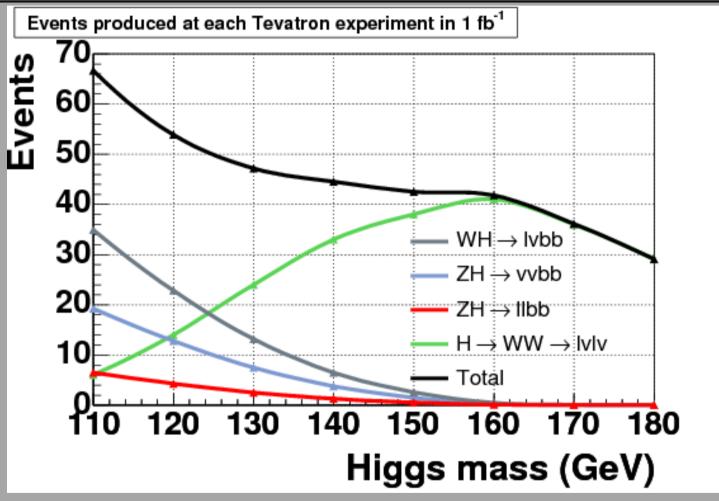






### Higgs Production Rates



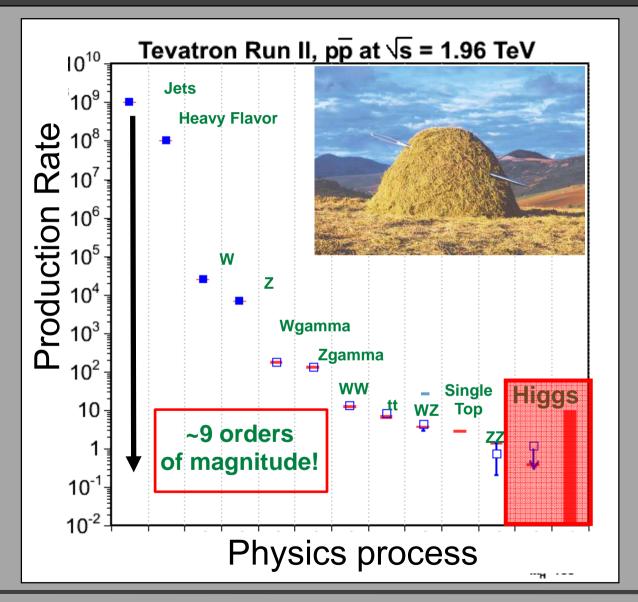


About 1000 Higgs events expected in the main search channels with the full Tevatron dataset (10 fb^-1)



#### Rates of Physics Processes at the Tevatron









## Analysis Flow



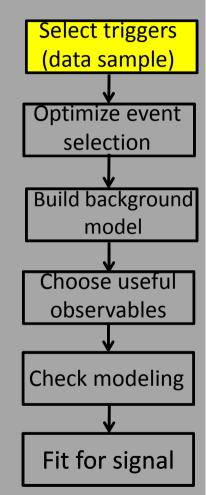
## Triggers



- •Collisions occur at a rate of  $\sim$  2.5 MHz!
  - → More than 99.9% jet events
- •We select (trigger) potentially useful events and throw the rest away!
- •There are many triggers to choose from: We can trigger on the lepton ( $e \text{ or } \mu$ ):
  - WH  $\rightarrow \ell \nu bb$ , ZH  $\rightarrow \ell \ell bb$ , H  $\rightarrow WW$

Or, MET + jets (MET= missing transverse energy):

•WH  $\rightarrow \ell \nu bb$ , ZH  $\rightarrow \nu \nu bb$ 

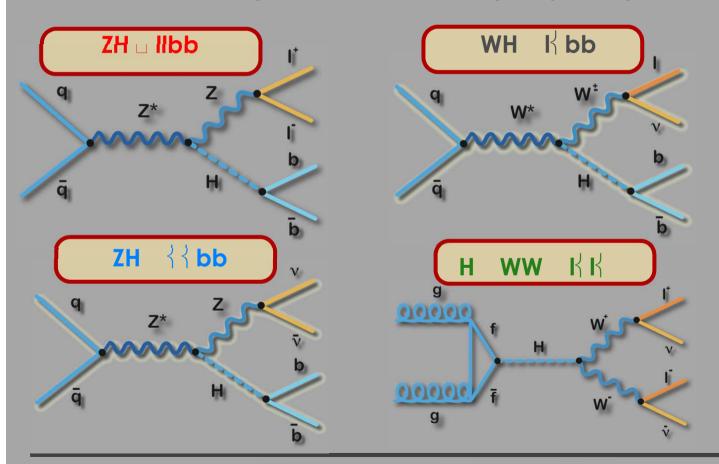


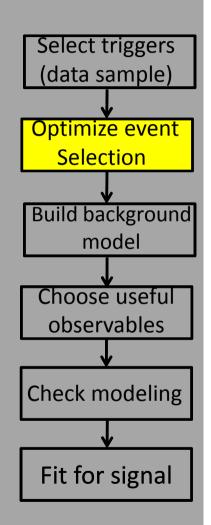


#### **Event Selection**



Based on the final state content, event selection is optimized to maximize signal acceptance and sample purity





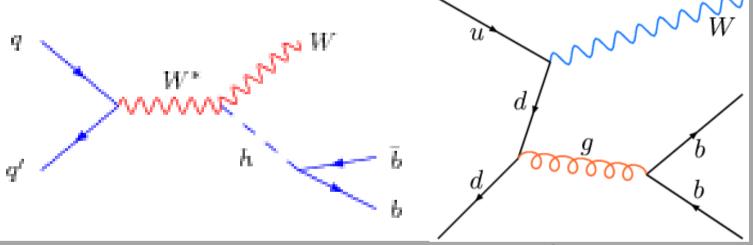


#### Backgrounds to the Higgs Boson Signal



Backgrounds are events from other processes that pass Higgs event selection.

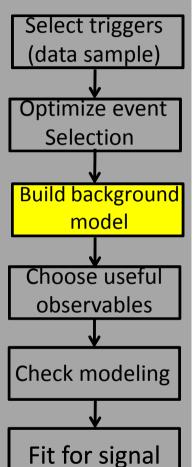
Examples from WH | | bb

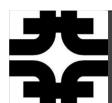


Signal

Physics Background

Other backgrounds from top, dibosons, ...



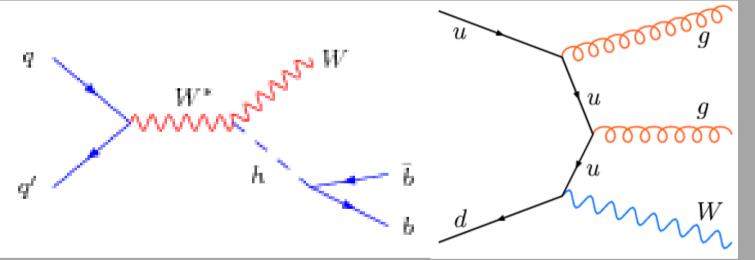


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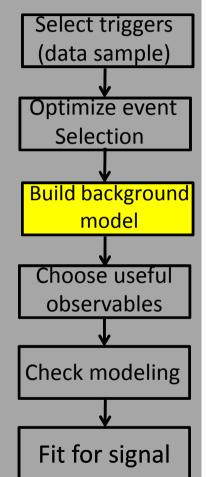
Examples from WH K bb



Signal

Instrumental Background

Other backgrounds from top, dibosons, ...





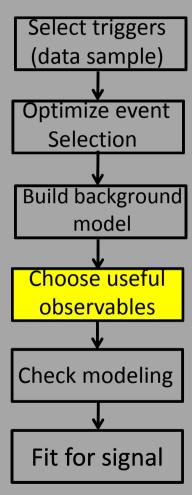
#### Variable Selection

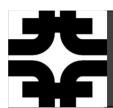


Uncertainties on background are larger than expected signal

Simple counting experiment won't work!

Need to use shape information...





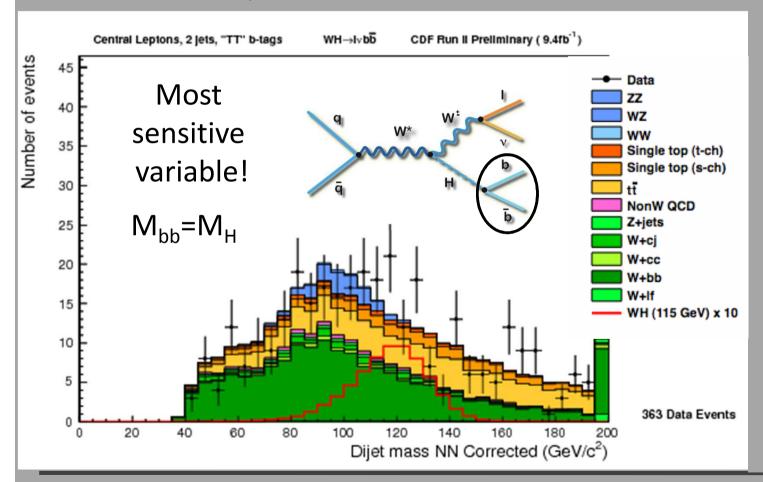
#### Variable Selection

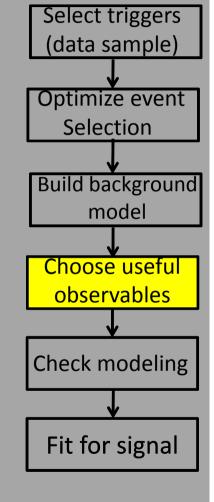


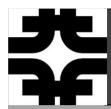
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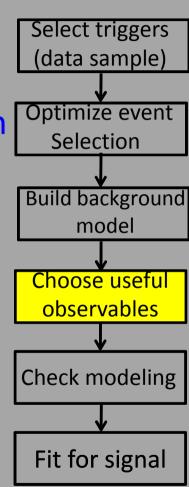




#### Other Variables



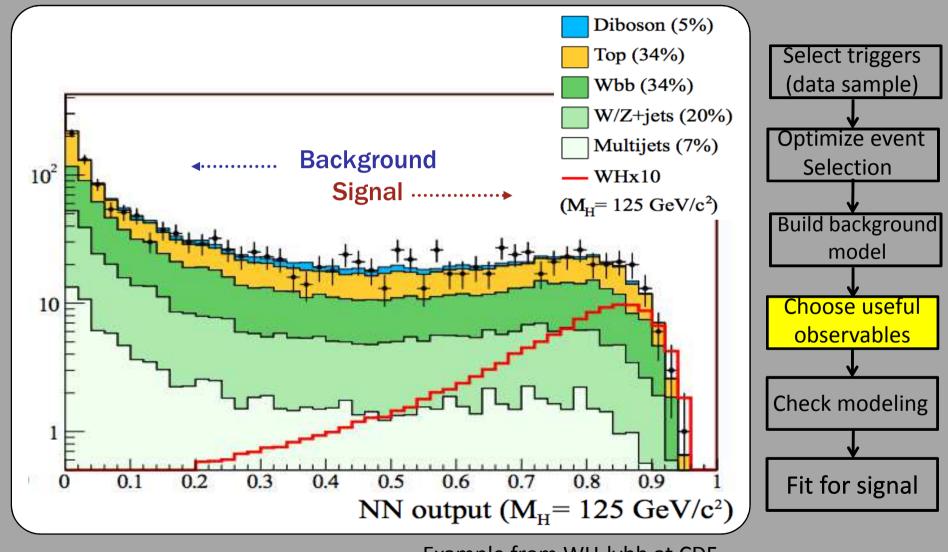
- For low-mass Higgs searches, M<sub>jj</sub> is the most sensitive, but other variables can add separation power between signal and background
- Multivariate discriminant: Just a function with multiple parameters optimized to discriminate signal from background
- Using multivariate techniques improves sensitivity by ~25% over just using M<sub>ii</sub>!

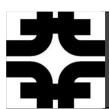




## Multivariate Techniques





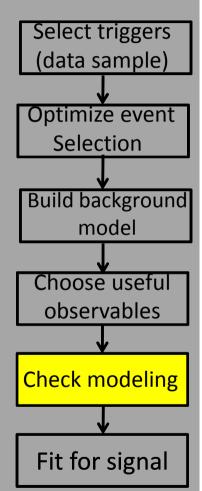




Multivariate techniques are only as good as the modeling of the input variables...

We check our background model in many ways:

- Compare input variables and MVA output with data in control regions and signal regions
- Build confidence in the robustness of techniques and background model by successful measurements of other processes in similar final states.
  - Examples:
    - Single top production
    - Diboson production

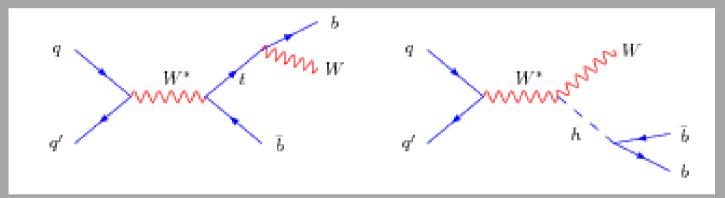




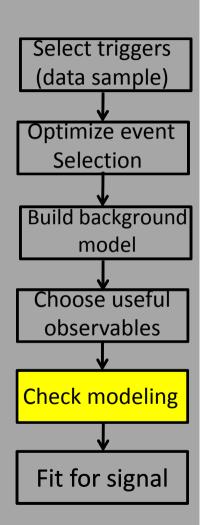


#### Single Top Example

 Measuring single top quark production is a benchmark for WH-lvbb!



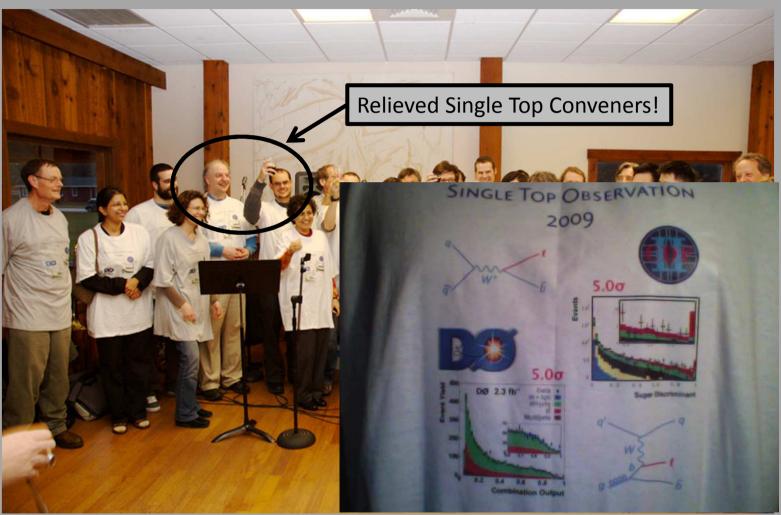
- Same final states, and therefore same analysis strategy, modulo different definitions of signal.
- Same tools and techniques
- Single top is a background for WH
- $\sigma_{WH} \sim 1/10 \, \sigma_{Singletop}$





## Single Top Celebration



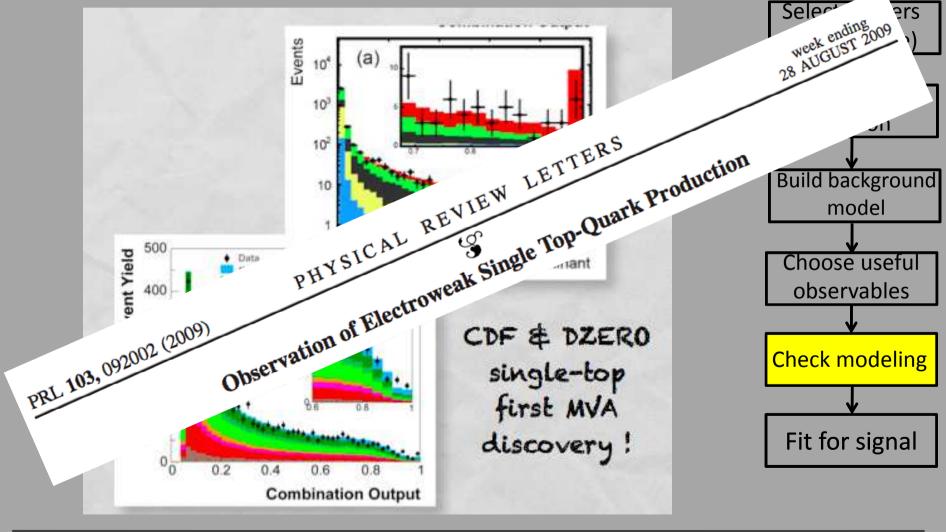


We all got T-shirts.





#### Single Top Example

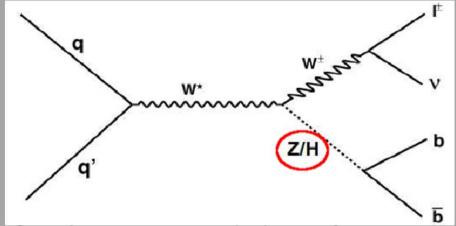




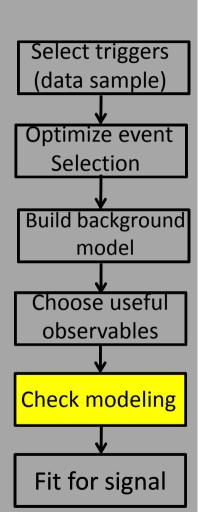


#### **Diboson Example**

 Feynman diagrams are topologically equivalent to Higgs channels



- Same final states, and therefore same analysis strategy, modulo different definitions of signal.
- We combine the main Higgs search channels retraining MVA discriminants for dibosons
- $\sigma_{WH} \sim 1/4 \sigma_{WZ-lvbb}$







Select triggers

(data sample)

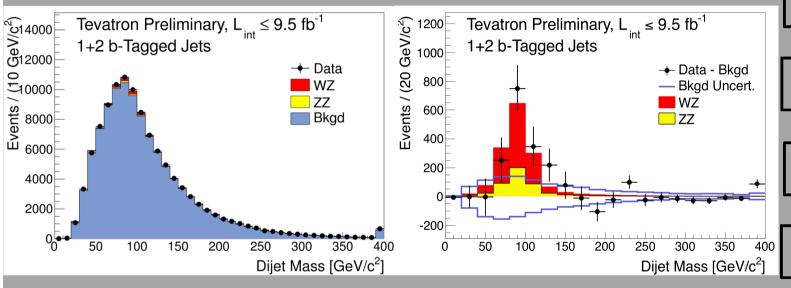
Selection

model

Choose useful

observables

#### **Diboson Example**



Optimize event Build background

**Check modeling** 

Fit for signal

 $\sigma(WZ+ZZ) = 3.9 \pm 0.6(stat) \pm 0.7 (syst) pb$ (SM Prediction =  $4.4 \pm 0.3$  pb)

(Note: MVA outputs used for cross section fit)



## Fit For Signal

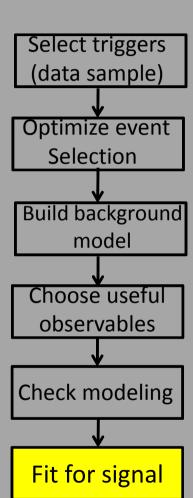


- We set limits on the Higgs boson production rate
- Use a combined binned likelihood fit:

$$L(R, \vec{s}, \vec{b} | \vec{n}) = \prod_{i=1}^{N_{\text{channel}}} \prod_{j=1}^{N_{\text{bin}}} \frac{\mu_{ij}^{n_{ij}} e^{-\mu_{ij}}}{n_{ij}!}$$

- Uncertainties incorporated as nuisance parameters
   -shape and normalization of background and signal
- These 95% C.L. limits say:

If the true Higgs production rate was at the limit value, we would see evidence of the Higgs signal more significant than what we observed in 95 % of experiments





## Fit For Signal

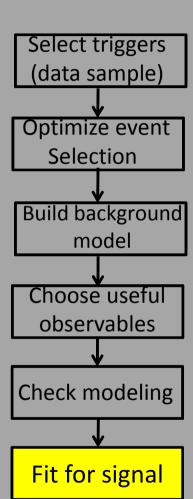


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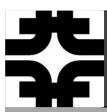
## **Combining All Channels**



Low-Mass Channels	Sensitivity (m <sub>H</sub> = 125 GeV)	
	CDF [SM]	D0 [SM]
WH → Iv bb	2.8	4.7
VH → MET + bb	3.6	3.9
ZH → II + bb	3.6	5.1
$H \rightarrow \gamma \gamma$	9.9	8.2
VH→eμ + X	-	5.9
VH→ eeμ/μμe/ττμ + X	-	7.1
VH, VBF, ggH → 1/2 jet + тт	12	-

etc.

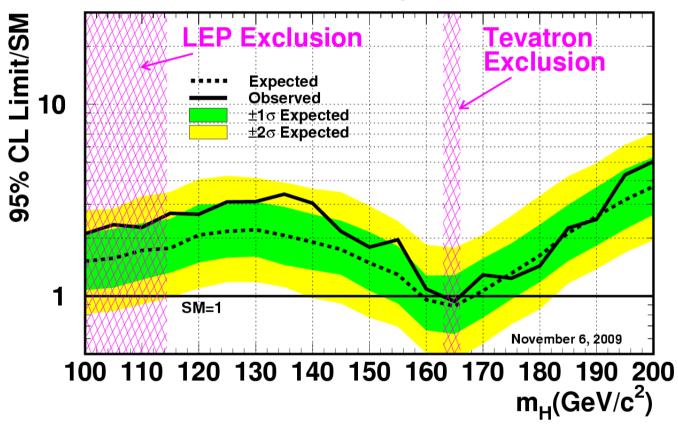
High-Mass Channels	Sensitivity (m <sub>H</sub> = 165 GeV)	
	CDF [SM]	D0 [SM]
$gg \rightarrow H \rightarrow WW \rightarrow IvIv$	0.7	0.8
H→WW→Ivqq'	-	5.2
H→ZZ→ IIII	20	-





(The result that I showed at my interview in early 2010)



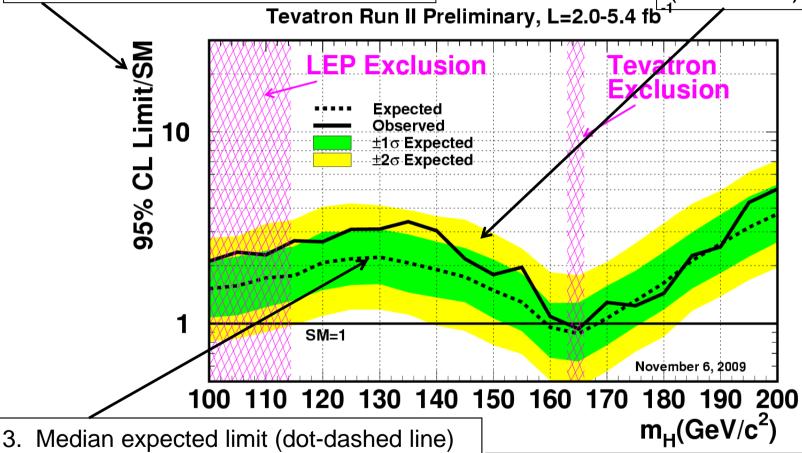






1. Upper cross section limit for Higgs production relative to SM prediction

Observed limit (solid line) from data

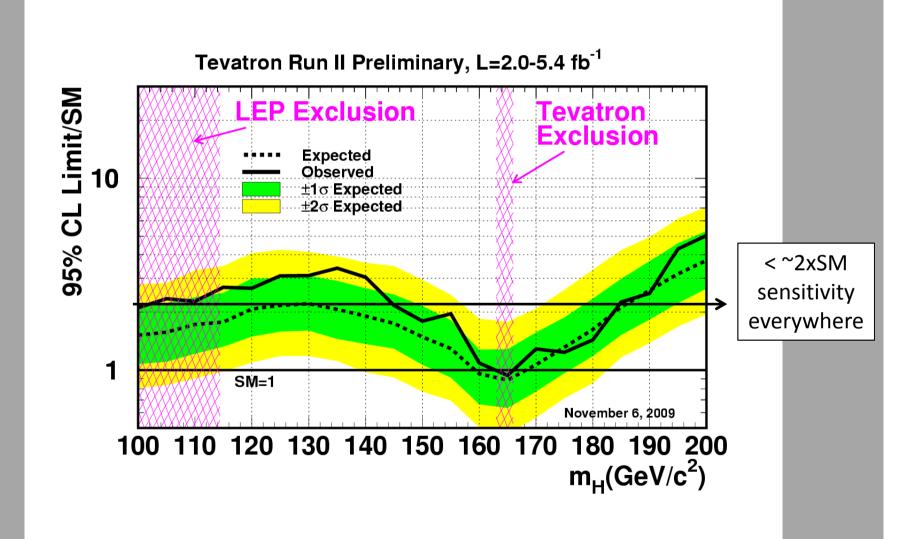


- 3. Median expected limit (dot-dashed line) and predicted 1σ/2σ (green/yellow bands) excursions from background only pseudo-experiments

  C. Group UVA 2012 -
  - 4. Analysis repeated using different signal templates for each m<sub>H</sub> between 110 and Hiஜூசெலிங்வுக் GeV steps 40

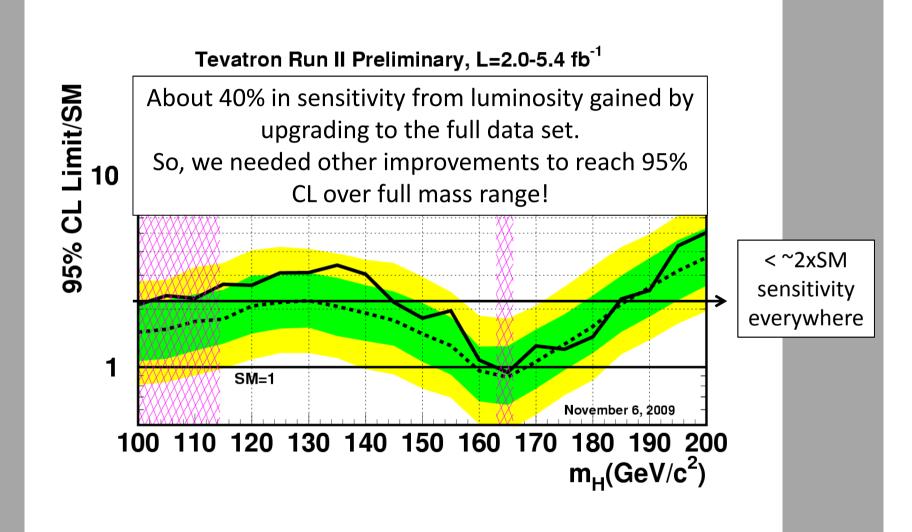
















## **UVA** Improvements

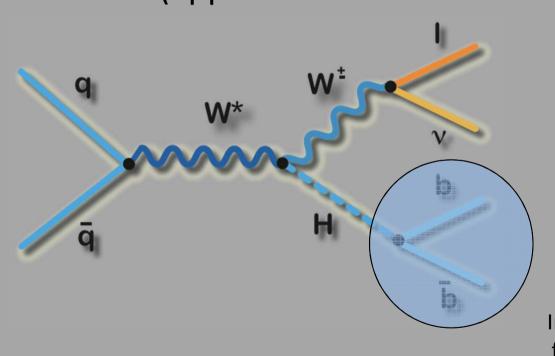


## Improvement Examples



Gain sensitivity by improving b-jet ID efficiency gains.

(applies to all h-bb channels)





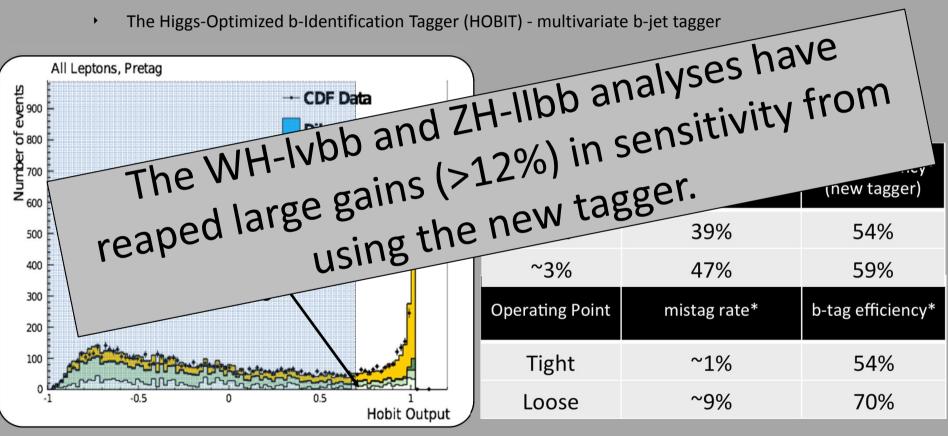
Yuri Oksuzian
Improved b-quark identification strategy
for the low mass Higgs searches at CDF



### New b-tagger at CDF



- By 2010 CDF had at least 5 types of b-taggers
- We incorporate the knowledge from previous taggers into:
  - The Higgs-Optimized b-Identification Tagger (HOBIT) multivariate b-jet tagger



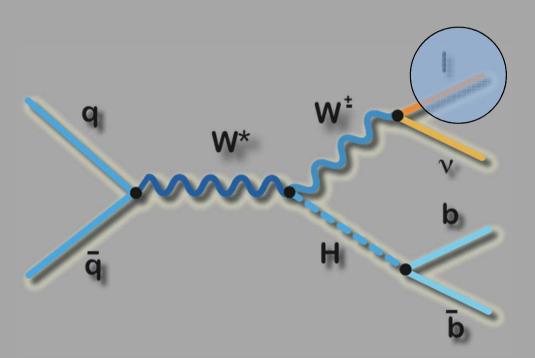
Form 5 orthogonal tagging categories: TT, TL, T, LL, L



## Improvement Examples



Gain sensitivity by improving lepton trigger acceptance.



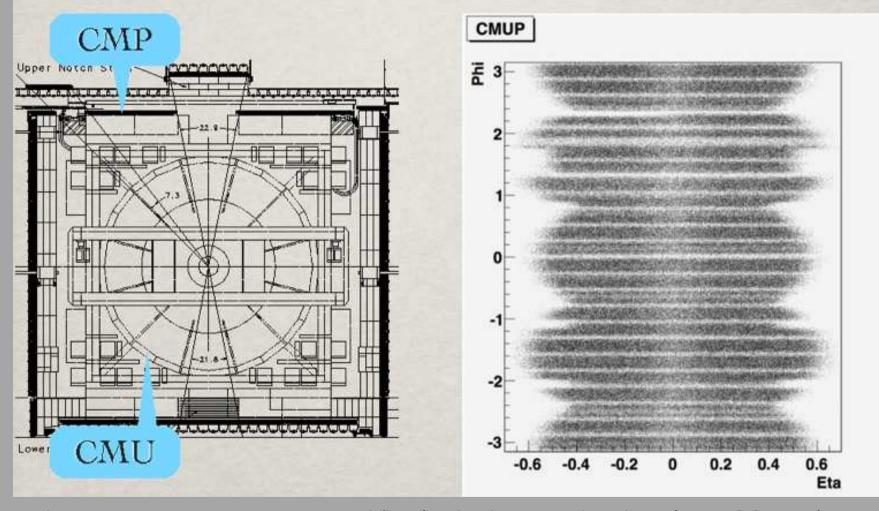


Hao Liu
Improving the Trigger Efficiency for the
WH-lvbb analysis at the CDF experiment



## Muon Trigger Improvement





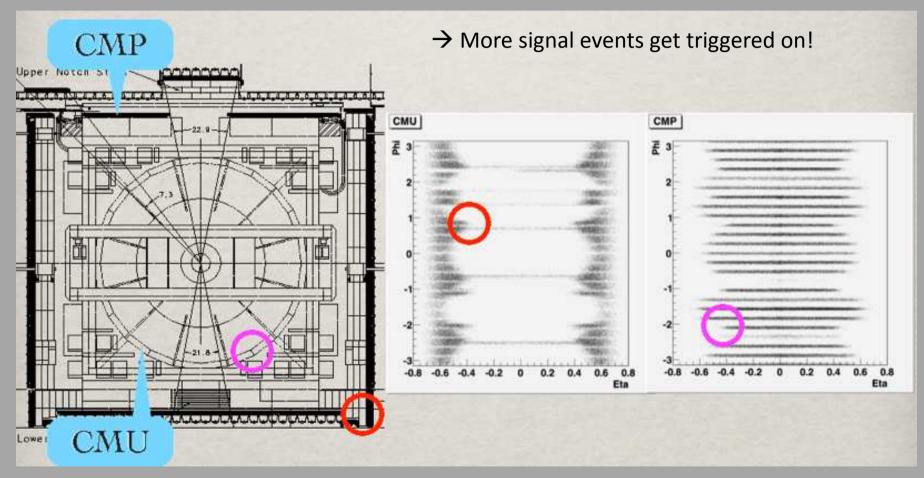
The primary muon trigger required 'hits' in both muon chambers (CMU && CMP) leaving many cracks in the trigger coverage.



## Muon Trigger Improvement



In 2007 I created triggers based on CMU-only and CMP-only to improve trigger acceptance.



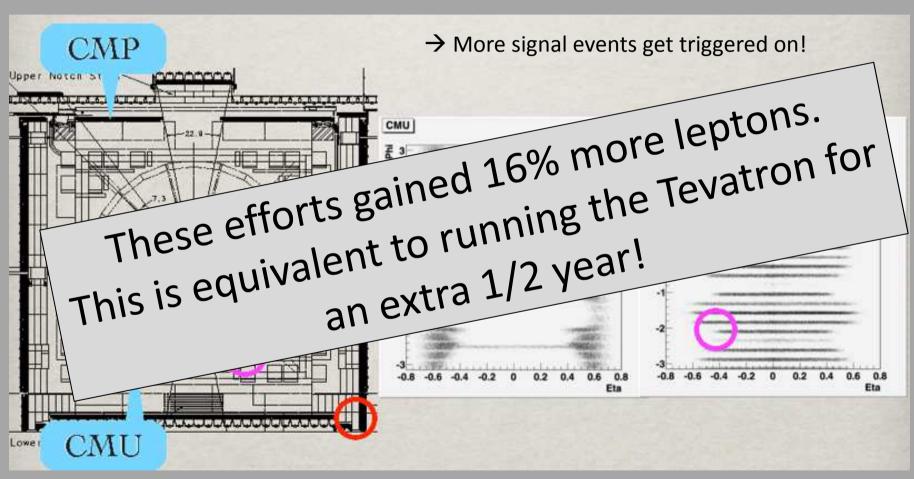
In 2011 Hao Liu added these extra events to the CDF WH-lvbb Higgs search. Hao also added an "inclusive" method to measure efficiency for a suite of triggers



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In 2011 Hao Liu added these extra events to the CDF WH-lvbb Higgs search. Hao also added an "inclusive" method to measure efficiency for a suite of triggers







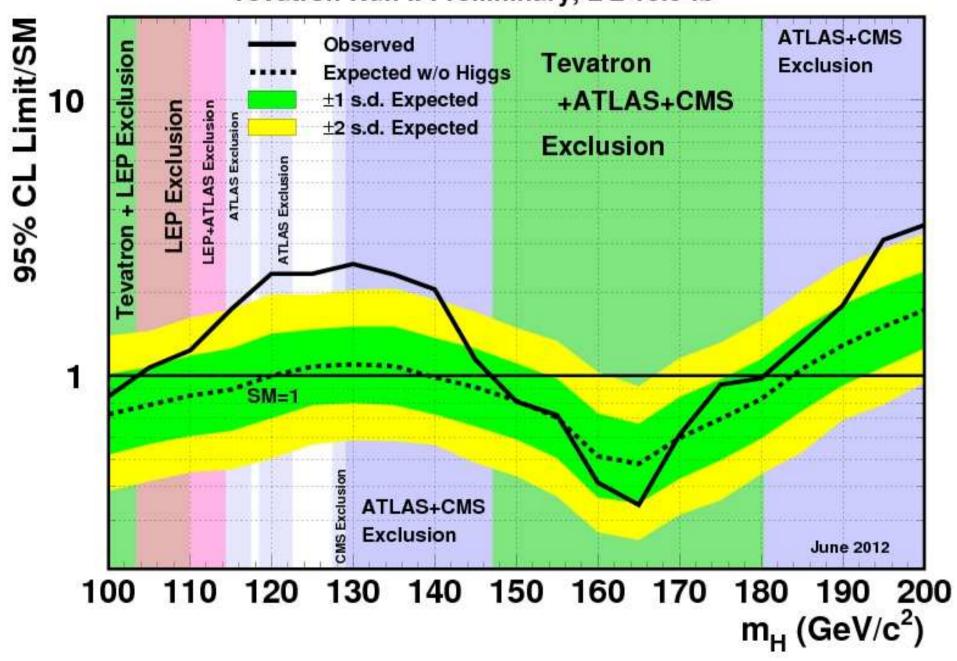
The 10 Million Dollar Man!



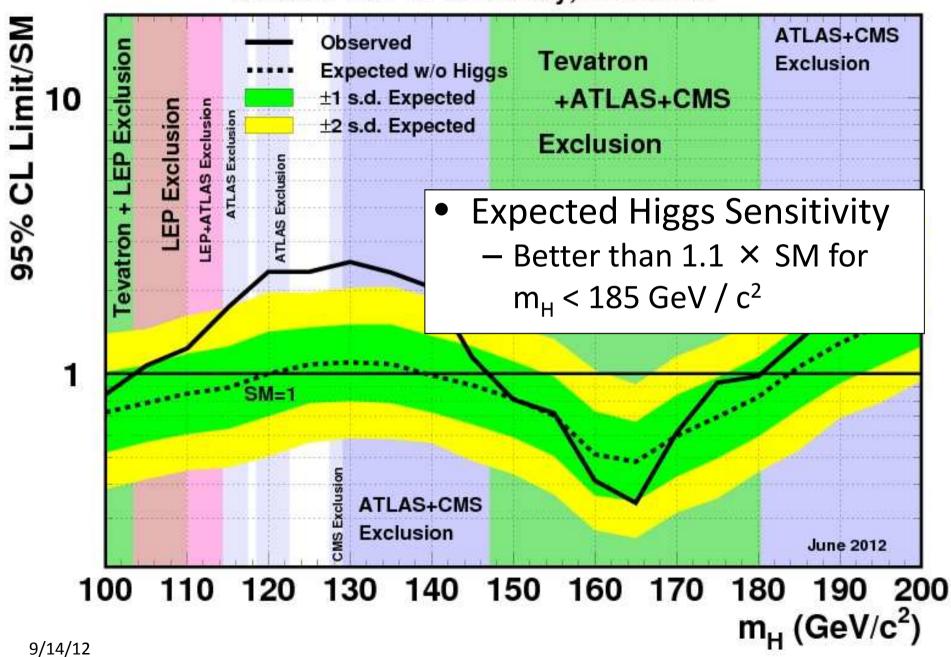


## Results

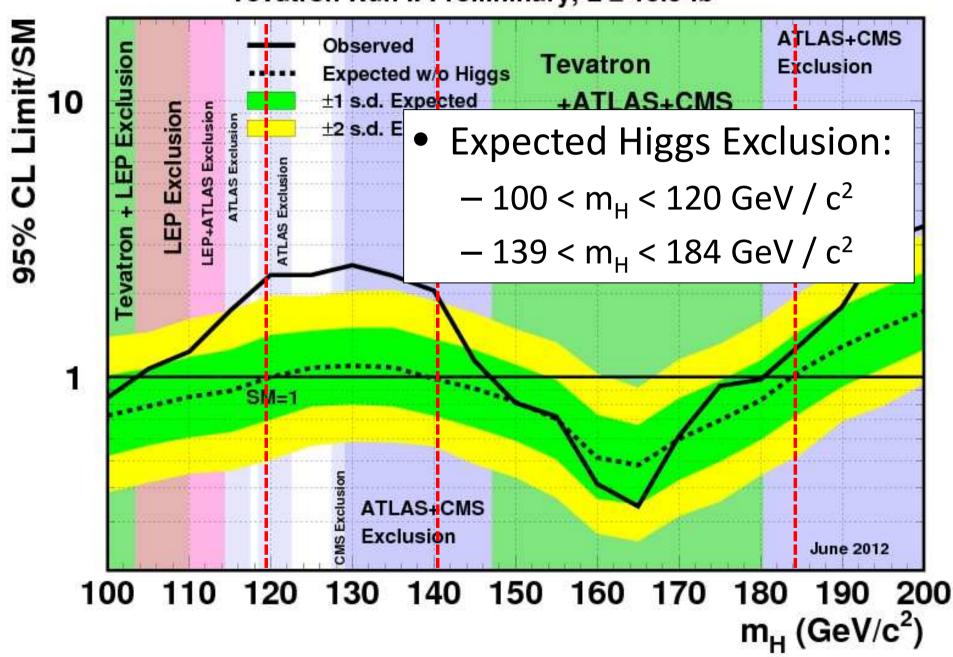
Tevatron Run II Preliminary, L ≤ 10.0 fb<sup>-1</sup>



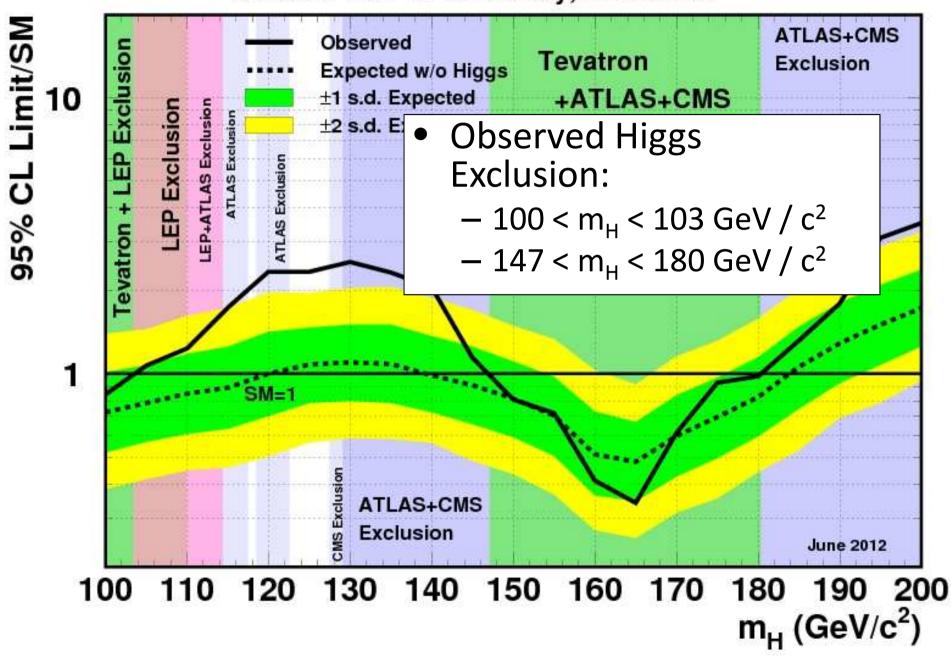
Tevatron Run II Preliminary, L ≤ 10.0 fb<sup>-1</sup>



Tevatron Run II Preliminary, L ≤ 10.0 fb<sup>-1</sup>



Tevatron Run II Preliminary, L ≤ 10.0 fb<sup>-1</sup>





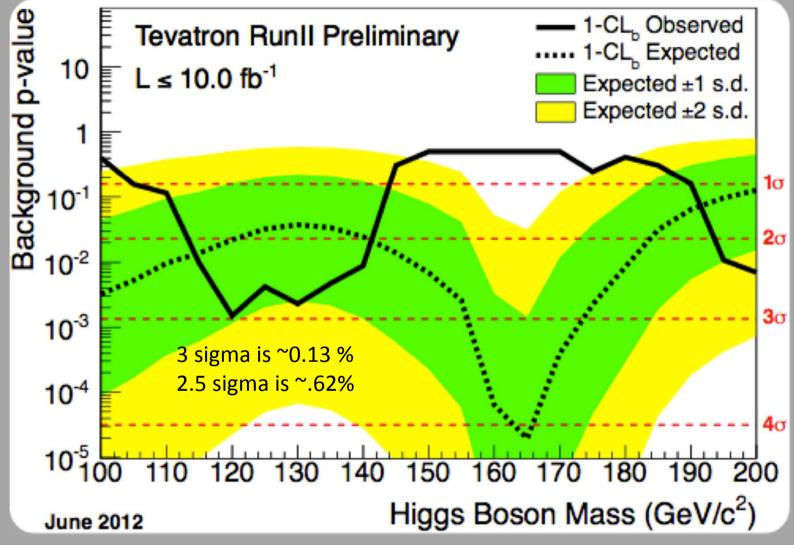


## The Excess



## Is it consistent with background only?





Local: 3.0 sigma  $\rightarrow$  Global: 2.5 sigma (LEE of 4 for the region 100-200 GeV/c<sup>2</sup>)

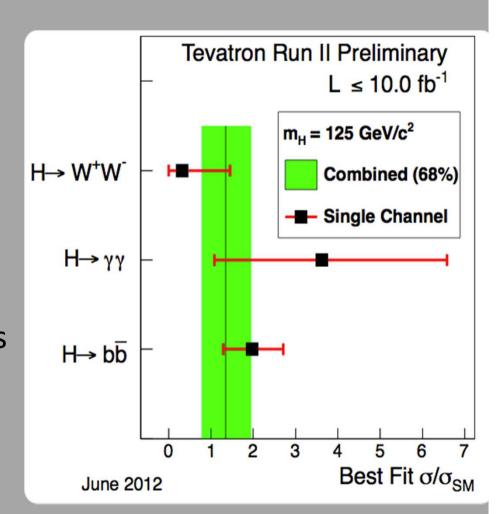


### Is it consistent with SM Higgs signal?



Perform fit of S+B model to data

Compare combined best fit Higgs production cross section to result from individual production modes

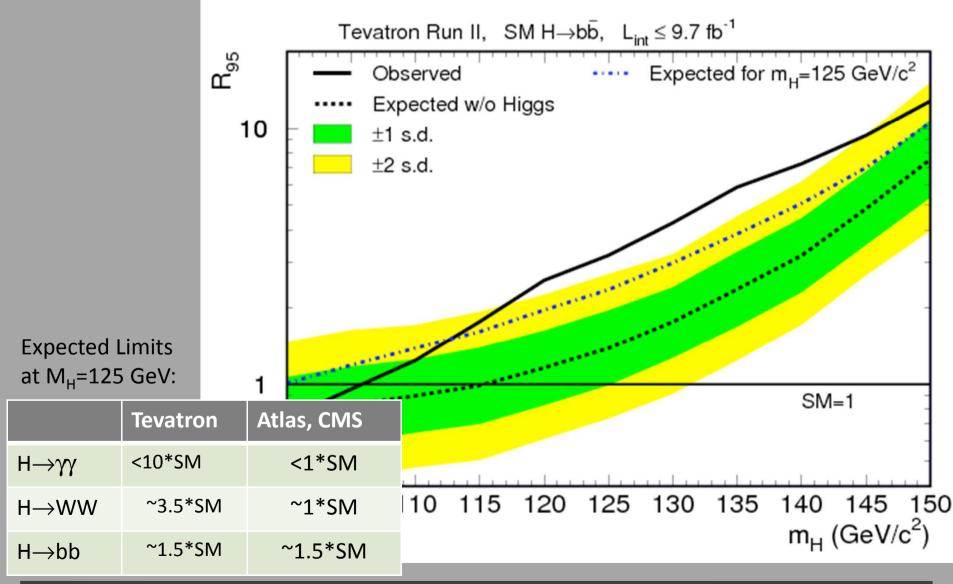


Consistent with SM values within the uncertainties



## Strength of Tevatron is H→bb

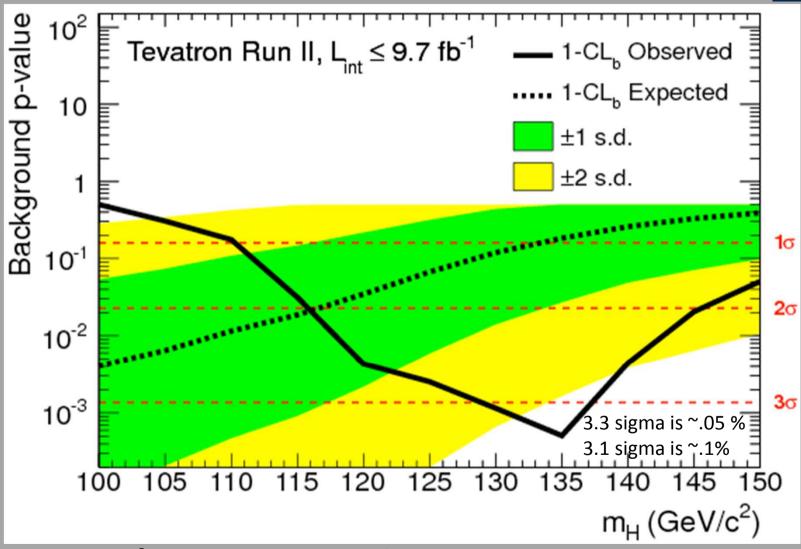






## Strength of Tevatron is H→bb



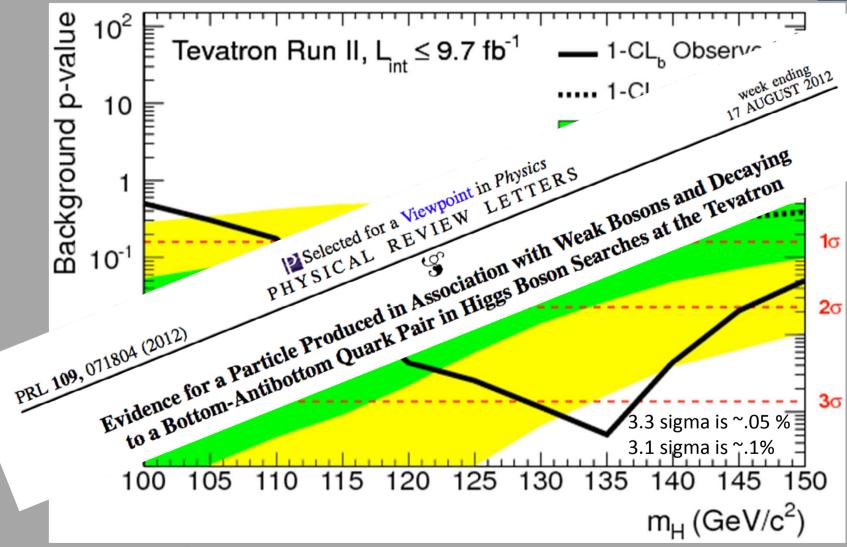


Local: 3.3 sigma  $\rightarrow$  Global: 3.1 sigma (LEE of 2 for the region 100-150 GeV/c<sup>2</sup>)



## Strength of Tevatron is H→bb





Local: 3.3 sigma → Global: 3.1 sigma (LEE of 2 for the region 100-150 GeV/c²)



## Higgs Summary



#### Excluded values of M<sub>H</sub> @ 95% CL

Expected	Observed
100-120 GeV	100-103 GeV
139-184 GeV	147-180 GeV

#### Significance of observed excess:

Channels	Local	Global
All Tevatron	3 σ	2.5 σ
H→bb	3.3 σ	3.1 σ



## Higgs Conclusion



- We have incorporated the full Tevatron dataset into most of the major search channels.
- The UVA group made major contributions and improvements to the CDF Higgs search.
- There is a broad excess in the data that is inconsistent with the background only hypothesis at greater than the 2sigma level.
- Significant excesses in complimentary channels to those present at the LHC.
- This is the year of the Higgs boson!





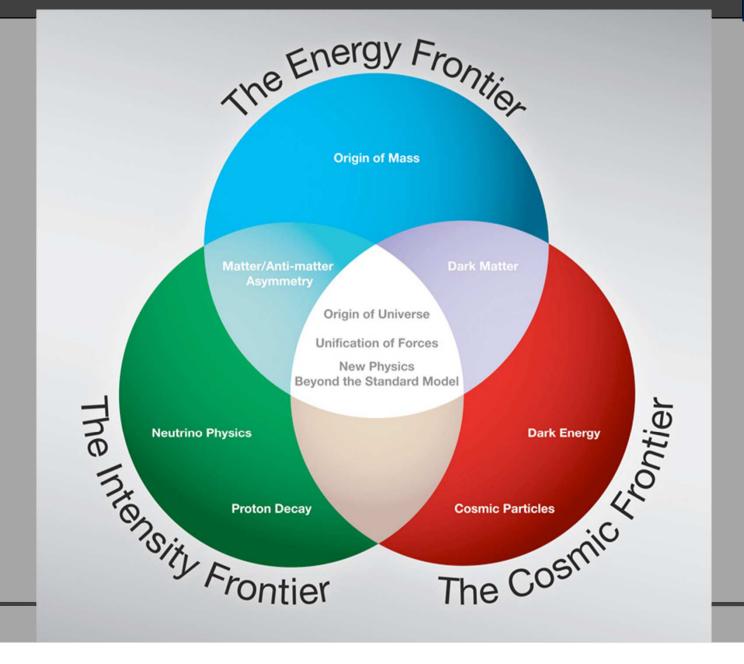


## The Future



#### Research at Fermilab in 3 Frontiers

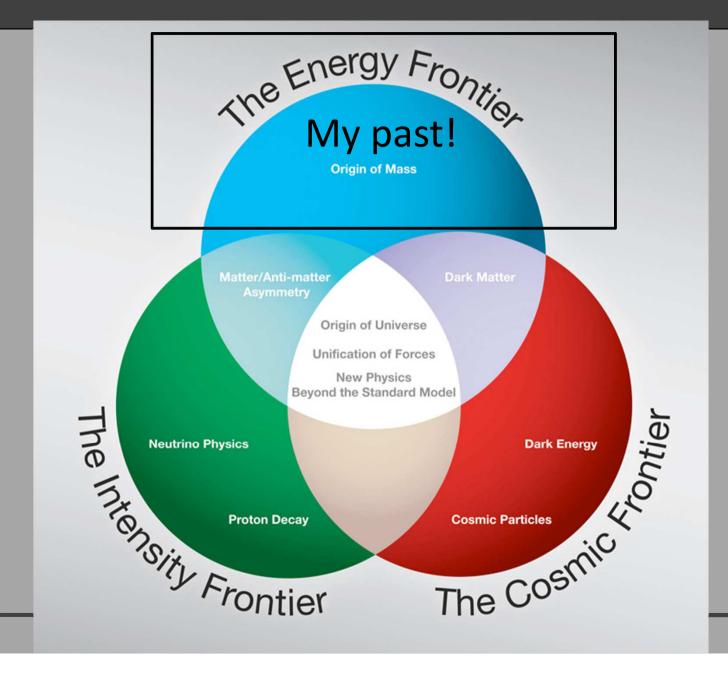






#### Research at Fermilab in 3 Frontiers

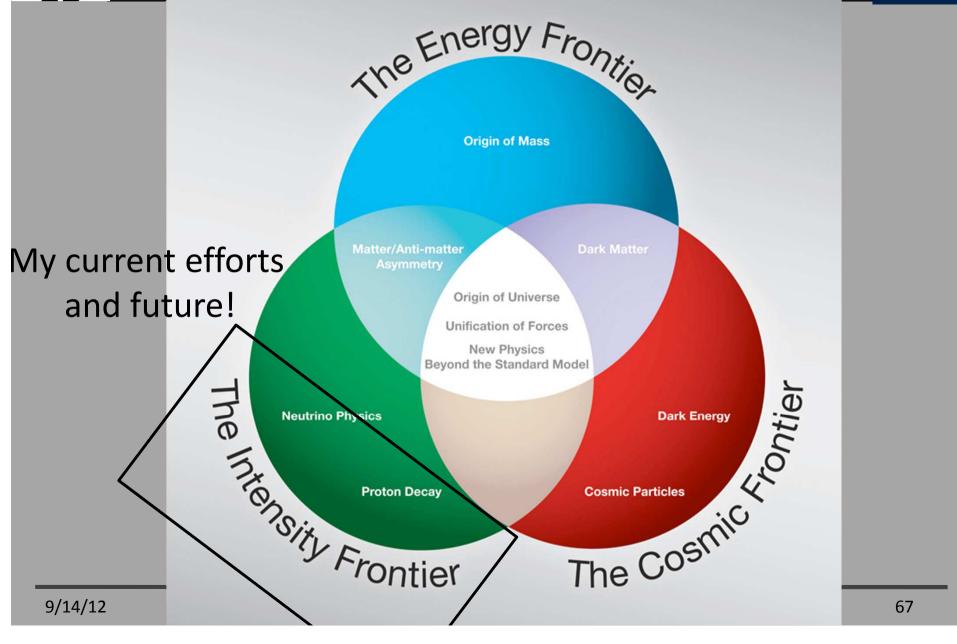


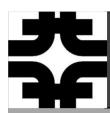




#### Research at Fermilab in 3 Frontiers







## The Intensity Frontier



- One can probe the properties of the universe by looking for extremely rare processes.
- Looking for rare processes using intense beams is an alternative to using higher energies.
  - The Intensity Frontier! (The future of Fermilab)
- The medium term future of particle physics on US soil is dedicated to the intensity frontier:
  - Rare decays (Mu2e for example)
  - Neutrino experiments (NOvA for example)



#### Mu2e



- ullet Goals: Search for:  $\mu^- N o e^- N$ 
  - Measure ratio:

$$R_{\mu e} = \frac{\Gamma(\mu^{-} + (A, Z) \to e^{-} + (A, Z))}{\Gamma(\mu^{-} + (A, Z) \to \nu_{\mu} + (A, Z - 1))}$$

- With sensitivity to R<10<sup>-17</sup>
  - → 3 orders of magnitude better than the competition
  - → Need more than 10<sup>18</sup> mouns!
    - → How to make muons? More than 10<sup>21</sup> protons on target!
  - → Also need very small background
- UVa is the largest University group on Mu2e
  - → Particle Physics and Nuclear Experimentalists!

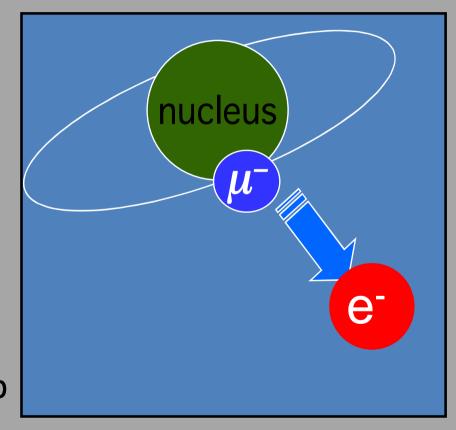


## **Experimental Signature**



$$\mu^- N \rightarrow e^- N$$

- Signal: Single Monoenergetic Electron
- If N = AI,  $E_e = 105$ . MeV
  - Electron E depends on Z
- Nucleus coherently recoils off outgoing electron, no breakup



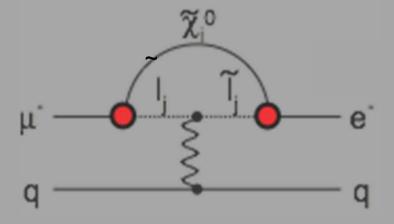


#### **CLFV from SUSY**



## Supersymmetry

rate ~ 10<sup>-15</sup>



# Access SUSY through loops:

Signal of
Terascale at LHC
implies
~40 event signal
and only 0.4 bkg
in this
experiment

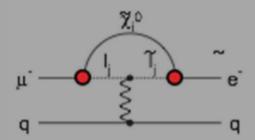


### Other "new physics" probed at Mu2e



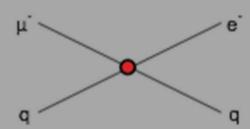
#### Supersymmetry

rate ~ 10-15



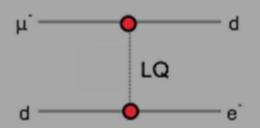
#### Compositeness

 $\Lambda_c \sim 3000 \text{ TeV}$ 



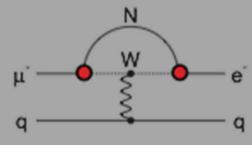
#### Leptoquark

 $M_{LQ} =$  3000  $(\lambda_{ud}\lambda_{ed})^{1/2}$  TeV/c<sup>2</sup>



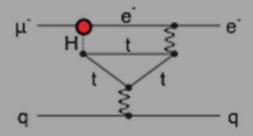
#### Heavy Neutrinos

 $|U_{\mu N}U_{e N}|^2 \sim 8x10^{-13}$ 



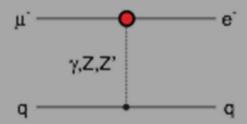
#### **Second Higgs Doublet**

 $g(H_{\mu e}) \sim 10^{-4} g(H_{\mu \mu})$ 



#### Heavy Z' Anomal. Z Coupling

 $M_{Z'} = 3000 \text{ TeV/c}^2$ 





### Cosmic Ray Background to Mu2e



- In order to obtain single-event sensitivity we need to keep background well below one event.
- One dangerous background is from cosmic rays producing electrons in the Mu2e detector.
- Simulation shows that we need about 99.99% veto efficiency to keep this background to a negligible level.

UVA Goal: Design and build a cosmic ray veto (CRV) system that can obtain this level of efficiency.

(~ 5M \$ project)



## CRV prototype R&D efforts









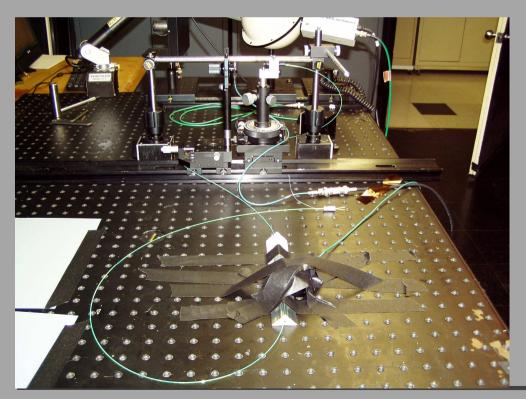
## More R&D (fibers)

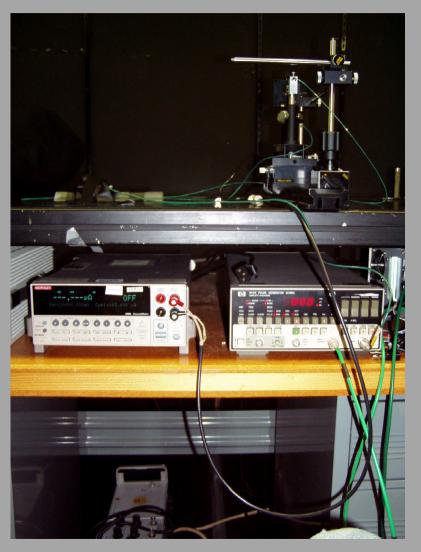


Fiber

**SiPM** 

The idea is to study the angular (radial?) distribution Of light leaving the fiber to optimize the light yield.







## More R&D (neutrons)



#### What?

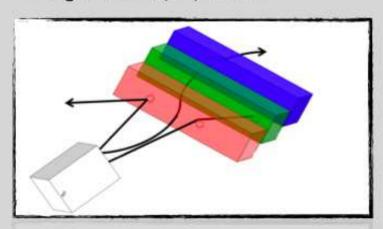
Estimate the CRV efficiency to neutrons

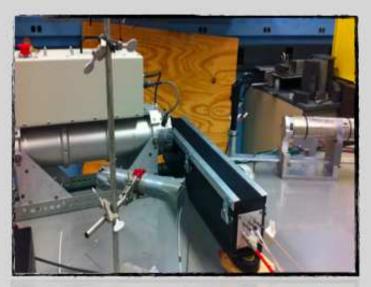
#### How?

- Direct measuring the response of the CRV scintillator to neutron flux
- Estimation from MCNP simulation

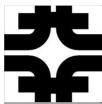
#### So...?:

- Data from neutron gun at UChicago
- MCNP samples simulating our setup with neutrons energies 2.8MeV, 10, 20 MeV



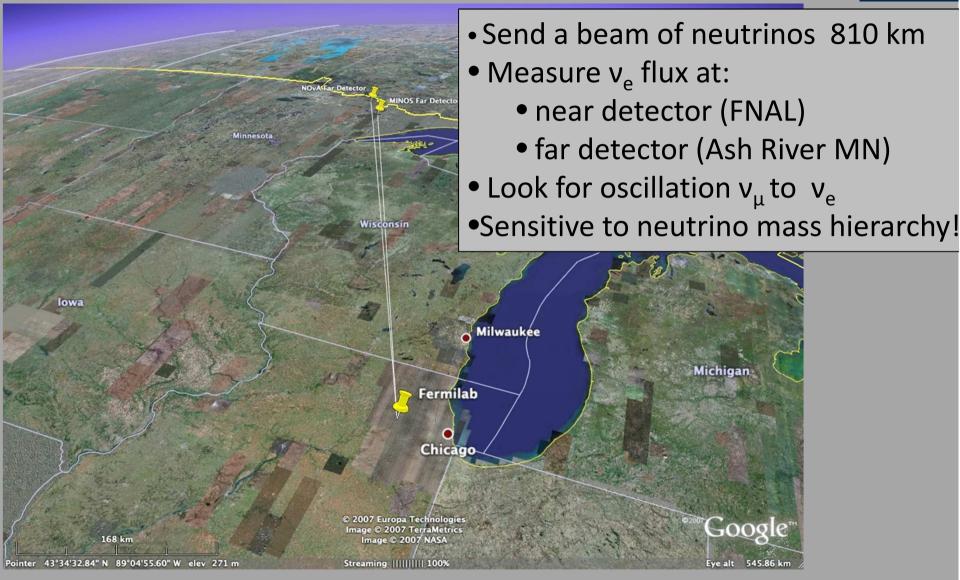


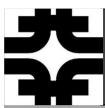




#### NOVA







## NOvA will be enormous!







## 1st block was installed this week!







## NOvA plan

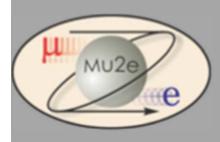


- Far detector being installed now!
- UVa playing a major role in design, construction and commissioning of both detectors and will contribute to physics analysis and operation.
- CG joined NOvA in early 2012 and currently convening computing effort to prepare the collaboration for data analysis.



## Intensity Frontier Conclusion





Mu2e Experiment: Actively contributing already to design and fabrication plan for cosmic ray veto system.



NOvA Experiment: Leading the computing effort to prepare for data analysis. The far detector is being constructed and analysis will start within the next year.



## The Group Group



