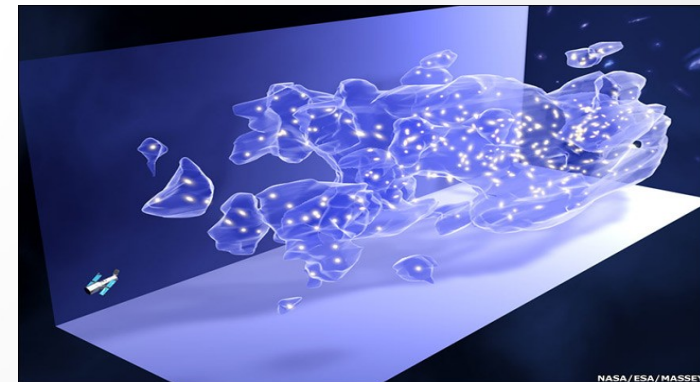


A Search For Dark Forces At The Jefferson Lab Free Electron Laser

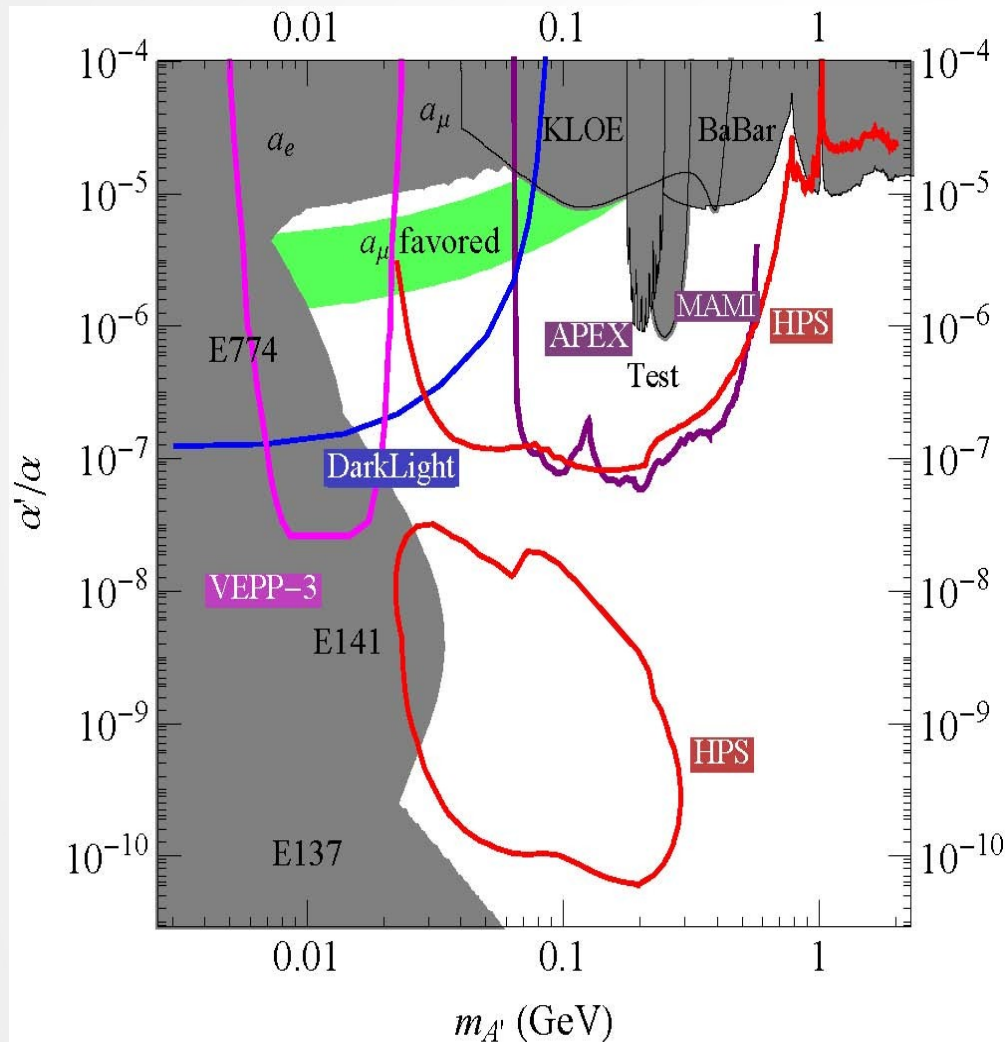


Narbe Kalantarians
Hampton University

*UVA Nuclear Physics Seminar
26 February 2013*

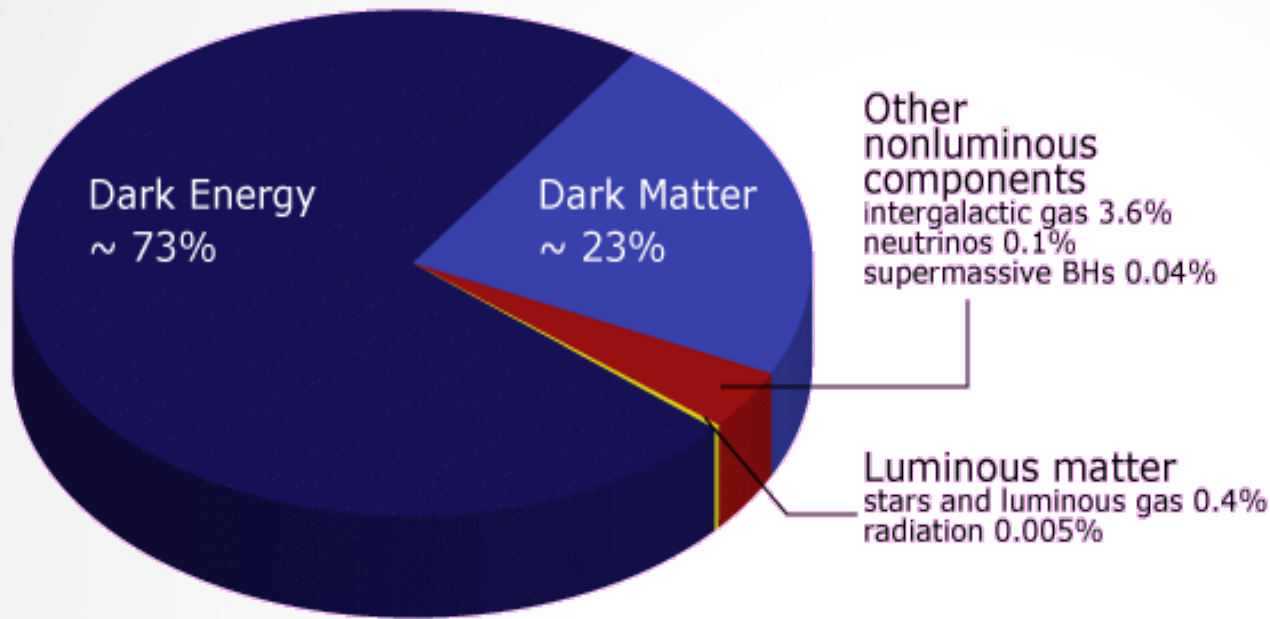


Outline



- Background/Motivation
- DarkLight Experiment
- JLab PAC Proposals
- Beam Test Run & Background Rad. Measurements
- Timeline
- Summary

Much Ado About (Almost) Nothing



The ordinary matter we observe is a trace part of the Universe.

**Another way of picturing this:*
Universe ~ Cup of Coffee
Dark Energy ~ Coffee
Dark Matter ~ Creamer
Observable Matter ~ Sweetener

Dark a. [OE. dark, derk, deork, AS. Dearc, deorc; cf. Gael. & Ir. dorch, dorch, dark, black, dusky.]

1. Destitute, or partially destitute, of light; not receiving, reflecting, or radiating light; wholly or partially black, or of some deep shade of color; not light-colored; as, a dark room; a dark day; dark cloth; dark paint; a dark complexion.

[1913 Webster]

2. Matter, not (observed) interacting with “ordinary” matter, other than gravitationally.



Dark Matter (mass \sim GeV – TeV)

χ^0

Germanium

3V

h

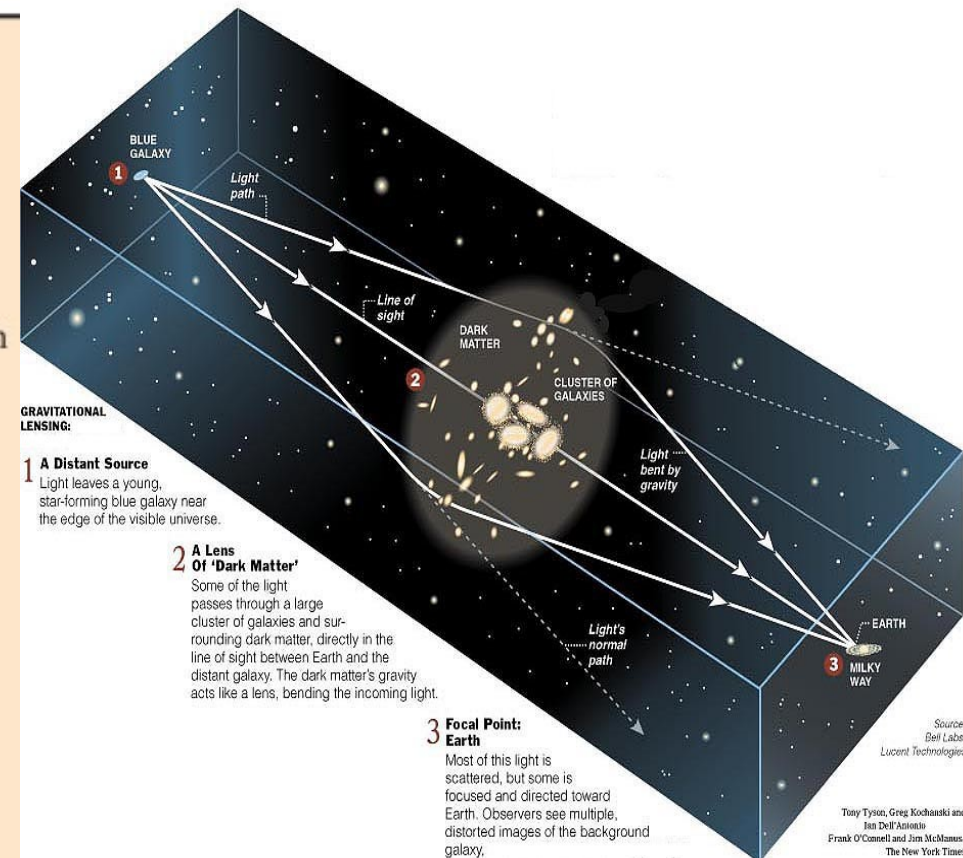
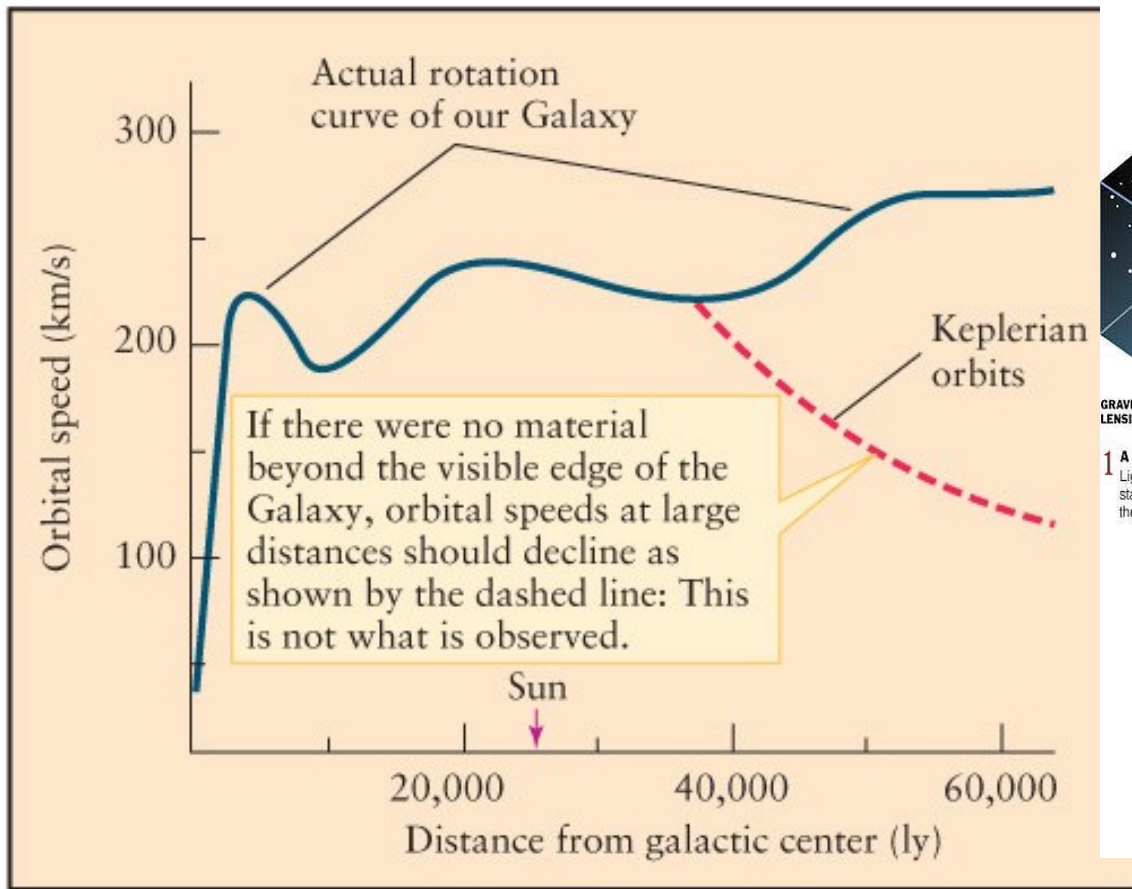
e

phonons

recoil energy (tens of keV)

Indicators for Dark Matter

Astrophysics



Galaxies: Rotation, Gravitational Lensing

Indicators for Dark Matter

Feldman, Liu, Nath
[arXiv:0810.5762v3](https://arxiv.org/abs/0810.5762v3)

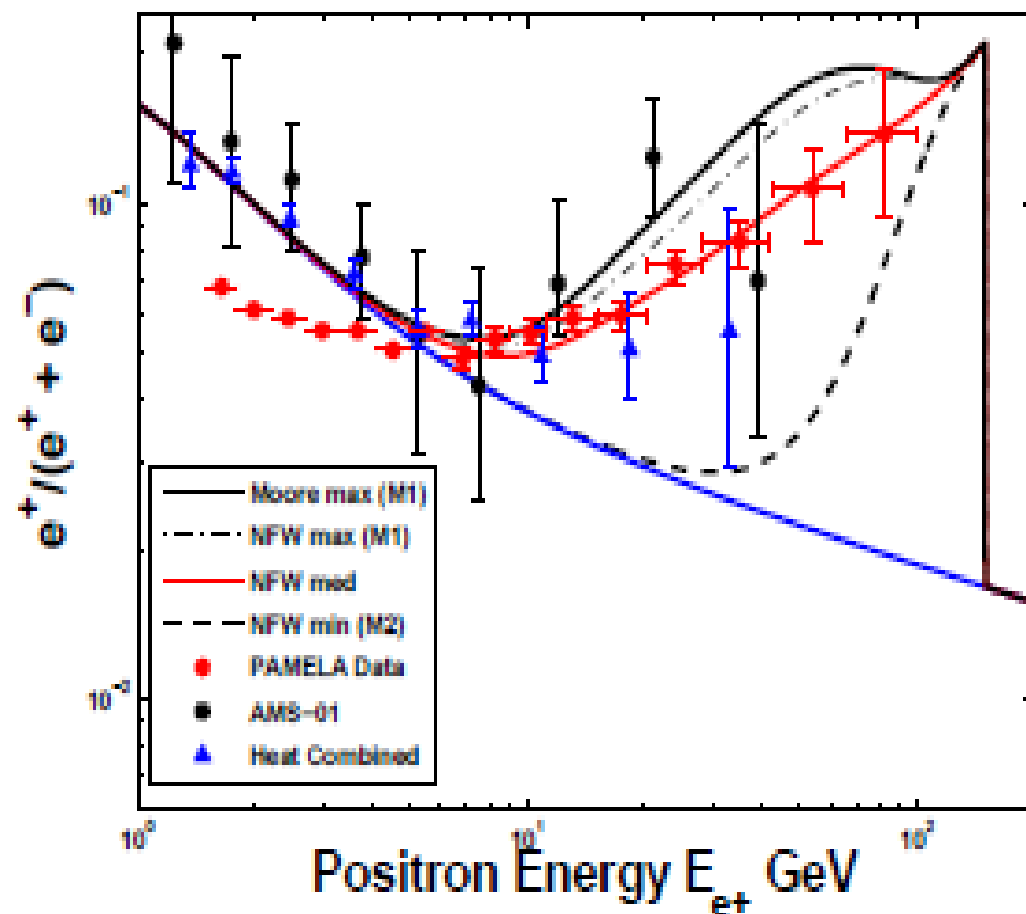
Nuclear/Particle Physics

PAMELA Results:
 e^+ excess
(not for p^- , though)

Sizable cross-section

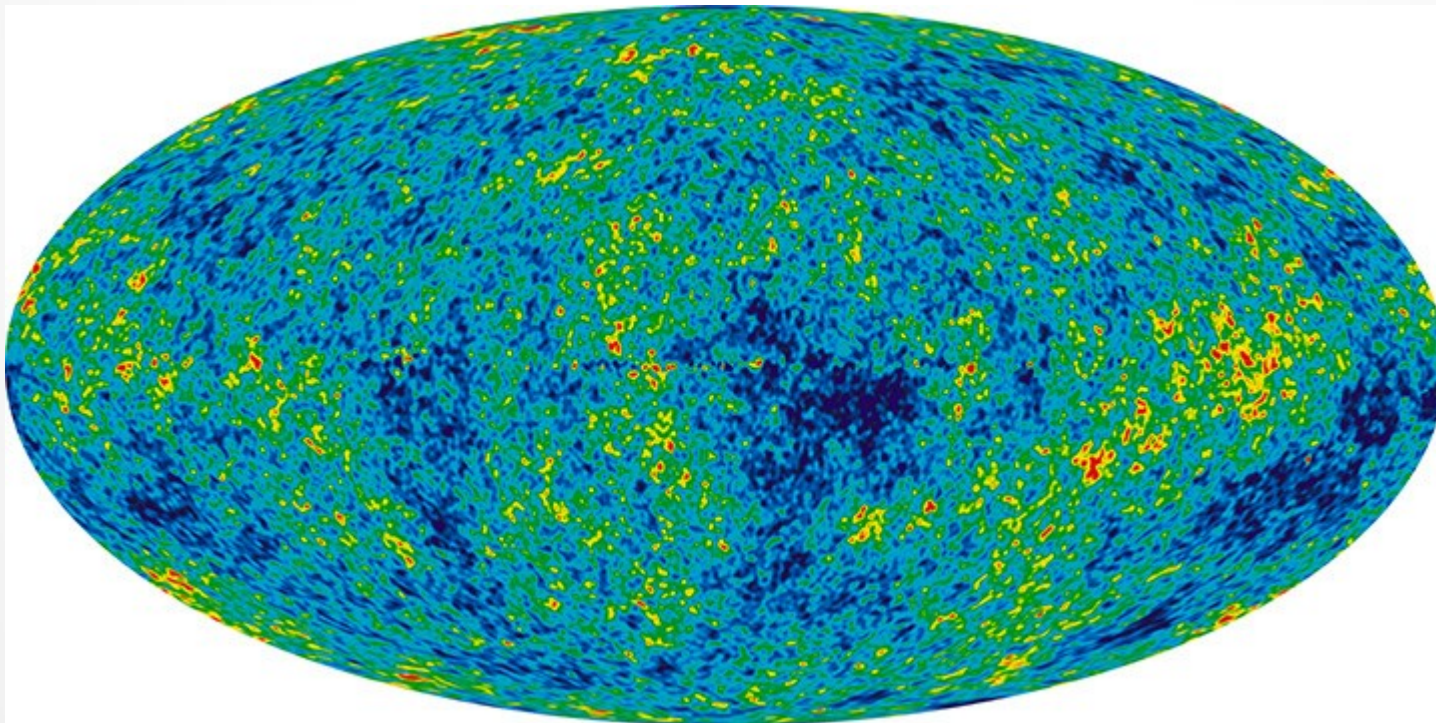
Similar result from ATIC

***Leads to an interesting idea**



WIMP “Miracle”

Currently observed dark matter density is consistent with massive particle production in Big Bang theory with WIMPS of 100 GeV mass and weak interaction cross-section (~ 1 pb).



Interesting Idea

A Theory of Dark Matter

Nima Arkani-Hamed,¹ Douglas P. Finkbeiner,² Tracy R. Slatyer,³ and Neal Weiner⁴

¹*School of Natural Sciences, Institute for Advanced Study, Princeton, NJ 08540, USA*

²*Harvard-Smithsonian Center for Astrophysics, 60 Garden St., Cambridge, MA 02138, USA*

³*Physics Department, Harvard University, Cambridge, MA 02138, USA*

⁴*Center for Cosmology and Particle Physics, Department of Physics,
New York University, New York, NY 10003, USA*

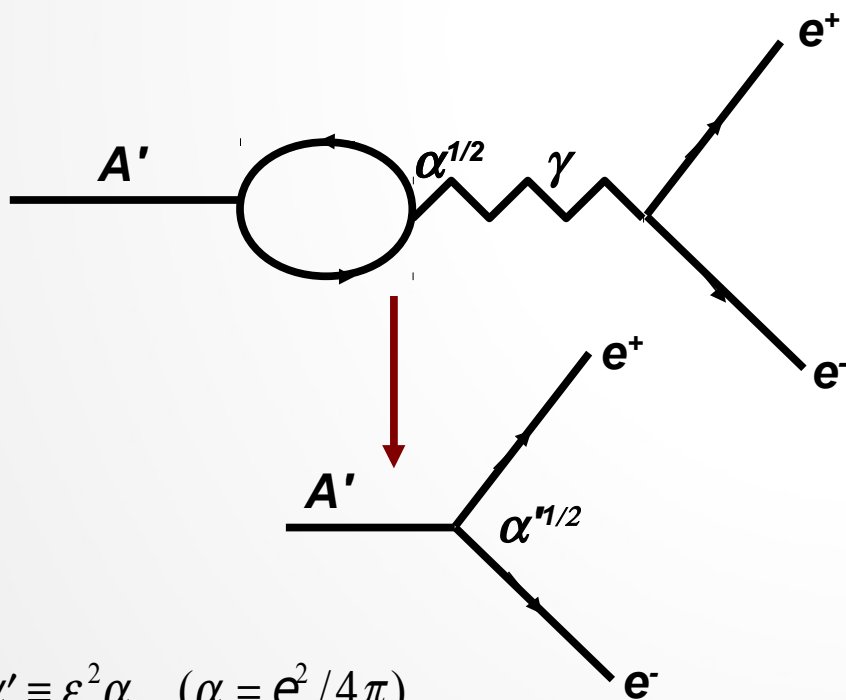
(Dated: January 20, 2009)

[arXiv:0810.0173v3](https://arxiv.org/abs/0810.0173v3)

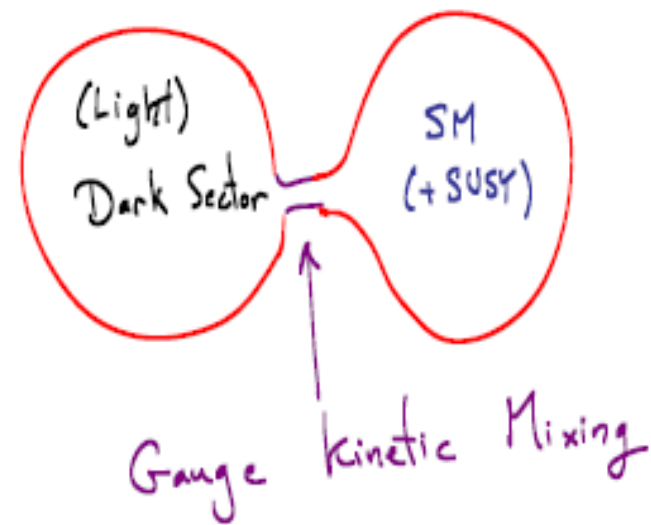
*The idea for a dark sector was proposed in 1986
([B. Holdom Phys. Lett. B 166, 196-198; 1986](#))
But remained largely unexplored.

Interesting Idea: Gist/Pith

- PAMELA data \rightarrow WIMP Mass $M_\chi \sim 500\text{-}800$ GeV (reinforced by WMAP “Haze”)
- Upper limits on hadronic channels from p^- limits
- These can imply possibility of hidden/”dark” force, $\lambda_{\text{Compton}} \sim m_{A'}^{-1} \sim 1 \text{ GeV}^{-1}$
- Predict carrier to be gauge boson A' , in mass range 10-100 MeV, coupling like a photon via kinetic mixing



$$\alpha' \equiv \varepsilon^2 \alpha, \quad (\alpha = e^2 / 4\pi)$$



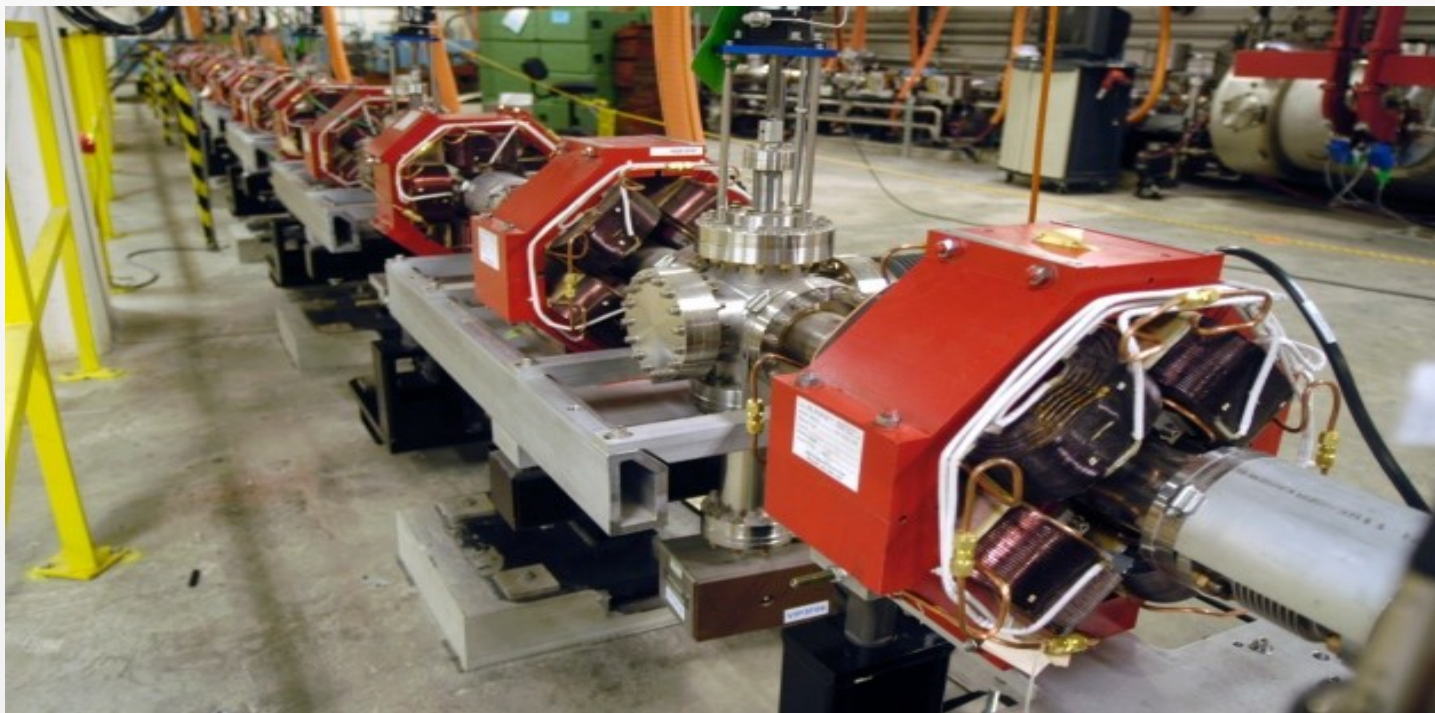
Arkani-Hamed

Interesting Idea: Testing

Beyond Observation: ***Provides a 3rd, independent study.**

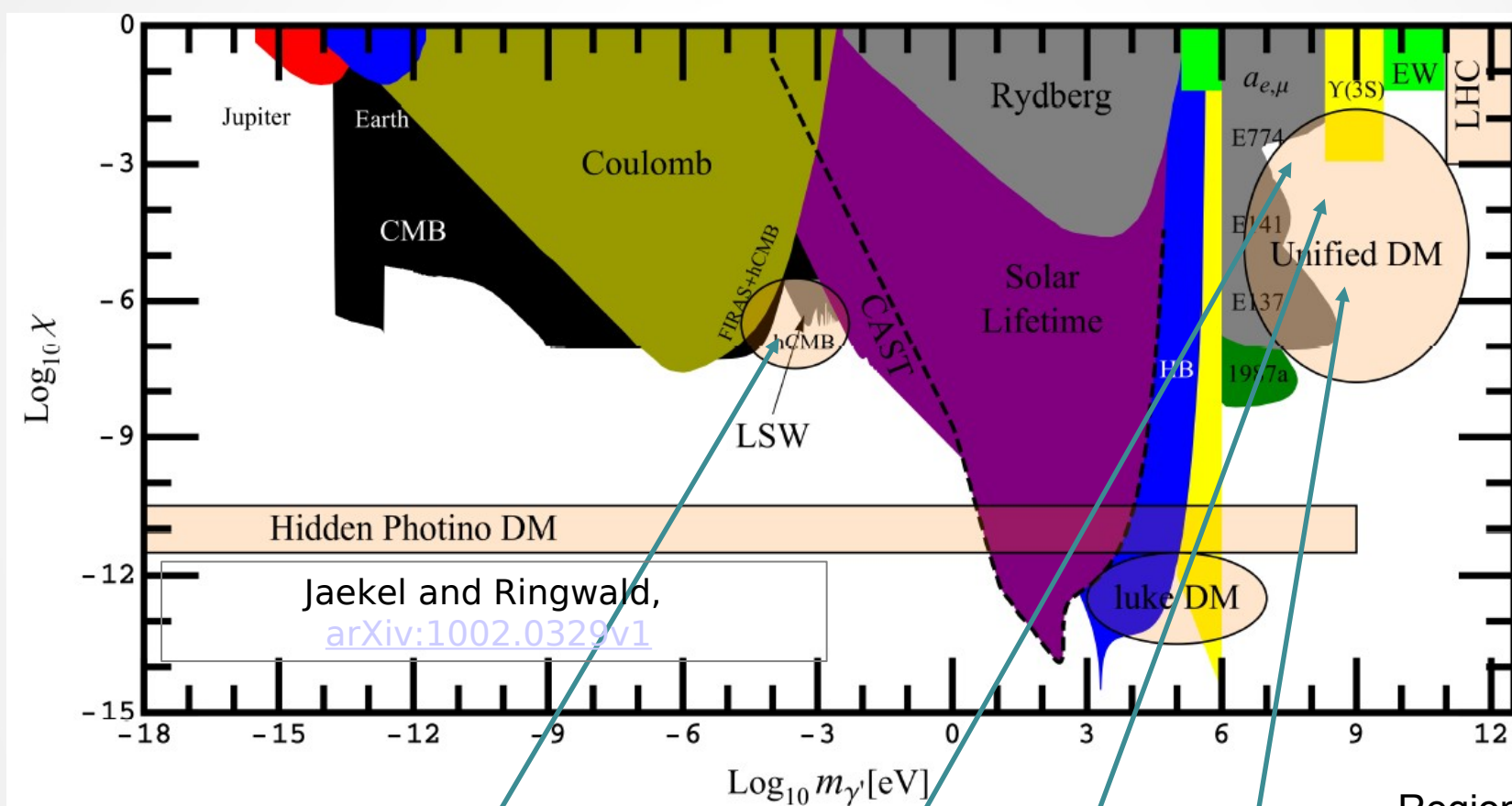
- Hadron Colliders; High Energy, Low Luminosity
- e^+e^- Colliders; Higher/Lower Energy, Lower/Higher Luminosity (arXiv:0903.3941v2)
- **Fixed Target; Low Energy, High Luminosity** (arXiv:0906.0580v1)

<http://www.nature.com/news/physicists-hunt-for-dark-forces-1.10386>



Previously Explored Dark Matter Regions

χ (kinetic mixing) vs. $m_{\gamma'}$ (mass of Dark Matter mediator)



**JLAB BSM
Experiments:**

LIPSS

DARKLIGHT

APEX

HPS

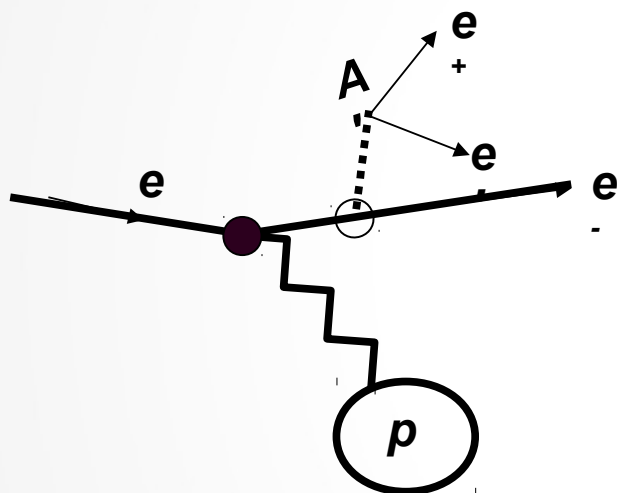
Regions not yet
explored, but
reachable

DarkLight: Goal



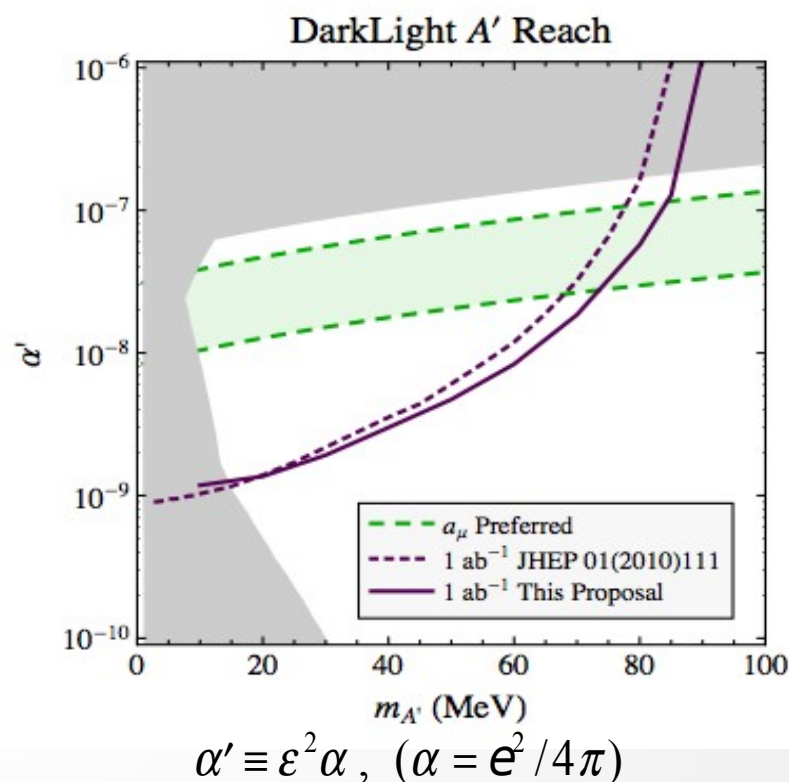
Detecting **A** Resonance **K**inematically with
electrons Incident on a **G**aseous **H**ydrogen **T**arget

A Search for new light bosons using the Jefferson Lab FEL facility.



Goal: Explore e^+e^- invariant mass spectrum using
the process $e^- + p \rightarrow e^- + p + e^- + e^+$
High Intensity, Low Energy Electron Beam
Using JLab's FEL on Thick Hydrogen Gas Target

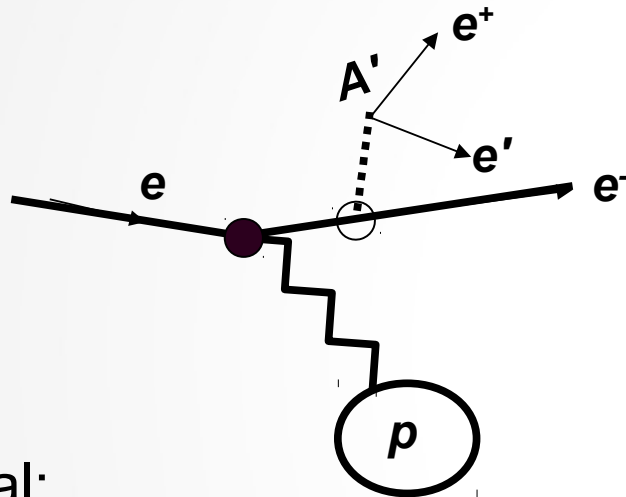
=> Luminosity: 1 ab⁻¹/month
(Formal request ~60 days)



“Dark Force Detection in Low Energy e-p Collisions”
[Freytsis, Ovanesyan, Thalar: arXiv:0909.2862 (JHEP 1001;111)]

DarkLight: Physics Processes

For $\alpha' \sim 10^{-8}$ the expected signal is 10^{-4} of the irreducible QED background:



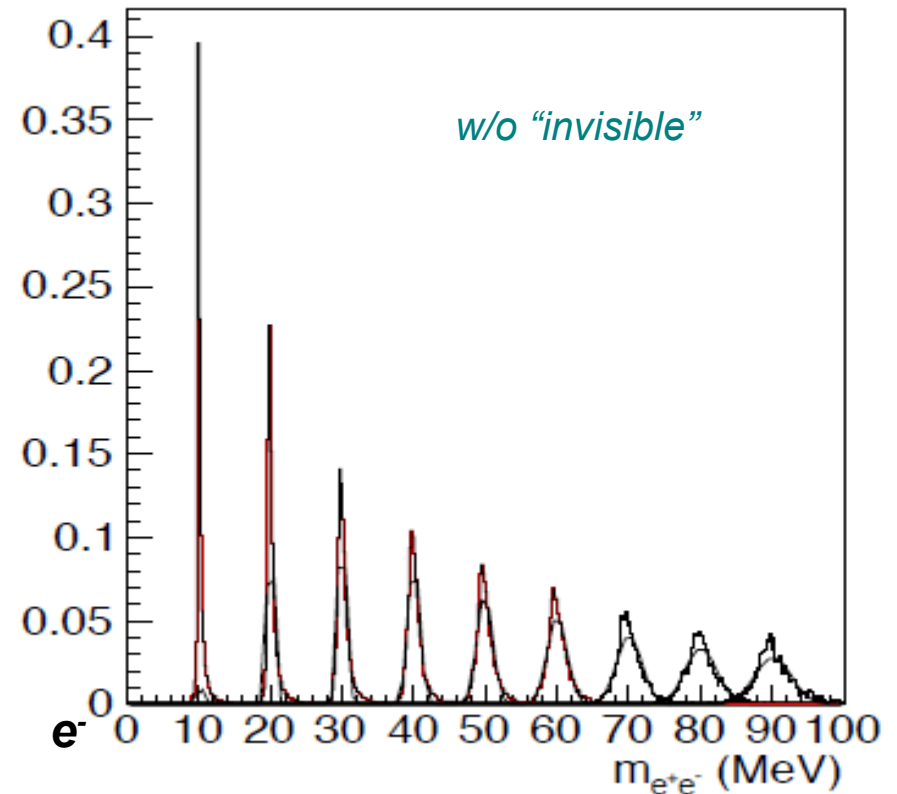
Signal:

$ep\gamma/epA'$, ep/epA' ("invisible")

Backgrounds:

ep/ep , $ep/ep\gamma$, $ep/ep\gamma\gamma$, $ep\gamma/ep\gamma$

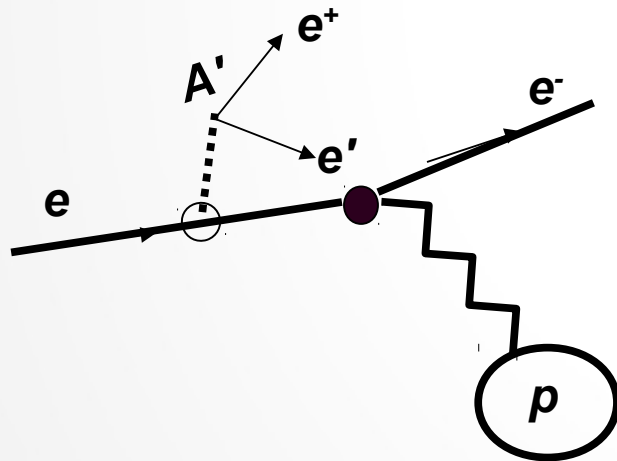
Reconstructed mass - all events



The experiment is basically a measurement of the QED background with 0.1ppm precision. *The detection of all 4 final states is essential.*

DarkLight: “Invisible”

- $ep \rightarrow epA'$ (“invisible”) observe only final state electron and proton
- Backgrounds' kinematics different enough that they can be controlled
- Requires photon tagging; scintillator

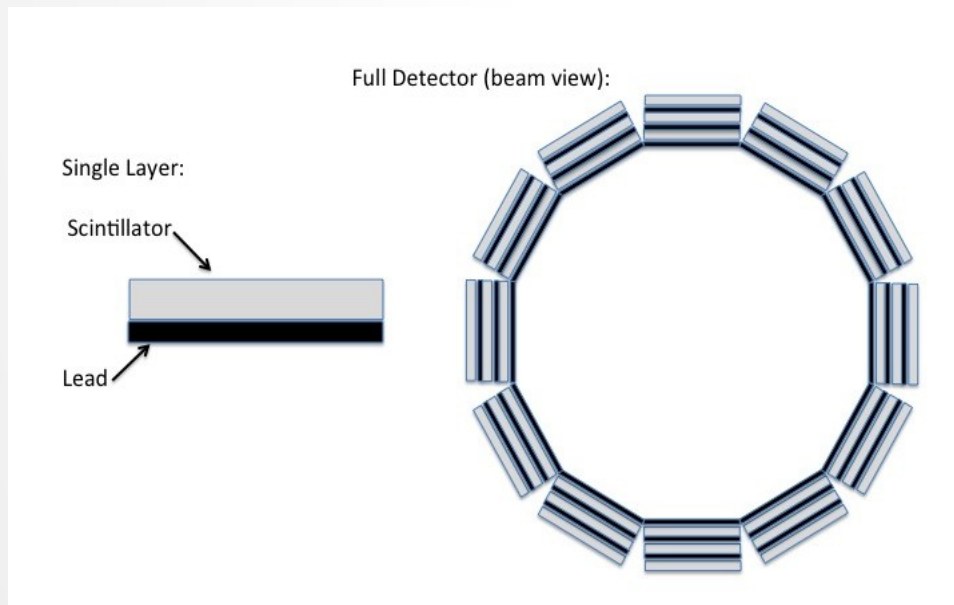


Kahn, Thaler arXiv:1209.0777

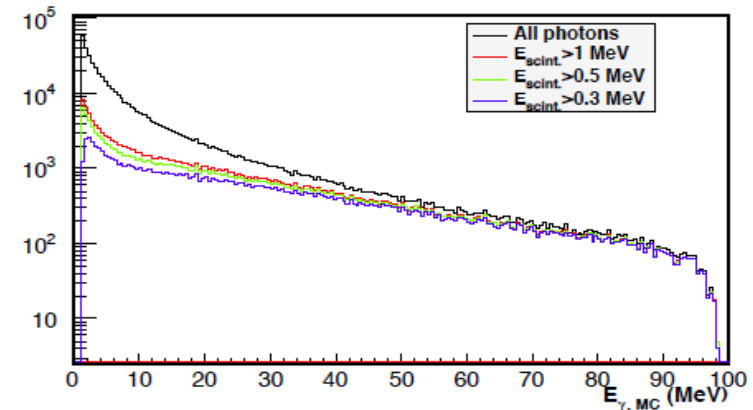
DarkLight: “Invisible”

Detector:

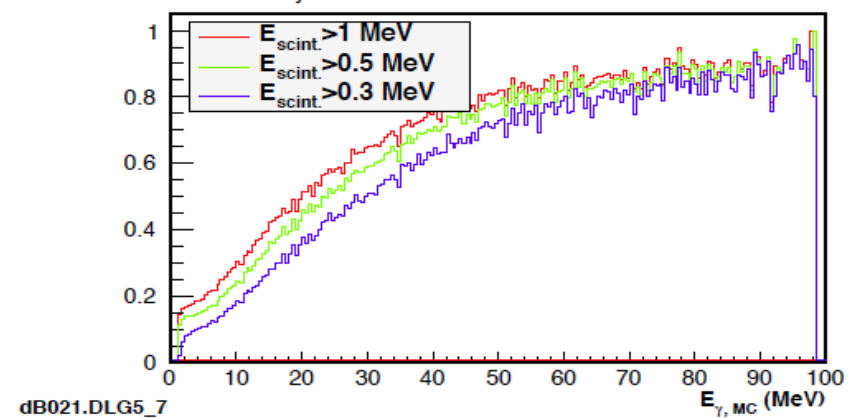
- Cylindrical Array
60 cm (diam) x 150 cm (length)
- Composed of 10 segments (10 cm wide)
- Segment = Pb (0.5 cm thick) + scintillator (1 cm thick) x 3+
- 3 layers => 90% efficiency.



Photons detected above threshold



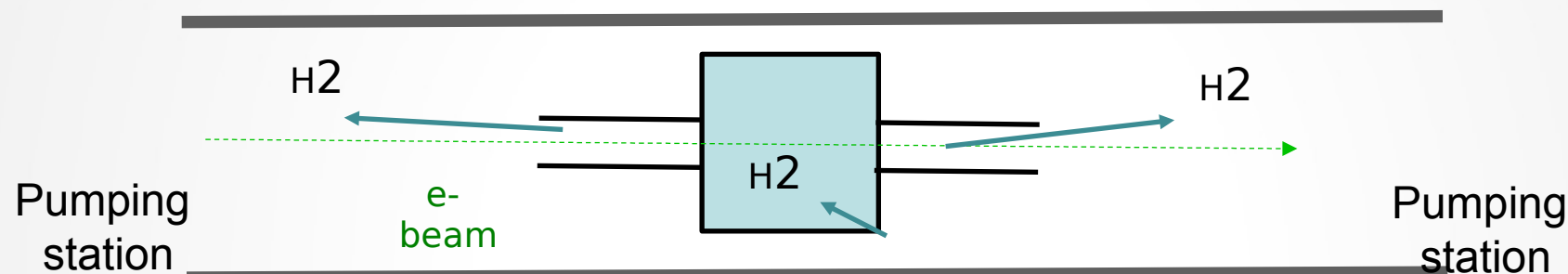
Photon detection efficiency



Jlab PAC Proposals

1. PAC 37: DarkLight Collaboration, PAC 37, November 30, 2010.

Early concept: electron beam scattering off H₂ in a windowless chamber.

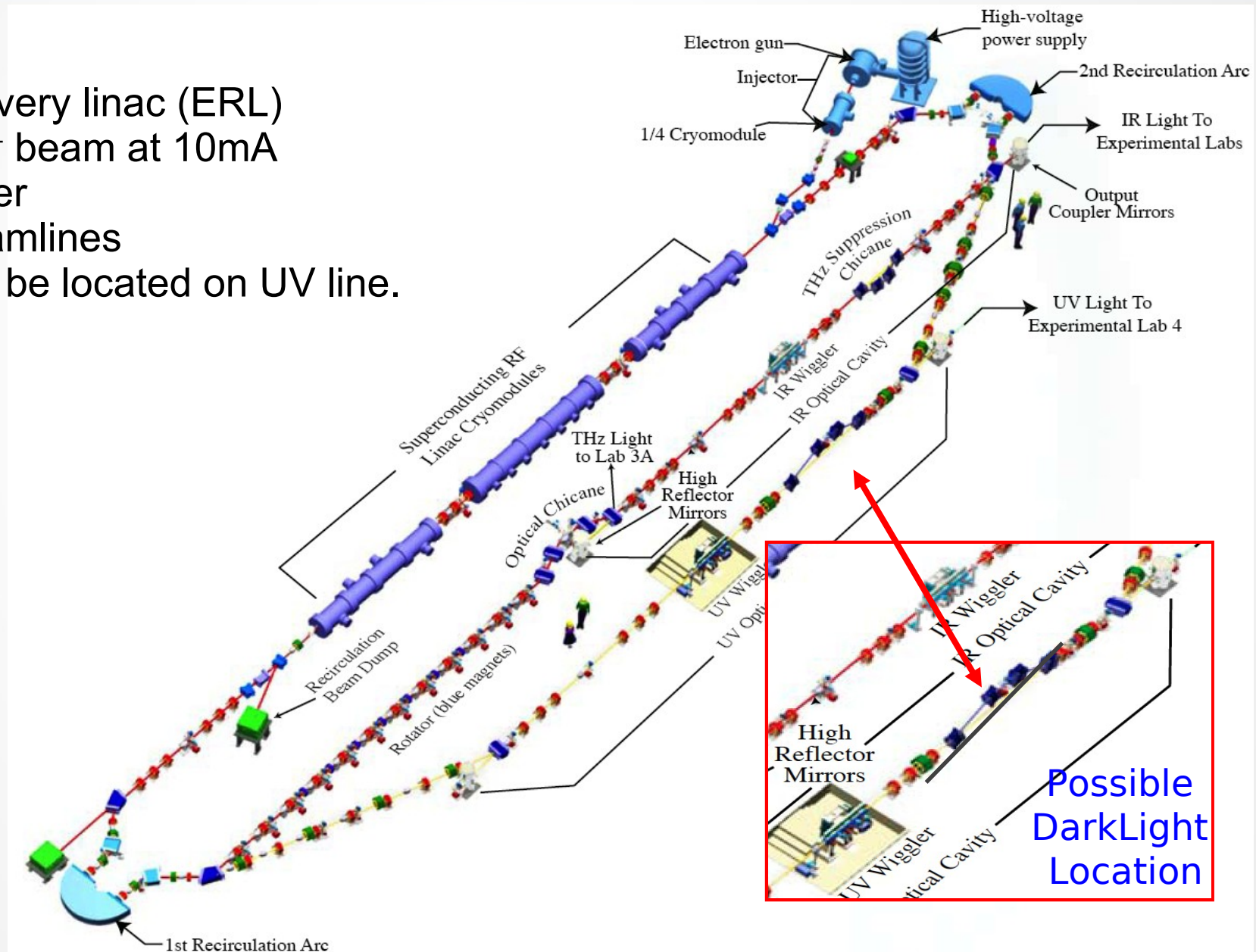


2. PAC 39: DarkLight Collaboration, PAC 39, May 4, 2012.

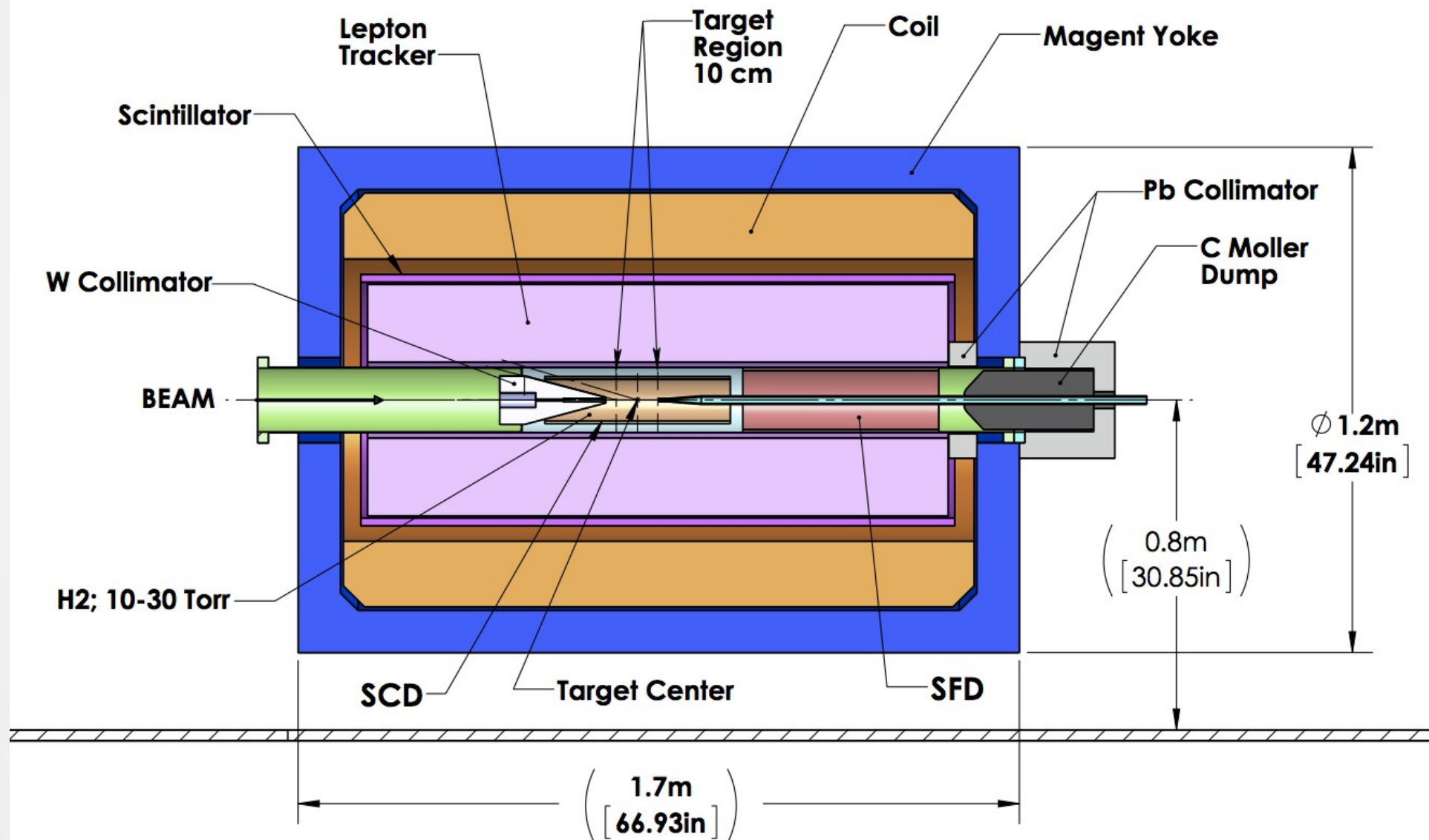
- Can DarkLight identify and shield against ambient FEL Vault background radiation.?
- Can the FEL beam be threaded through the proposed H₂ target?
- Can beam halo be managed?
- Are there any RF heating/effects on the target entrance/exit?

DarkLight: At FEL

- Energy recovery linac (ERL)
 - 100 MeV e^- beam at 10mA
 - 1 MW power
- IR & UV Beamlines
- DarkLight to be located on UV line.



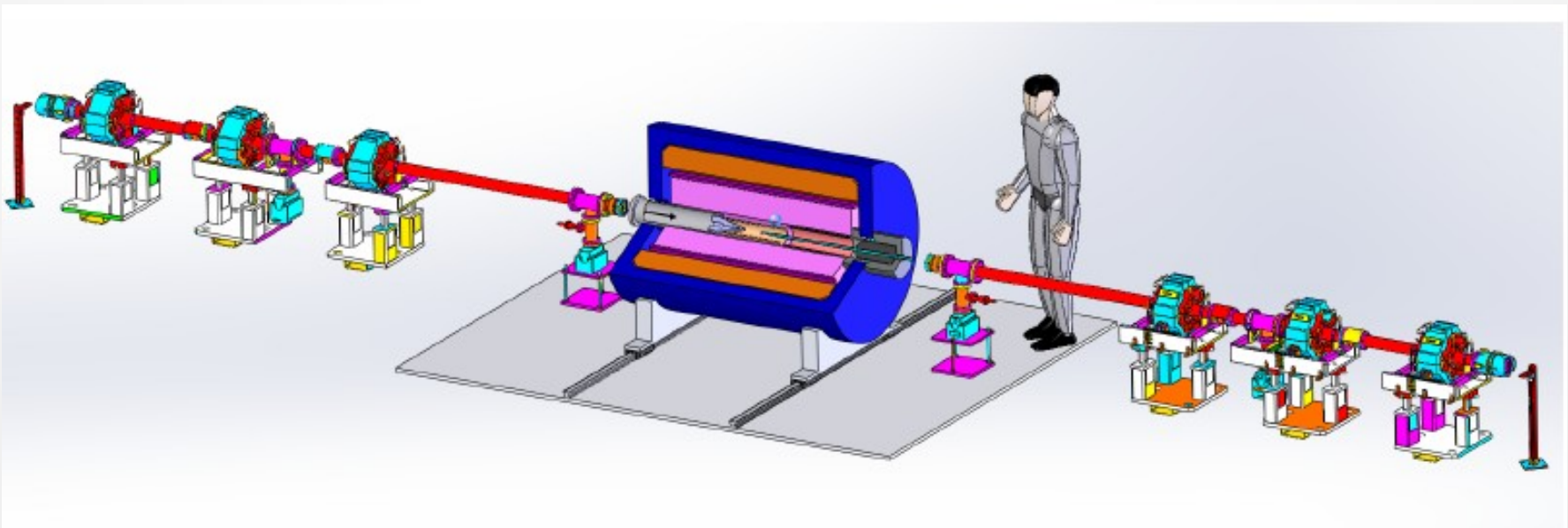
DarkLight: Schematic



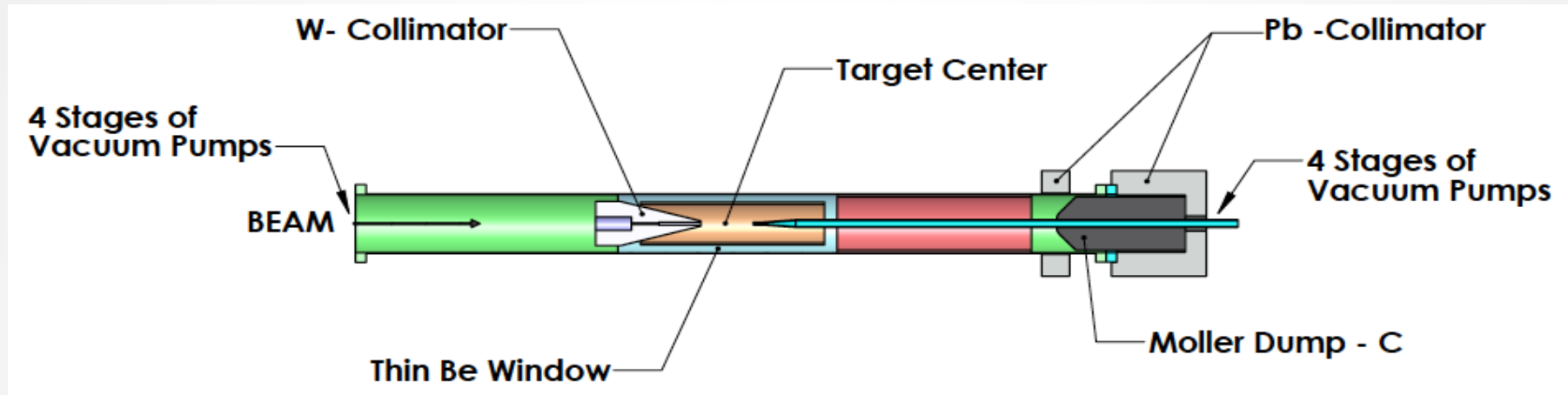
DarkLight: Components

DarkLight has 4 primary components:

- Target – Differentially pumped hydrogen gas target $10^{19}/\text{cm}^2$, 10 cm long.
- Silicon proton detector - ~ 3.5 cm from beam, single layer of silicon micro-strip detector. Measure energy and angle of recoil proton.
- Lepton tracker – 10-25 cm radius TPC, based on PANDA design.
- Magnet – Solenoid provides 0.5 T B-field to focus Moller e^- and measure lepton momentum and direction.



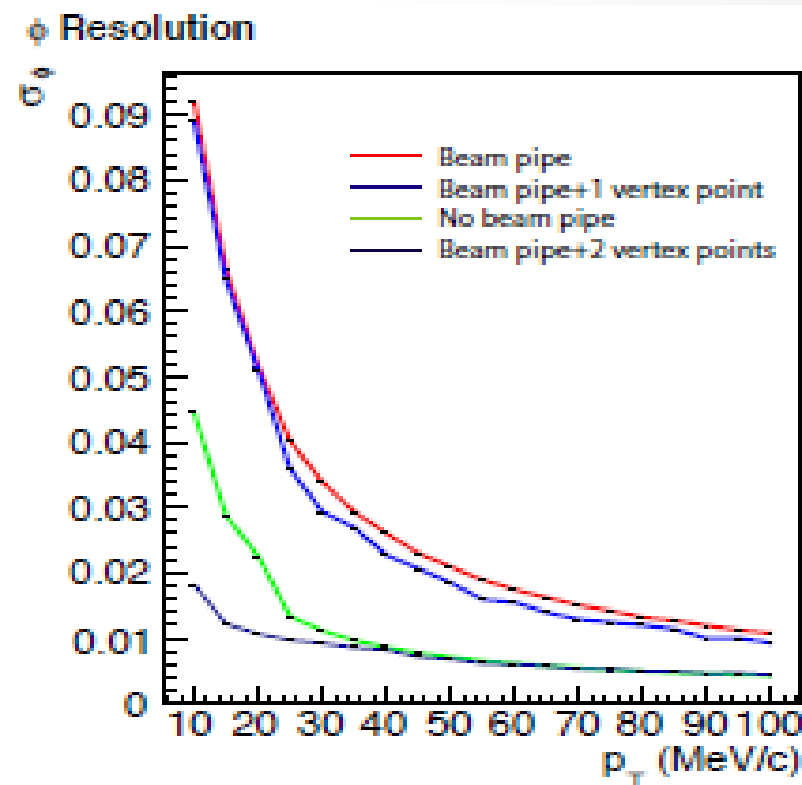
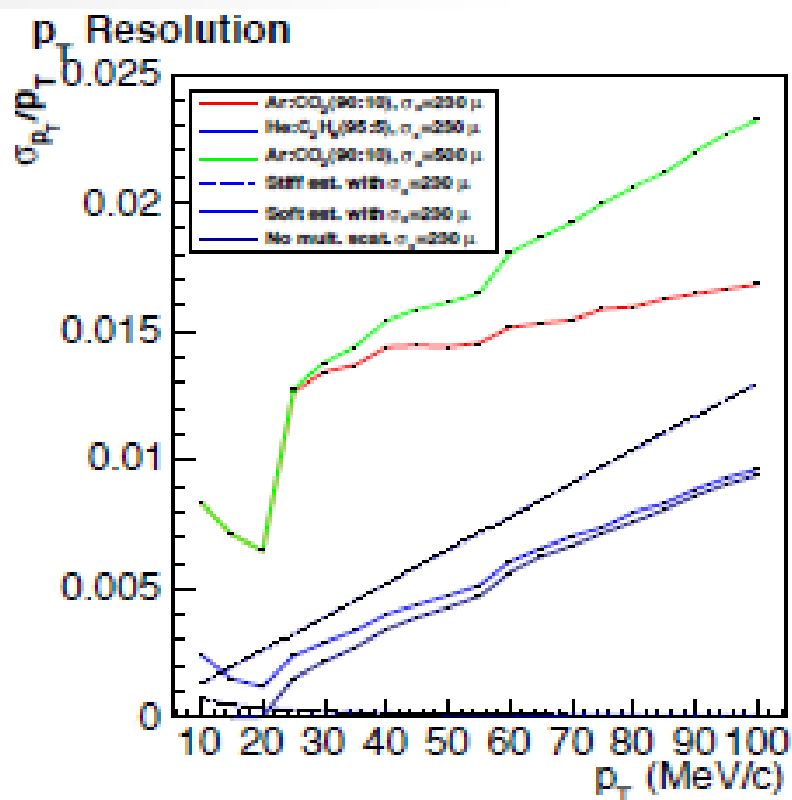
DarkLight: Target



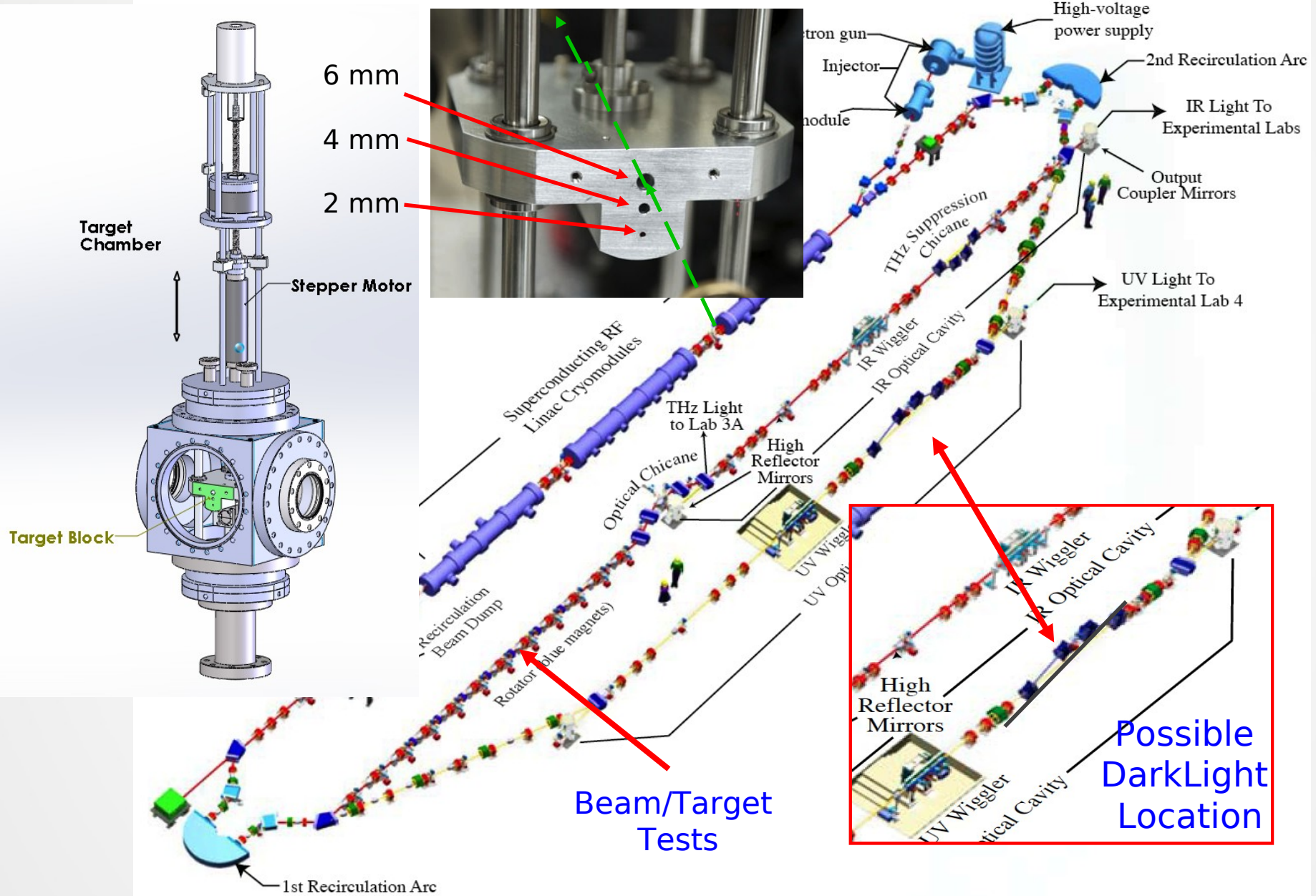
- Hydrogen target realized by flowing gas through narrow apertures
- Aperture diameter: 2 mm
- Aperture length: 50 mm
- Thickness: 10^{19} Hydrogen atoms cm^{-2}
- Flow rate: $24 \text{ Torr-liter s}^{-1}$
- Viscous subsonic flow regime
- Multiple stages of differential pumping required
- Plasma windows under consideration

DarkLight: Lepton Tracker

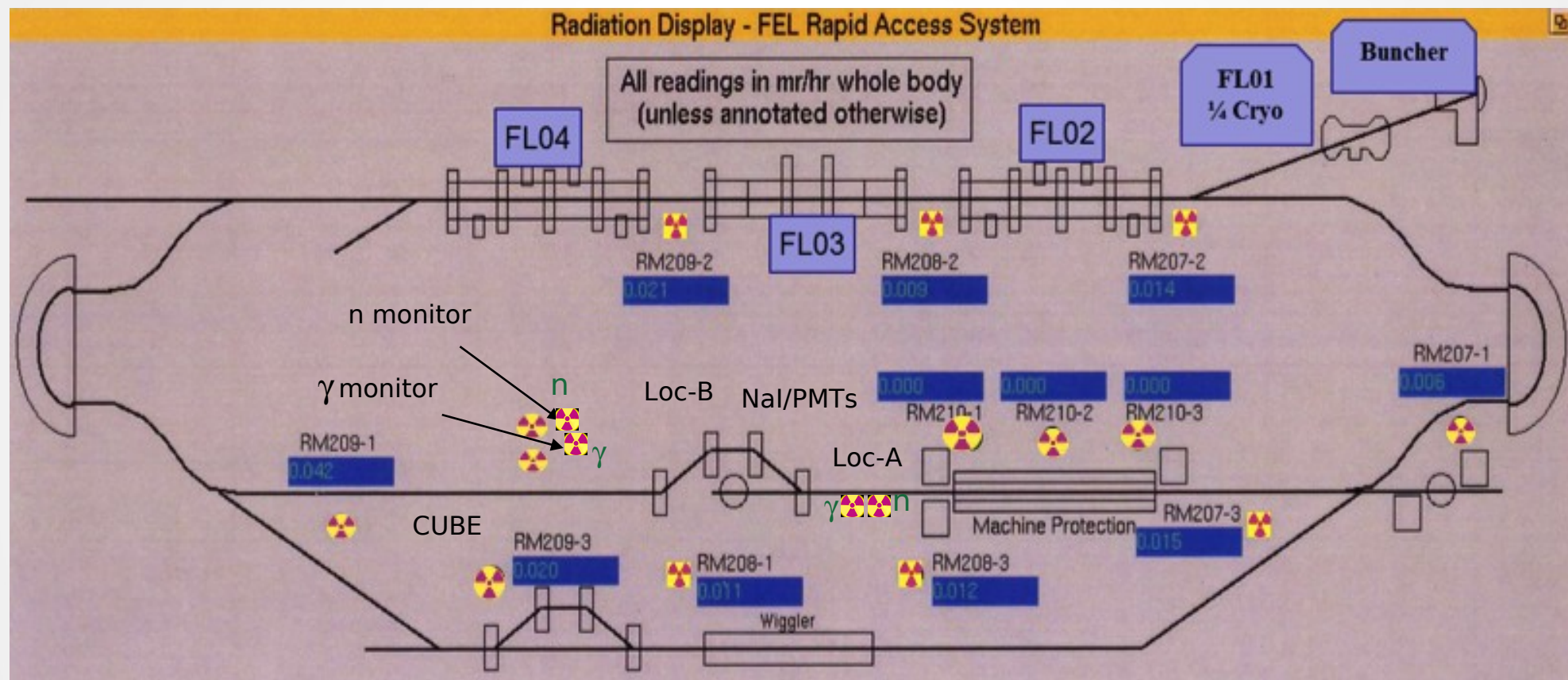
- Similar to PANDA TPC or STAR Forward Tracker.
- Gas Ar(Ne)/CO₂ (90/10)
- $\sim 10^4$ channels
- Triple GEM; gain ~ 10
- Drift length 725 mm
- Inner/Outer diameter 105/300 mm



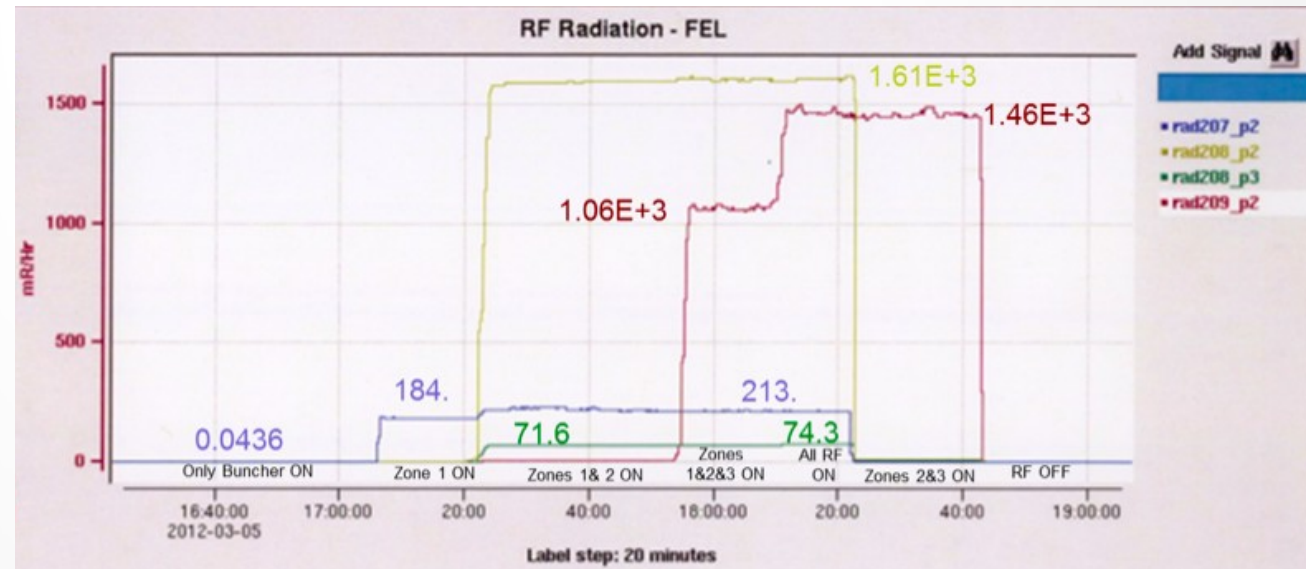
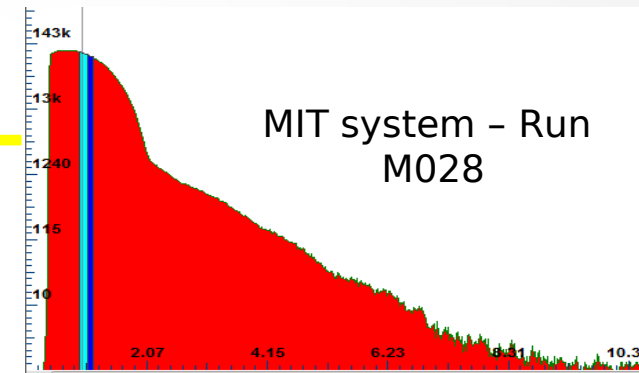
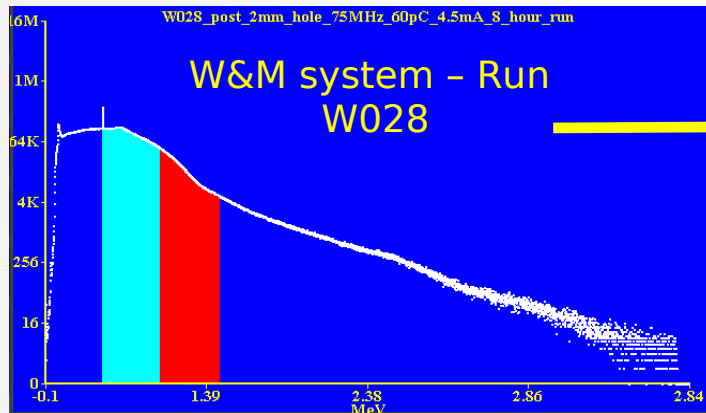
DarkLight: Beam Tests



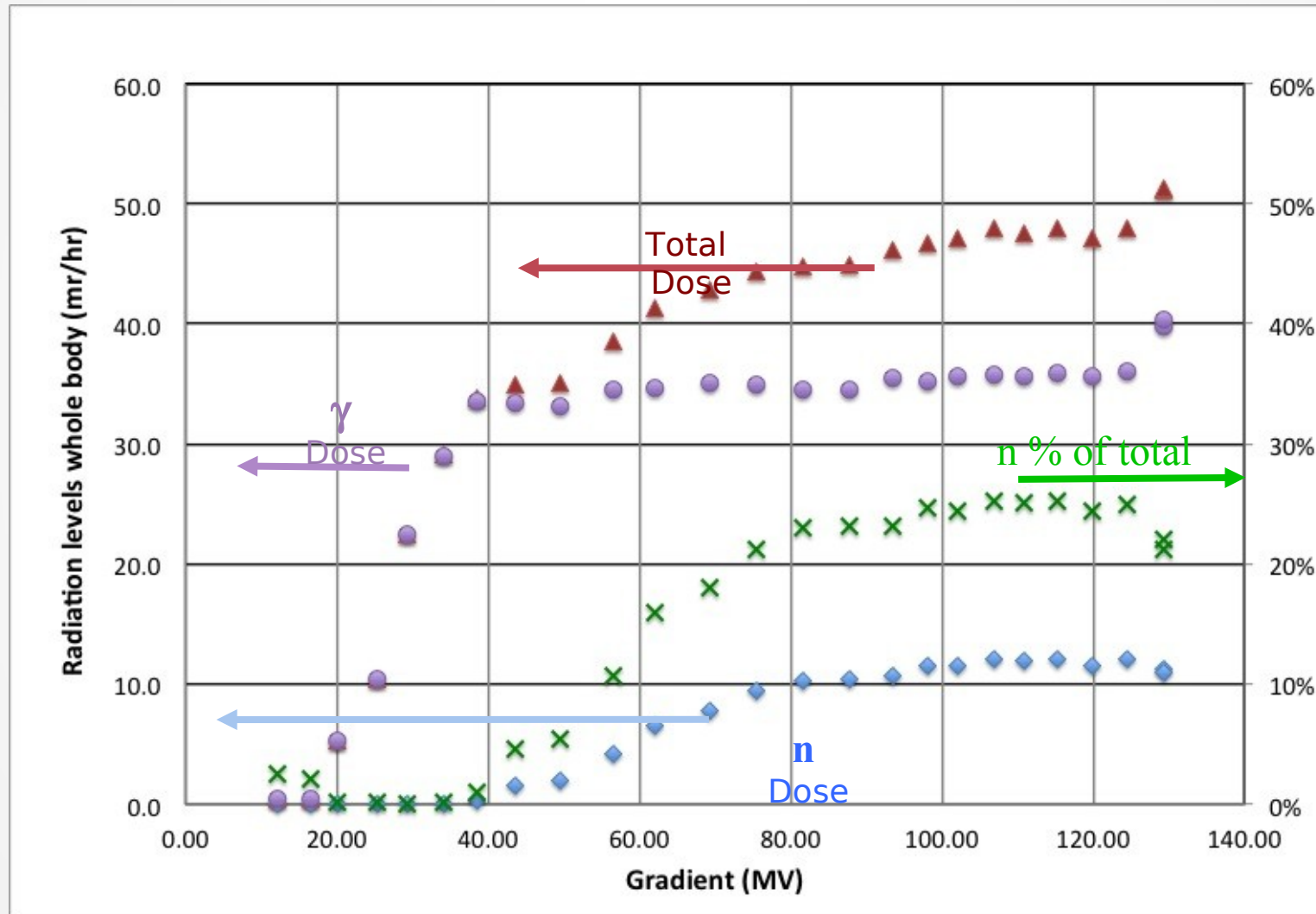
FEL Rapid Access System



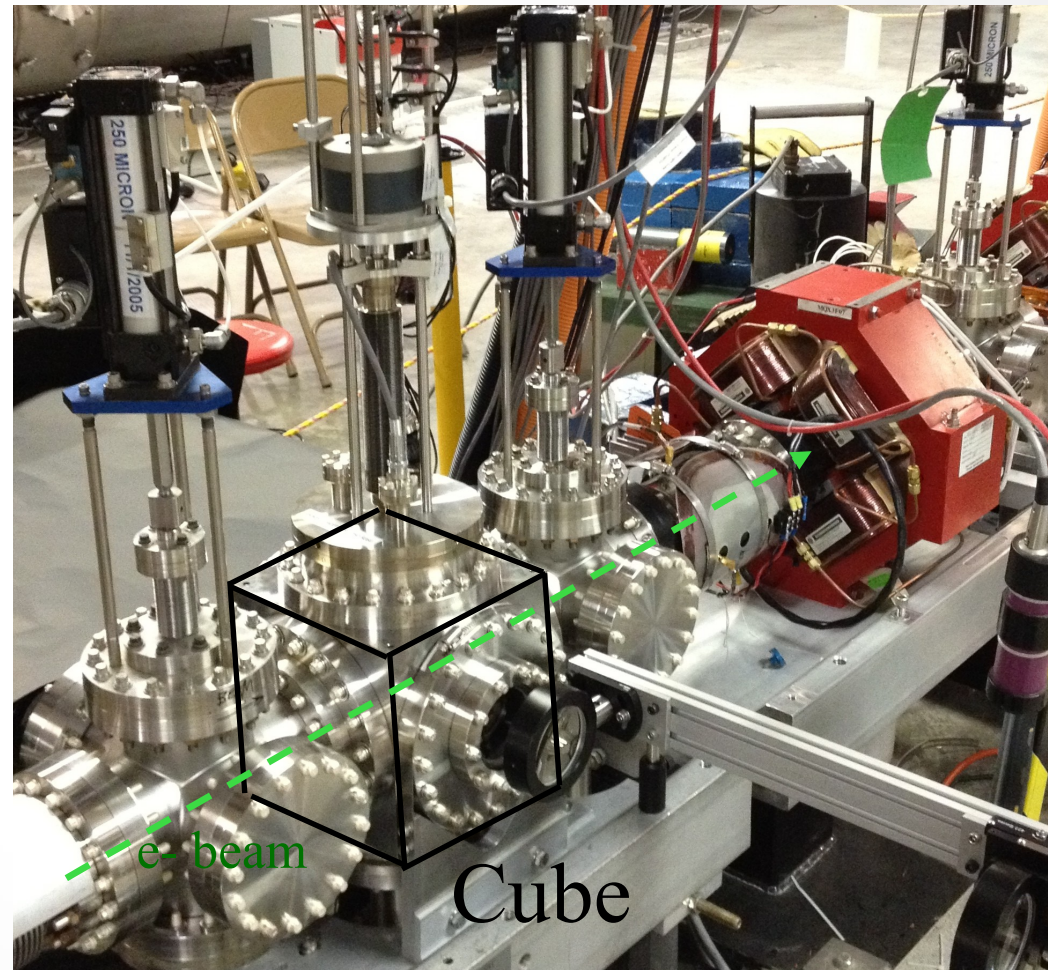
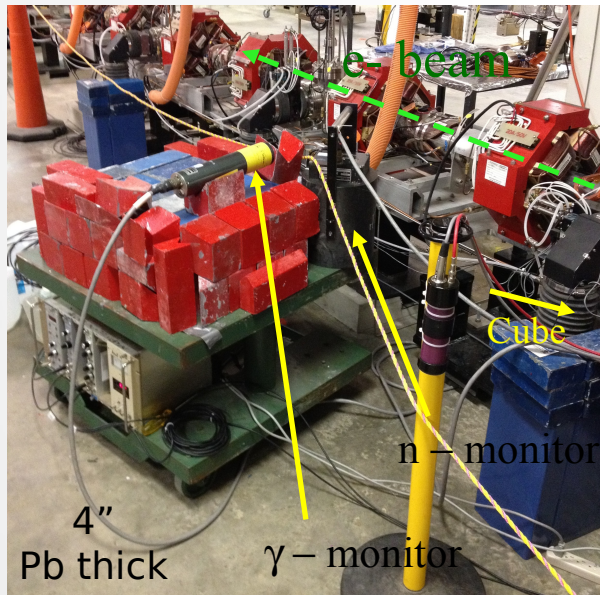
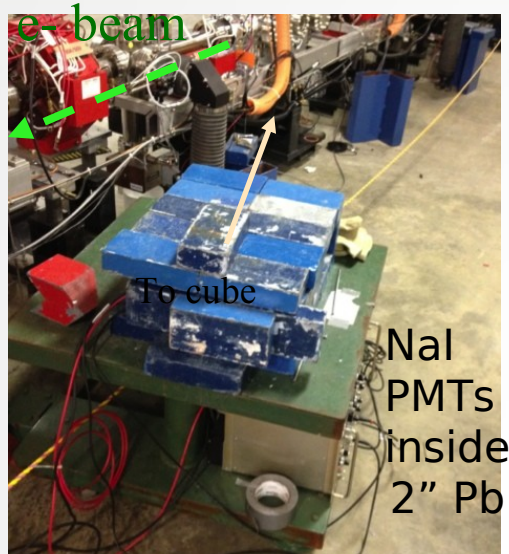
FEL Vault Background Radiation



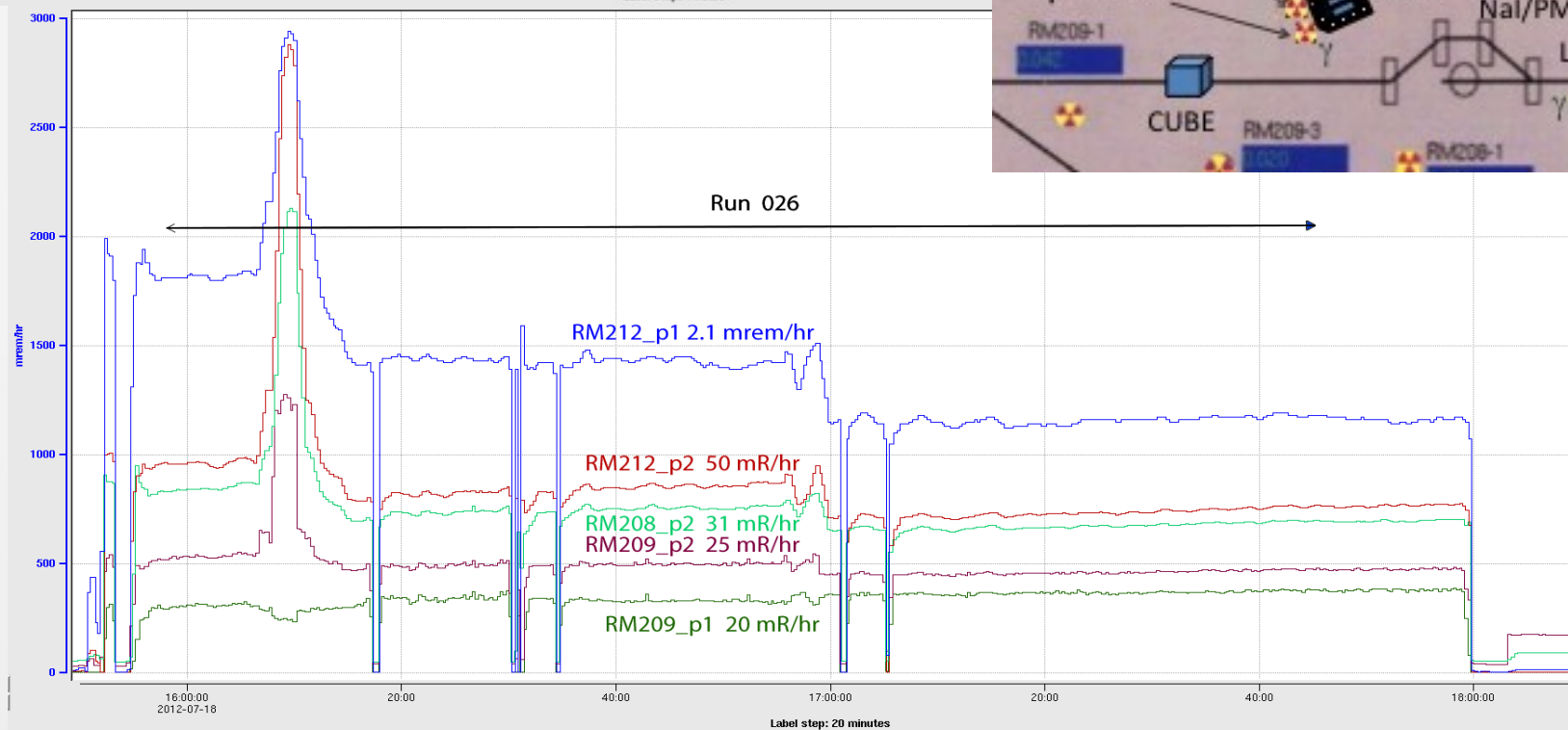
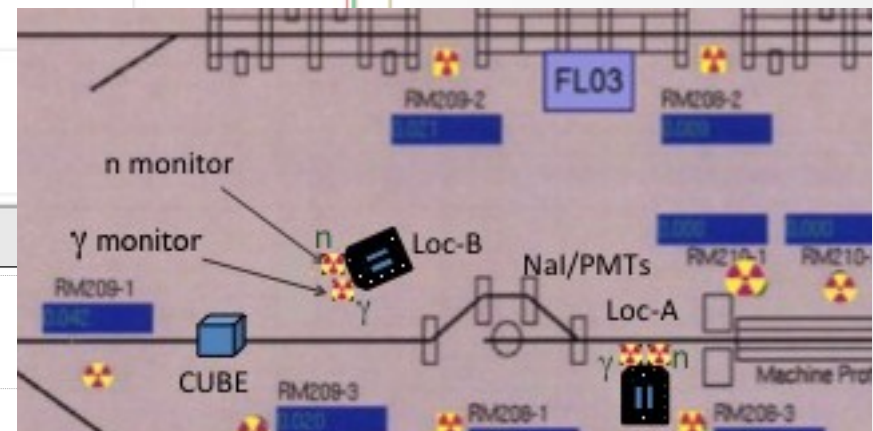
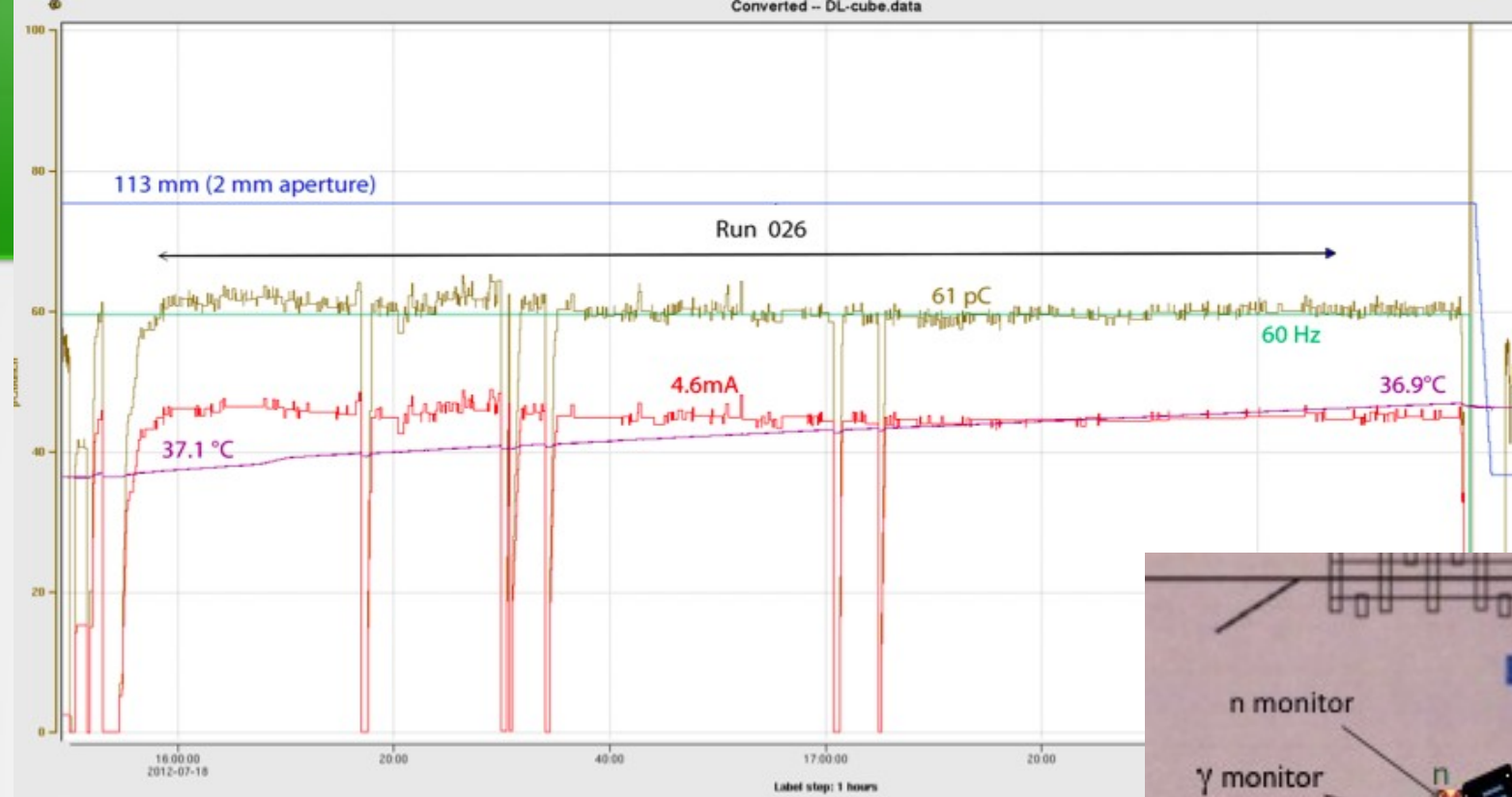
FEL Vault Radiation Levels vs. Total RF Gradient



FEL Beam-Target Tests & Radiation Measurements



DL Run 26 Location B



Beam Tests Results

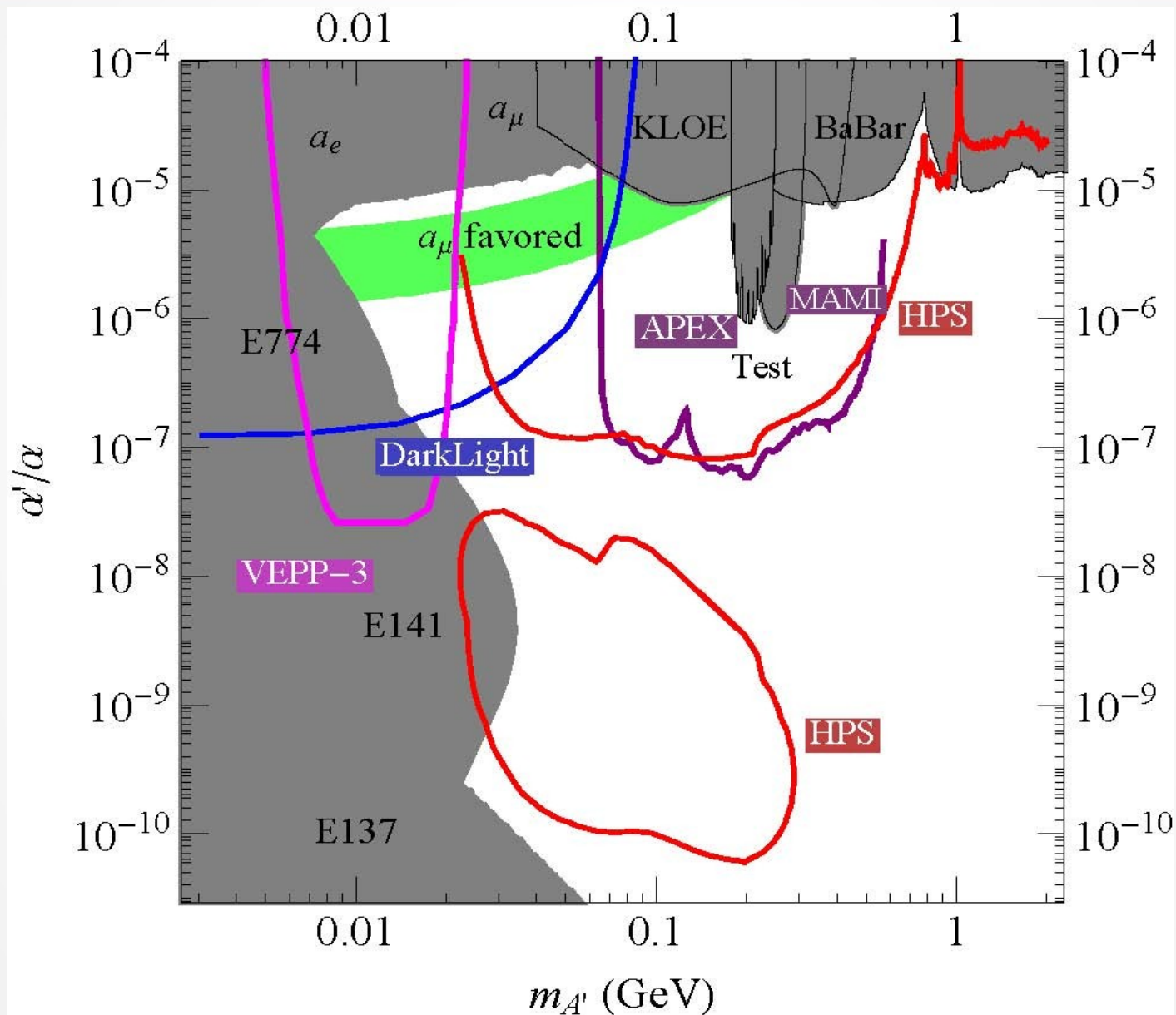
A test e^- beam 100 MeV, 4.5 mA (450 kW power) was successfully transmitted through a 2mm hole, 12.7 cm long, with max loss of 7 ppm for 7 hours. This showed that

- e^- beam bunch **CAN** be threaded through a 12.7 cm long, 2 mm hole.
- Halo **CAN** be minimized.
- The FEL has the stability required for a successful DarkLight experiment.
- Radiation in the vault is manageable.

Possible Timeline

Focus	Year	2012	2013	2014	2015	2016
FEL beam & Radiation limits						
Finalize Design Secure funding						
Technical Review Start Construction						
Detector Commissioning						
DarkLight data taking begins						

Projected Results



Collaboration

Spokespersons: Peter Fisher and Richard Milner

J. Balewski, J. Bernauer, W. Bertozzi, J. Bessuille, B. Buck, R. Cowan, K. Dow, C. Epstein, P. Fisher ², S. Gilad, E. Ihloff, Y. Kahn, A. Kelleher, J. Kelsey, R. Milner, C. Moran, L. Ou, R. Russell, B. Schmookler, J. Thaler, C. Tschalaer, C. Vidal, A. Winnebeck
Laboratory for Nuclear Science, Massachusetts Institute of Technology, Cambridge, MA 02139, USA and the Bates Research and Engineering Center, Middleton MA 01949

S. Benson, C. Gould, G. Biallas, J. Boyce, J. Coleman, D. Douglas, R. Ent, P. Evtushenko, H. C. Fenker, J. Gubeli, F. Hannon, J. Huang, K. Jordan, R. Legg, M. Marchlik, W. Moore, G. Neil, M. Shinn, C. Tennant, R. Walker, G. Williams, S. Zhang *Jefferson Lab, 12000 Jefferson Avenue, Newport News, VA 23606*

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Institute for Research in Electronics and Applied Physics University of Maryland, College Park, MD

R. Alarcon, R. Dipert
Physics Department, Arizona State University, Tempe, AZ

G. Ovanessian *Los Alamos National Laboratory, Los Alamos NM*

T. Gunter, N. Kalantarians, M. Kohl
Physics Dept., Hampton University, Hampton, VA 23668 and Jefferson Lab, 12000 Jefferson Avenue, Newport News, VA 23606

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D. S. Gunarathne, C. J. Martoff, D. L. Olvitt, B. Surrow, X. Li
Physics Dept., Temple University, Philadelphia, PA 19122

E. Long *Physics Dept., Kent State University, Kent, OH, 44242*

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University Bonn, D - 53115 Bonn Germany

K. Brinkmann, H. Zaunick
II. Physikalisches Institut Justus-Liebig-Universitt Giessen, D-35392 Giessen Germany

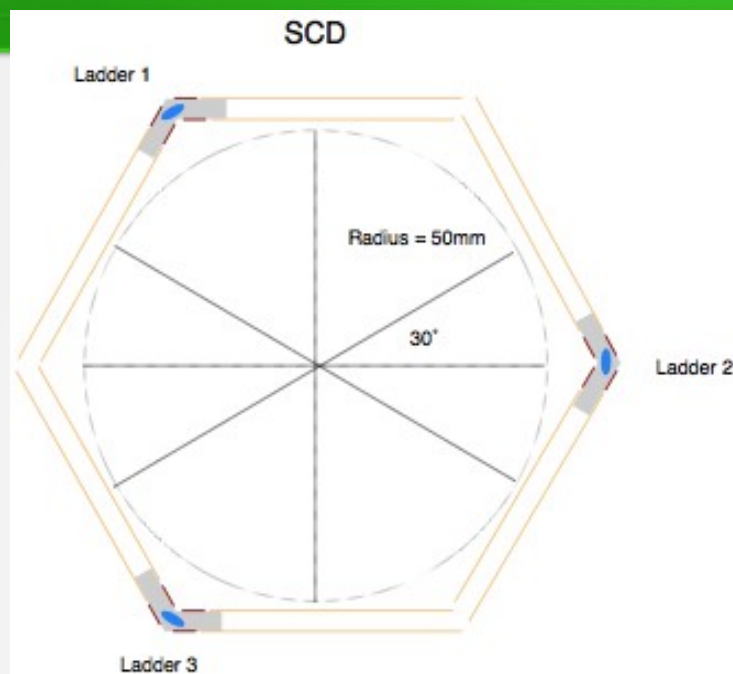
W.J.Kossler
Physics Dept., College of William and Mary, Williamsburg VA 23185

Summary and Conclusion

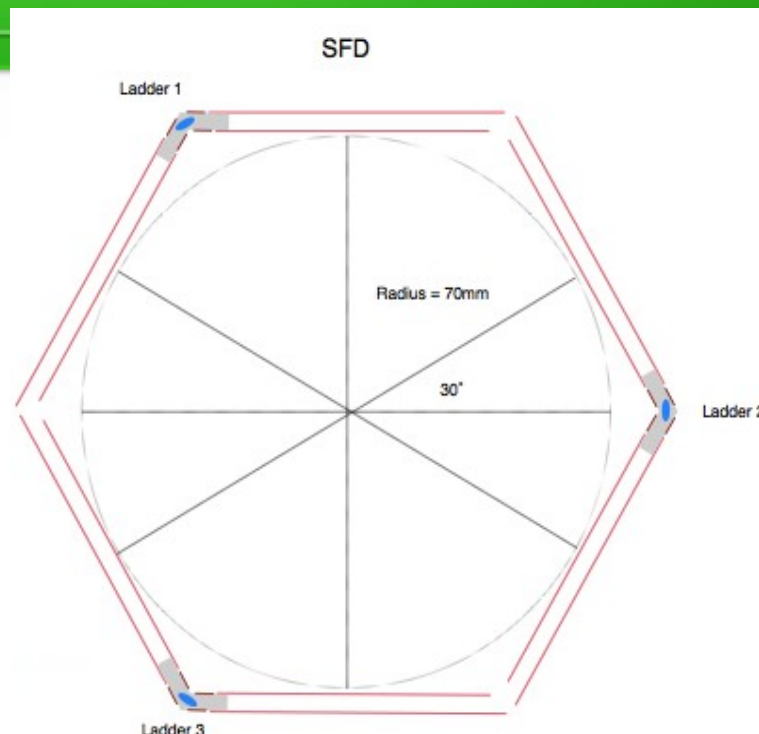
- Fairly recent theoretical developments suggest interaction between dark matter and SM via gauge boson.
- Lepton scattering, fixed-target provides yet another method for studying dark matter.
- DarkLight intends to use the FEL beam (~ 1 mA, 100 MeV) incident on a H₂ gas target. Collect 1/ab in ~ 60 days of beam time.
- High acceptance detector inside a 0.5T solenoid: Si-strip recoil detector, TPC for lepton tracking.
- Satisfied condition for PAC to recommend approval.

Support Slides

DarkLight: Central & Forward Detectors



- Radius: 50mm
- Φ -coverage: 360°
- θ -coverage: 17° - 163°
- Number of ladders: 3
- Ladder length: 430mm
- Sensor dimensions: 56.5mm X 60.0mm / 52.5mm X 60.0mm
- Sensor thickness: 300 μ m
- Total number of sensors: 84
- Power dissipation: 0.3W per chip / 50W per ladder
- Radiation tolerance of sensors: 1MRad
- Radiation tolerance of readout chip: \gg 1MRad



- Radius: 70mm
- Φ -coverage: 360°
- θ -coverage: 6.1° - 19°
- Number of ladders: 3
- Ladder length: 458mm
- Sensor dimensions: 78.5mm X 64.0mm / 72.5mm X 64.0mm
- Number of sensors: 28 per ladder
- Sensor thickness: 300 μ m
- Total number of sensors: 84
- Power dissipation: 0.3W per chip / 50W per ladder
- Radiation tolerance of sensors: 1MRad
- Radiation tolerance of readout chip: \gg 1MRad

DarkLight Specs

$M_{A'}$ 1 MeV (< 1% Rad. Length)

Incident electron energy 100 MeV

Scattered lepton angle 25-165 deg

Scattered lepton energy 10-100 MeV

Recoil proton angle 6-163 deg

Recoil proton energy 1-6 MeV

Position: 250 μm

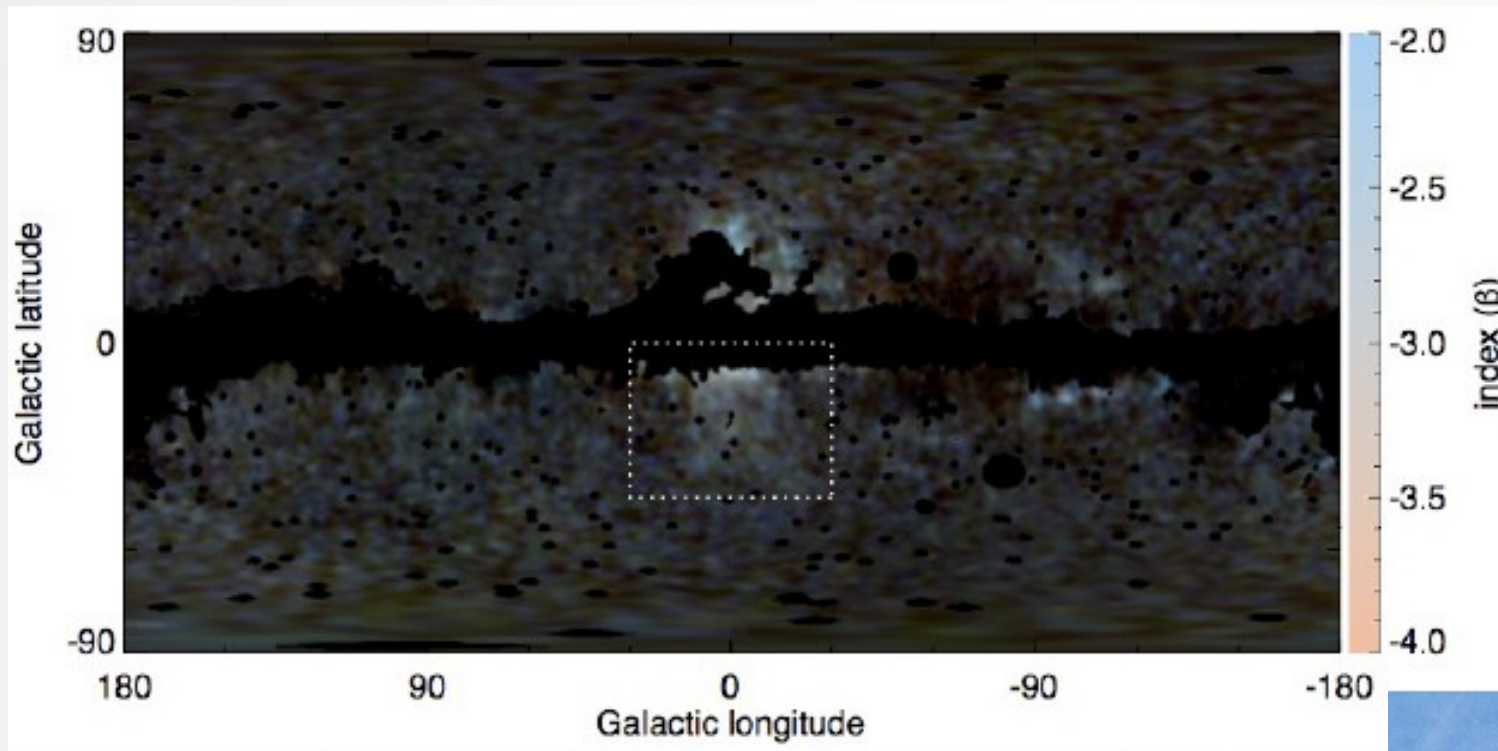
Elastic rate within acceptance 10 MHz

Trigger Rate \sim kHz

Jlab FEL Capabilities

	Near Term Capability, Dec. 2013	Full Capability	Internal Target (Near Term)
	external target	external target	internal target
E (MeV)	80-320	80-610	80-165
P _{max} (kW)	100	300	1650
I (mA)	0.31-1.25	0.5-3.75	10
f _{bunch} (MHz)	750 / 75	750 / 75	750 / 75
Q _{bunch} (pC)	1.67-0.4 / 16.7-4	5-0.67 / 50-6.7	13.5 / 135
$\epsilon_{\text{transverse}}$ (mm-mrad)	~1 / ~3	~2 / ~5	~3 / ~10
$\epsilon_{\text{longitudinal}}$ (keV-psec)	~5 / ~15	~10 / ~25	~15 / ~50
Polarization	No	Up to 600 μ A	No
	750 MHz drive laser; single F100	12 GeV RF drive; three F100s	12 GeV RF drive; three F100s

WMAP “Haze”



Dobler & Finkbeiner 2008;
New source of 10+ GeV e^+e^- in galactic center

Electrons spiraling in
B-fields create microwaves

