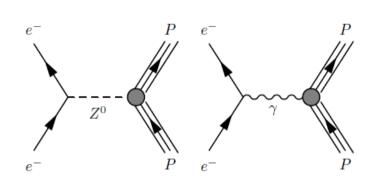
Parity Violating Electron Scattering at Jefferson Lab

Rakitha S. Beminiwattha Syracuse University

Outline

- Parity Violating Electron Scattering (PVES) overview
- Testing the Standard Model (SM) with PVES
 - Qweak, SoLID-PVDIS and MOLLER
- Nuclear structure physics with PVES
 - PREX/CREX
- PVES as a probe of nucleon structure
 - SoLID-PVDIS EMC proposal

Parity Violating Electron Scattering

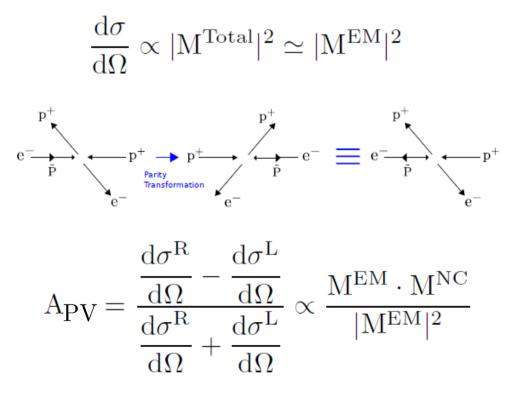


$$M^{EM} = \frac{4\pi\alpha}{Q^2} Q_\ell l^\mu J^{EM}_\mu$$
$$M^{NC} = \frac{-G_F}{2\sqrt{2}} (g_V^\ell l^\mu + g_A^\ell l^{\mu 5}) (J^{NC}_\mu + J^{NC}_{\mu 5})$$

Differential scattering cross section,

Due to PV nature of the neutral current, the differential cross section is dependent on the helicity of the electron

The difference in helicity correlated scattering cross section is known as the PV asymmetry,



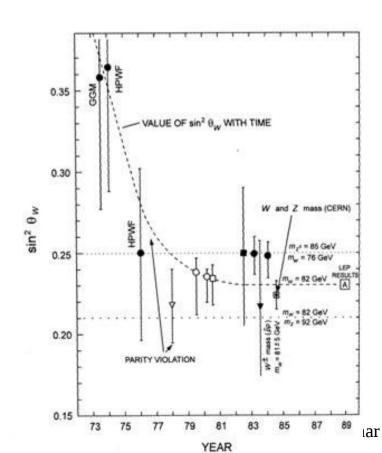
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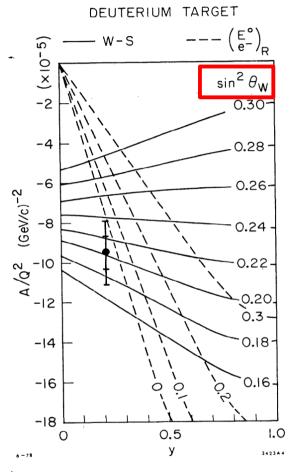
PVES Applications

- Testing the Standard Model (SM)
 - Qweak (e-p), MOLLER (e-e), SoLID-PVDIS (e-q) experiments
- Nuclear Structure
 - Neutron density measurements with PREX/CREX experiments (e-²⁰⁸Pb and e-⁴⁸Ca)
- Nucleon Structure
 - EMC with SoLID-PVDIS experiment using e-48Ca
 - Strangeness in proton (HAPPEX, G0 experiments) and etc.

PVES Historical Significance

- Confirmation of the EW SM from the first PVES experiment at SLAC by Prescott et. al.
- First measurement of parity-violation in the neutral weak current!
 - Which they found the weak mixing angle to be around 1/4 that amount to a small <u>axial vector(e)</u> X <u>vector(f)</u> weak neutral interaction!



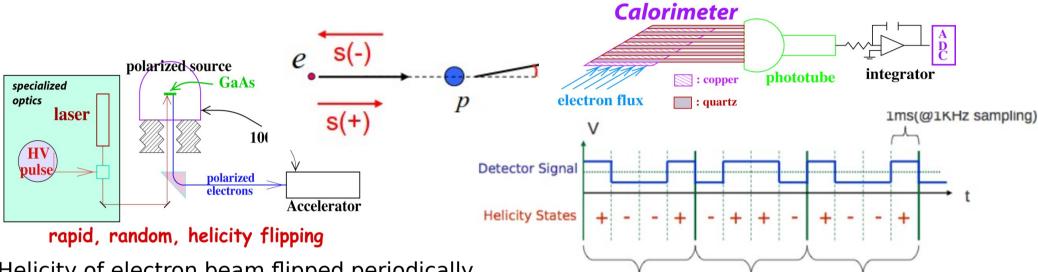


1st PVDIS at SLAC! first result in 1978: Prescott et al., PLB 77, 347 (1978) Prescott et al., PLB 84, 524 (1978)

Unique Nature of a PVES Experiment

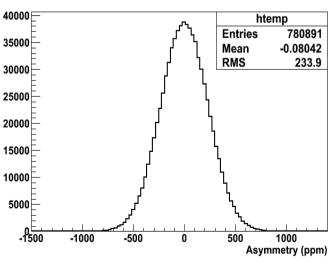
- The Injector + Accelerator + Apparatus or "The Whole Machine" becomes parts of the experiment
- Complete understanding of all the backgrounds is the key to successful PVES
- Monitor PVES asymmetries real-time to find issues and fix them
 - No second chance at offline after the experiment

How to Do A PVES Experiment



Helicity of electron beam flipped periodically, delayed helicity reporting to prevent direct electrical pick up of reversal signal by detectors

Detector signal integrated for each helicity window and asymmetry formed by quartet



Contribution	Expected width (ppm)	$\delta A = \frac{\sigma_A}{\overline{\sigma_A}}$		
Pure statistics	201	$\sqrt{\mathrm{N}_{\mathrm{QRT}}}$		
Detector resolution	92			
Current monitor resolution	50	σ _A = 230-260 ppm		
Target boiling	57	$A_{Phys} = -0.200 \text{ ppm}$ $\delta A_{Phys} = 0.006 \text{ ppm}$		
Total	233.7	on phys – those pph		

Α.

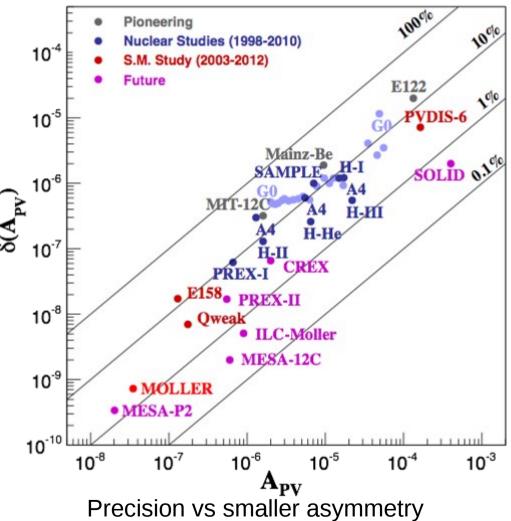
Main Detector All bars Asymmetry (Blinded)

11/10/15

A.

PVES Progress

- Looking to Future : Technical challenges :
- Statistics
 - High rate, beam polarization, beam current, high-power target, large acceptance detectors
- Noise
 - Electronics, target density fluctuations, detector resolution
- Systematics
 - Helicity-correlated beam asymmetry (false asym.), backgrounds, precision beam polarimetry, precise Q² determination



PVES Progress

Looking to Future : Technical challenges :

- Random beam fluctuations limits : present (Qweak) vs. Future (MOLLER)
- Beamline monitor precision : present (Qweak) vs. Future (MOLLER)

Beam property	MOLLER spec.	Qweak observed
Intensity	< 1000 ppm	500 ppm
Energy	< 108 ppm	6.5 ppm
Position	< 47 μm	48 µm
Angle	< 4.7 µrad	1.4 μrad
		Courtooy of Mark Ditt

Courtesy of Mark Pitt

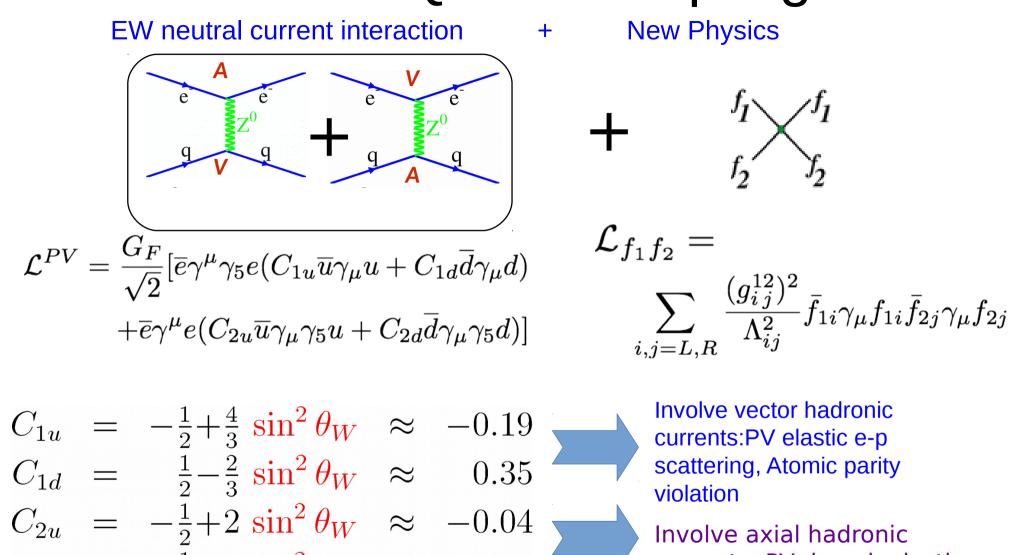
Monitor type	MOLLER spec.	Qweak observed
Beam charge	10 ppm	65 ppm
Beam position	3 μm	6 μm

Courtesy of Mark Pitt

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Electron-Quark Couplings

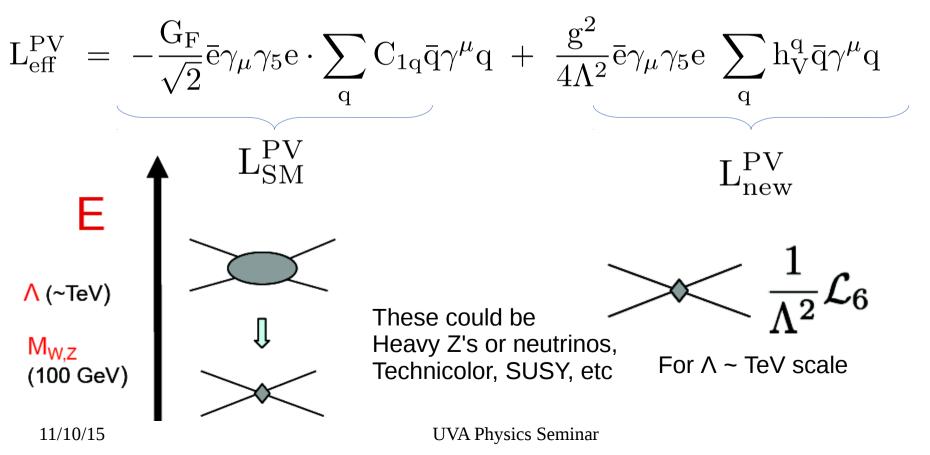


Involve axial hadronic currents: PV deep inelastic scattering

 $C_{2d} = \frac{\tilde{1}}{2} - 2 \sin^2 \theta_W \approx 0.04$

PVES in Search for New Physics

- New physics at high energies can be detected through precision PVES at low energies
 - At low energies new physics appear as a new contact interaction



PVES vs Colliders: Neutral Currents

- Both colliders and PVES can access $\Lambda > 10~\text{TeV}$ but...
- In PVES : both New physics and EW physics amplitudes interference with electromagnetic amplitude

$$|A_{\gamma} + A_Z + A_{New}|^2 \to A_{\gamma}^2 \left[1 + 2\left(\frac{A_Z}{A_{\gamma}}\right) + 2\left(\frac{A_{New}}{A_{\gamma}}\right) \right]$$

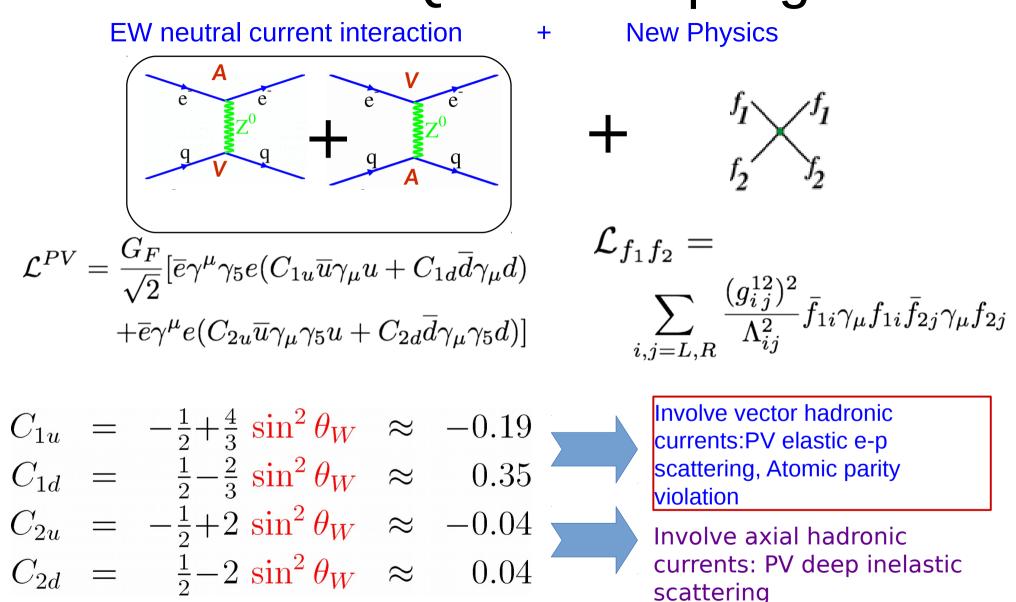
Can observe PV new physics interactions!

• In colliders : No interference

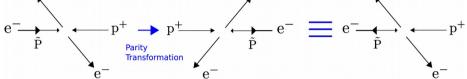
$$|A_Z + A_{New}|^2 \rightarrow A_Z^2 \left| 1 + \left(\frac{A_{New}}{A_Z}\right)^2 \right|$$

At Z resonance A_{γ} is imaginary and no interference observed!

Electron-Quark Couplings



Parity Violating Asymmetry for the Qweak Experiment $e^{-} \xrightarrow{p^{+}} e^{-} \xrightarrow{p^{+}} e^{-} \xrightarrow{p^{+}} e^{-} \xrightarrow{p^{+}} e^{-}$



The Qweak experiment determines the proton's weak charge by measuring the PV asymmetry in elastic scattering of longitudinally polarized electrons on unpolarized protons

$$A_{\rm PV} = \left[\frac{-G_{\rm F}Q^2}{4\sqrt{2}\pi\alpha}\right] \left[\frac{\varepsilon G_{\rm E}^{\gamma}G_{\rm E}^{\rm Z} + \tau G_{\rm M}^{\gamma}G_{\rm M}^{\rm Z} - (1 - 4\sin^2\theta_{\rm W})\varepsilon'G_{\rm M}^{\gamma}G_{\rm A}^{\rm Z}}{\varepsilon(G_{\rm E}^{\gamma})^2 + \tau(G_{\rm M}^{\gamma})^2}\right]$$

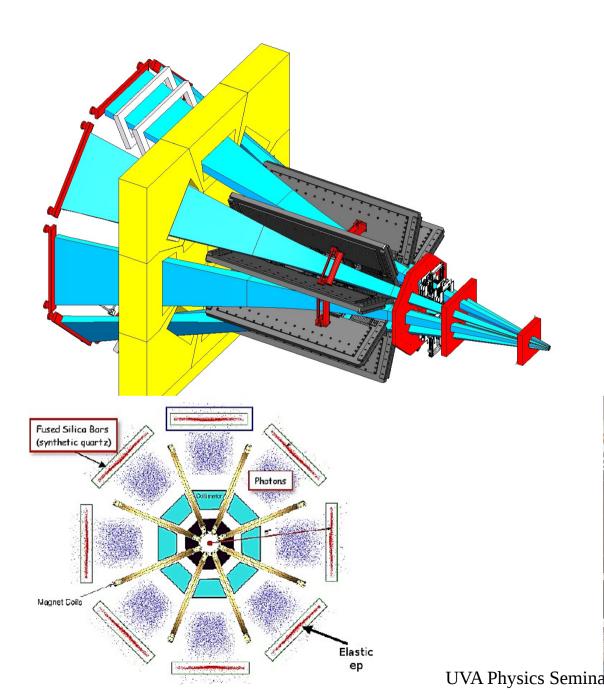
At forward angles and very small Q^2 ,

$$A_{PV}(\vec{e}p) = -\frac{G_F Q^2}{4\sqrt{2}\pi\alpha} [Q_W^P + F(Q^2, \theta)] = A_{Q_W^P} + A_{Had}$$

Proton's weak charge,
$$Q_W^P = 2g_V^u . g_A^e + g_V^d . g_A^e = 1 - 4 \cdot \sin^2\theta_W$$

Form factor term due to finite proton size \rightarrow Hadron structure (~ 30% of the asymmetry) By running the experiment at very small Q², sensitivity to the effects of the "Hadron structure" is minimized

Qweak Experimental Apparatus



Parameters

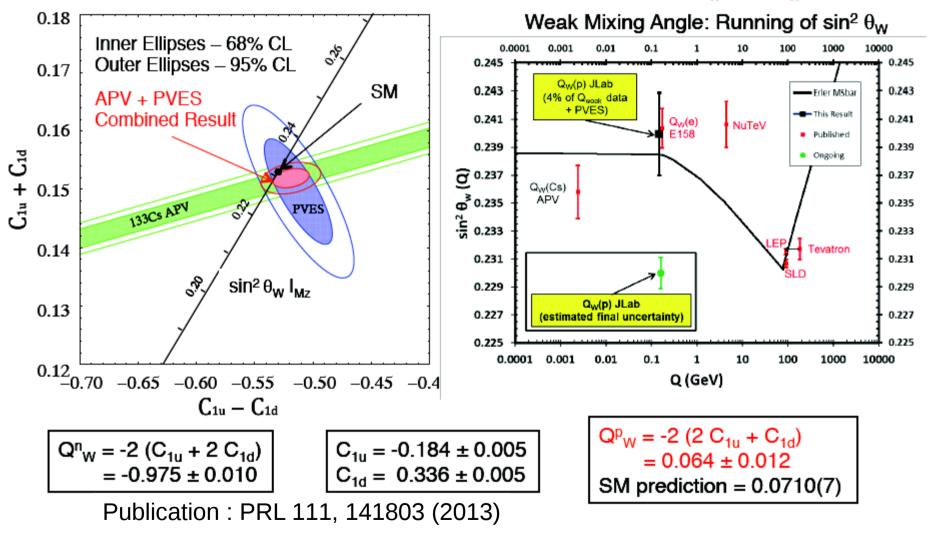
- E_{beam} = 1.165 GeV
- <Q²> = 0.025 GeV²
- $<\theta>=7.9\pm3$
- ϕ coverage = 50% of 2π
- $I_{beam} = 180 \ \mu A$
- Integrated rate = 6.4 GHz
- Beam polarization = 88%
- Target = 35 cm
- Cryo-power = 3 kW



Qweak Commissioning Run

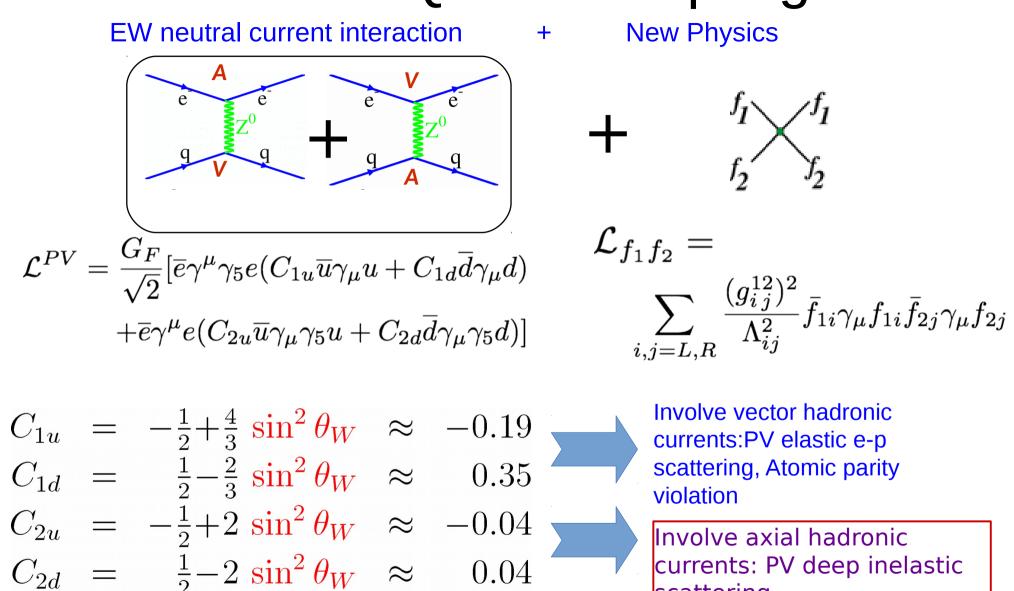
Combined Analysis Extract: C_{1u}, C_{1d}, Qⁿ_W

Qweak + Higher Q² PVES Extract: Q^p_w, sin² θ_w



More production data is still being analyzed : expect final results in 2016!

Electron-Quark Couplings



scattering

PV Deep Inelastic Scattering

Off the simplest isoscalar nucleus (deuterium) at high Bjorken x

$$A_{PV}^{DIS} = \frac{G_F Q^2}{4\sqrt{2}\pi\alpha} [2g_A^e Y_1(y) \frac{F_1^{\gamma Z}}{F_1^Z} + 2g_V^e Y_3(y) \frac{F_3^{\gamma Z}}{F_1^Z}]$$

At high x, deuterium PV asymmetry becomes independent of PDFs, x and W, with well defined SM predictions for given Q^2 and y = 1 - E'/E

$$\begin{split} & \text{For } Q^2 >> 1 \text{ GeV}^2 \text{ and } W^2 > 4 \text{ GeV}^2 \\ A_{PV}^D &= \frac{G_F Q^2}{4\sqrt{2}\pi\alpha} [a_1(x) Y_1(y) + a_3(x) Y_3(y)] \\ & \text{Where, } Y_1 \simeq 1; Y_3 \simeq \frac{1 - (1 - y)^2}{1 + (1 - y)^2} \end{split}$$

$$a_1^{D}(x) = \frac{6}{5}(2C_{1u} - C_{1d})(1 + \frac{0.6s^+}{u^+ + d^+})$$
$$a_3^{D}(x) = \frac{6}{5}(2C_{2u} - C_{2d})(\frac{u^- + d^-}{u^+ + d^+})$$

Where
$$f_i^{\pm} = f_i \pm \bar{f}_i, y = \frac{E-E}{E}$$

Interplay with QCD,

- Flavor dependent quark distributions (u,d, and s)
- Charge symmetry violations (CSV)
- Higher twist effects (HT)
- Nuclear medium effects (EMC)

e⁻

 $\frac{\gamma}{z^0}$

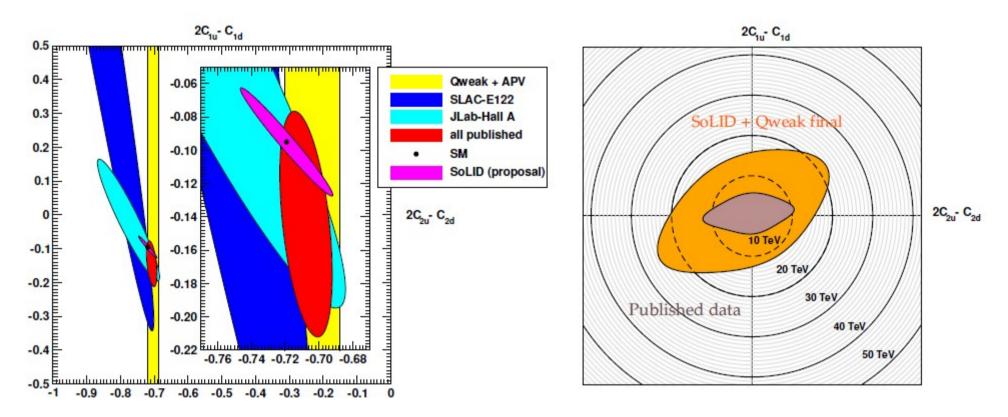
SoLID-PVDIS Physics Motivation

- A precision test of the Standard Model
- Search for Charge Symmetry Violation (CSV)
- Test of QCD higher twist corrections (quark quark correlations)
- Measurement of d/u quark ratio for proton

Attractive PVDIS feature

- Large PV asymmetries with manageable backgrounds
- Ability to reach higher precision beam polarimetry with 11 GeV electron beam energies

Projected Coupling Constraints from PVDIS



Constraint on quark coupling constants and updated limits on new physics beyond SM using SoLID-PVDIS projections

Charge Symmetry Violations

Sensitivity to CSV $R_{CSV} = \frac{\delta A_{PV}}{A_{PV}} = 0.28 \frac{\delta u(x) - \delta d(x)}{\delta u(x) + \delta d(x)}$ Where $\delta u \equiv u^p - d^n; \delta d \equiv d^p - u^n;$ BAG Model + OED Splitting OLOGED Splitting in MRST Uncertainty band, this proposed 0.00 0.0

Direct observation of parton level CSV

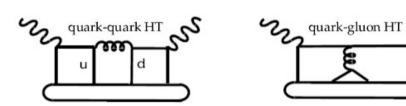
- Charge symmetry \rightarrow u^p = dⁿ; d^p = uⁿ
- Fractional change in APV due to CSV from different models shown
- The uncertainty band using PVDIS figure-of-merit is plotted

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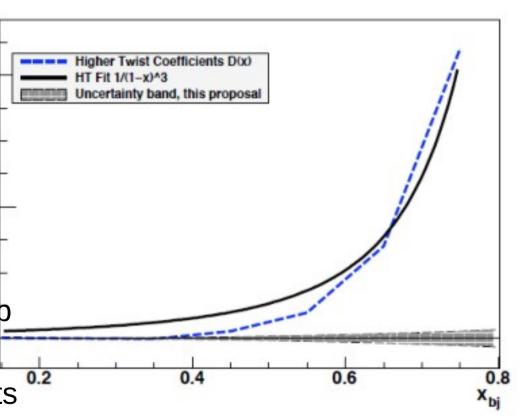
Higher Twists effects in PVDIS

~5

- In QCD, additional Q² dependence gives information on quark-quark and quark-gluon correlations
 - Higher Twist (HT) terms

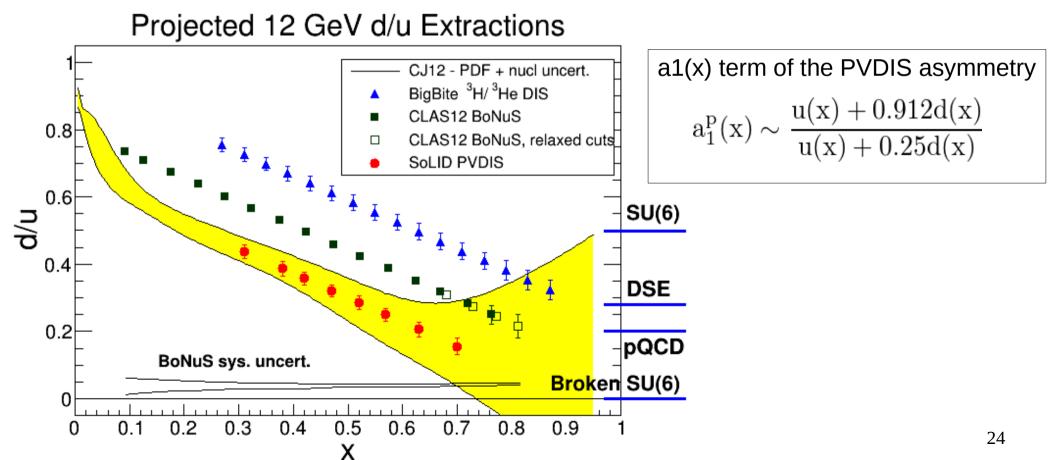


- With PVDIS asymmetry measurements, only Q² dependence of q-q HT can show up
- Large kinematic reach in SoLID allows for evaluation of higher twists
- PVDIS signature is the variation of Y_1a_1 term (of the APV) with x and Q^2

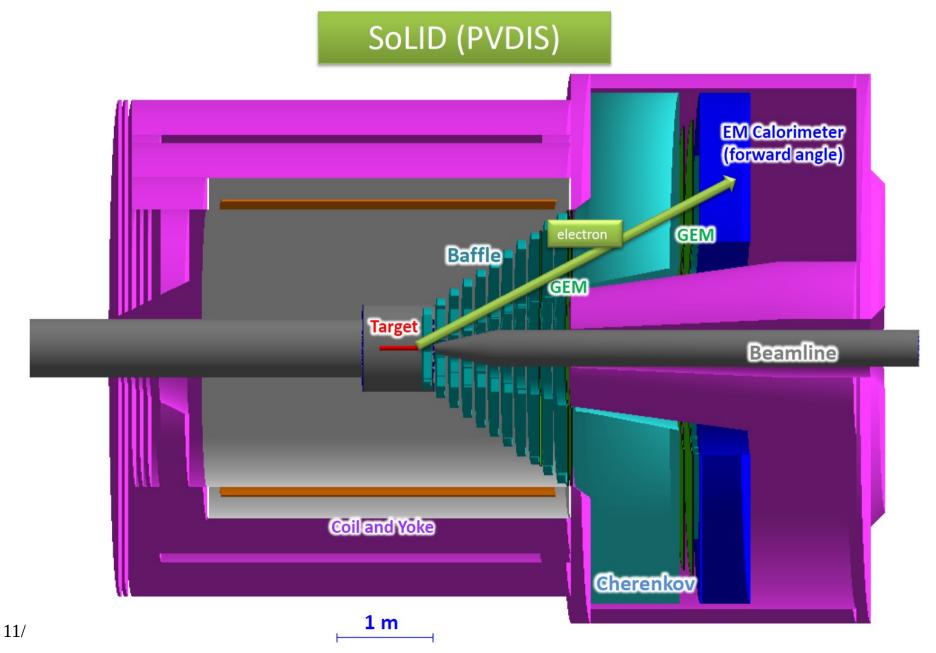


Quark Flavor Dependent Effects on Proton

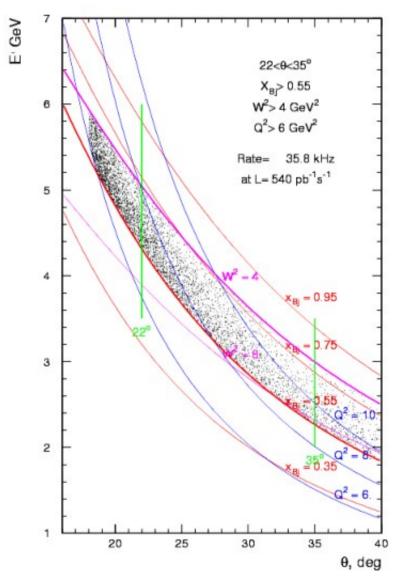
- Measurement of d(x)/u(x) ratio for the proton at high x
- A clean measurement free from any nuclear corrections
- Uncertainties of set of PVDIS measurements are shown in the plot (red points)
 - Provides a high precision measurements in range of x



Solenoidal Large Intensity Device (SoLID) Apparatus

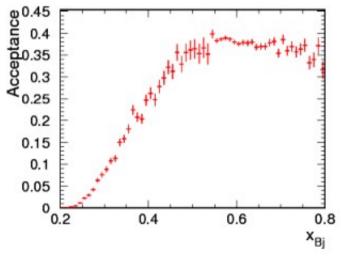


Solenoidal Large Intensity Device (SoLID) Acceptance



SoLID Specs. and Figure-Of-Merit

- High Luminosity (10³⁹ cm²/s)
- Beam current 50 uA and polarization 85%
- Large scattering angles for high x and y access



- With moderate running times,
 - X-range of 0.25 to 0.75
 - W² > 4 GeV²
 - Q^2 range a factor of 2 for each x

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SOLID-PVDIS Figure-Of-Merit

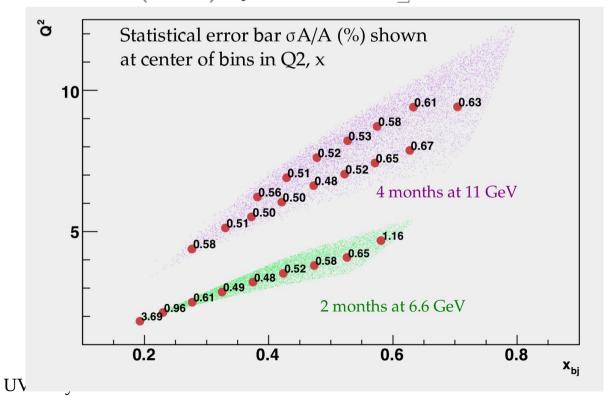
Sub. 1% precision over broad range of kinematic range: <u>A Standard Model test and a detailed study</u> of hadronic structure contributions

 $A_{PV}^{D} = A_{PV}^{EW} (1 + \beta_{HT} \frac{1}{(1-x)^3 O^2} + \beta_{CSV} x^2)$

If no CSV, HT, quark sea, or nuclear effects, All (Q2, x) bins should give the asymmetry within statistics and kinematic factors

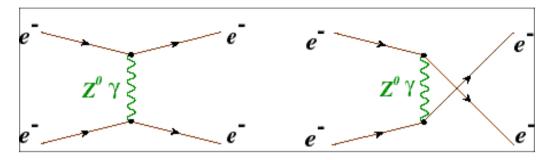
Fit to data :

Kinematics dependence of Physics					
х	У	Q^2			
	YES				
YES					
YES		YES			
	X	x y YES			



PV in MØller Scattering

A Search for New Physics at the TeV Scale



- Proposed MOLLER experiment will be the best contact interaction search for leptons at low OR high energy
 - Best current limit on contact interaction scales available from LEP2
 - LEP2 only sensitive to parity conserving quantities $(g_{RL}^2 and g_{RR}^2+g_{LL}^2)$

Where $g_{ij}=g_{ij}^{*}$ are contact interaction coupling constants for chirality projections of the electron spinor

- Model independent mass scale for parity violating interactions :

$$\mathcal{L}_{e_1 e_2} = \sum_{i \ i=L \ B} \frac{g_{ij}^2}{2\Lambda^2} \bar{e_i} \gamma_\mu e_i \bar{e_j} \gamma^\mu e_j \qquad \qquad \frac{\Lambda}{(g_{RR}^2 - g_{LL}^2)} = 7.5 \ \text{TeV}$$

The MOLLER measurement will extend the current sensitivity of 4electron contact interactions, both qualitatively and quantitatively

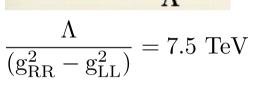
PV in MØller Scattering

A Search for New Physics at the TeV Scale

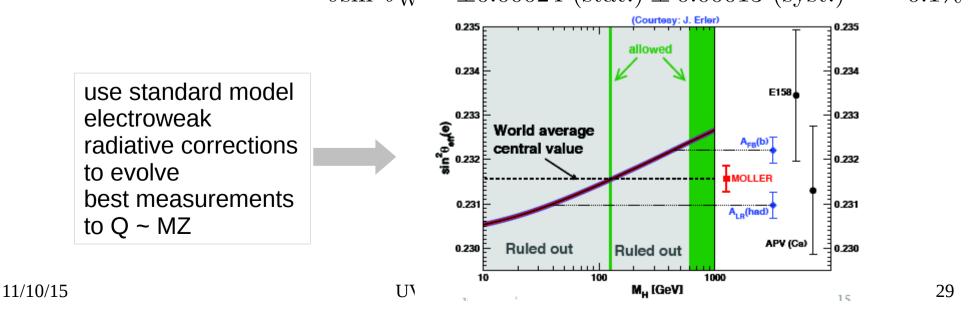
• Measure weak charge of electron precisely $Q_{w}^{e} \sim 0.045$

$$\begin{array}{c} e^{-} \\ e^{-} \\ e^{-} \\ e^{-} \end{array} \quad \frac{\delta Q_{W}^{e}}{Q_{W}^{e}} = 2.4\% \rightarrow A_{new} \sim 0.001 \cdot G_{F} + \\ e^{-} \end{array}$$

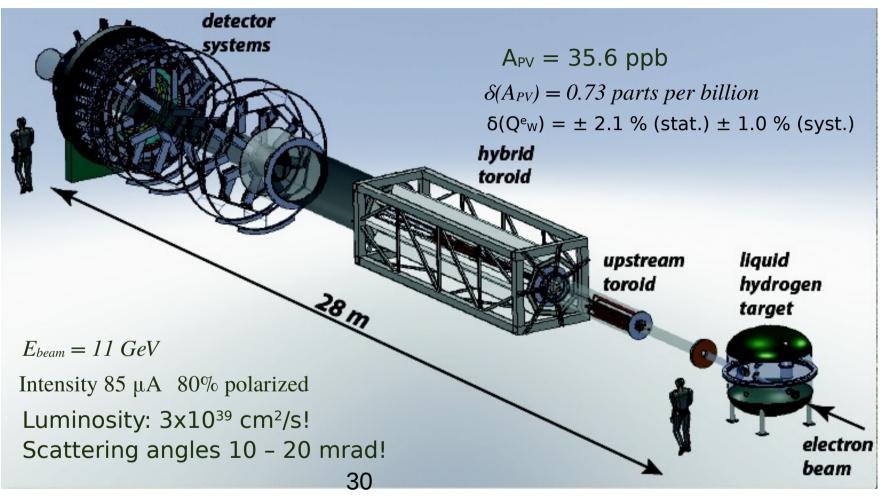
Unprecedented sensitivity



• Provide best projected uncertainty weak mixing angle at any energy scale $\delta \sin^2 \theta_W = \pm 0.00024 \text{ (stat.)} \pm 0.00013 \text{ (syst.)} \rightarrow \sim 0.1\%$



MOLLER Apparatus



MOLLER Context Summary

- Best contact interaction reach for leptons at any energy
 - Similar to LHC reach with semi-leptonic amplitudes
 - To do better for a 4-lepton contact interaction would require:
 - Giga-Z factory, linear collider, neutrino factory or muon collider
- If LHC sees any anomaly in runs 2 and 3 (~ 2022)
 - The unique discovery capability in MOLLER will be very important
- MOLLER also provides discovery scenarios beyond LHC signatures
 - Hidden weak scales
 - Lepton number violating interactions
 - Light dark matter mediators

PREX/CREX : Neutral Current as a Probe of the Neutron

- Y, Z⁰ Protons & Neutron Skin State Protons & Neutrons
- Weak neutral current : A clean probe couples mainly to neutrons

$$A_{\rm PV} = \frac{\frac{d\sigma^{\rm R}}{d\Omega} - \frac{d\sigma^{\rm L}}{d\Omega}}{\frac{d\sigma^{\rm R}}{d\Omega} + \frac{d\sigma^{\rm L}}{d\Omega}} = \frac{G_{\rm F}Q^2}{2\pi\alpha\sqrt{2}} \left[1 - 4 \cdot \sin^2\theta_{\rm W} + \frac{F_{\rm n}(Q^2)}{F_{\rm p}(Q^2)} \right]$$

$$_{\rm eak} = 1 - 4\sin^2\theta_{\rm W} \sim 0.076 \qquad \qquad Q_{\rm weak}^{\rm n} \sim -1 \qquad A_{\rm PV} \to 10^{-6}$$

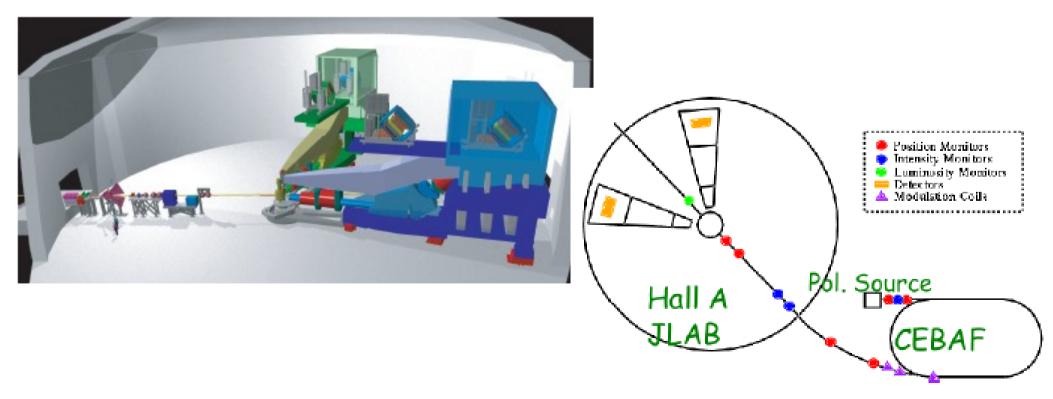
 It provides theoretically clean method to measure neutron radius and skin thickness

$$A_{PV} \rightarrow F_{W}(Q^{2}) \rightarrow R_{n} \rightarrow (R_{n} - R_{p})$$

 Q_w^p

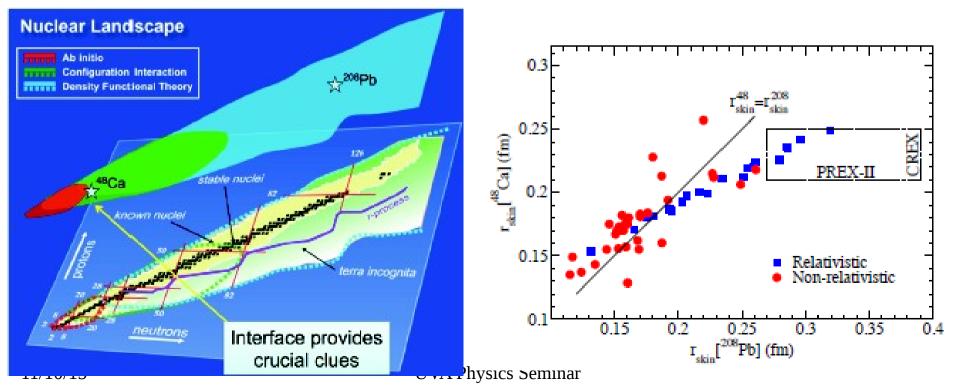
Experimental Setup

- Two High Momentum Spectrometers (HRS) to run simultaneously
 - Will require a Septum magnet to reach our acceptance
- PREX acceptance at about 5° Using E = 1.1 GeV beam
- CREX acceptance at about 4° Using E = 2.2 GeV beam
- Both ²⁰⁸Pb and ⁴⁸Ca provide large inelastic separation with HRS and have very long life time for a neutron excess nuclei

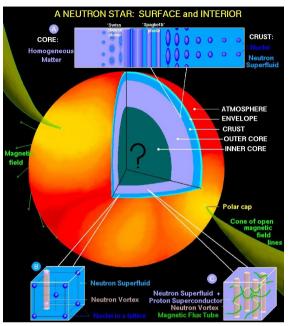


Why Two different Nuclei?

- Ab initio calculations only reach as far as medium nuclei such as ⁴⁸Ca
 - Experimental data from ²⁰⁸Pb and ⁴⁸Ca will provide a bridge between medium nuclei ab initio calculations and heavy nuclei Density Functional Theory (DFT) calculations.
- Correlations predicted between neutron skin of ²⁰⁸Pb and ⁴⁸Ca need experimental validations



PREX Implications : Neutron Stars





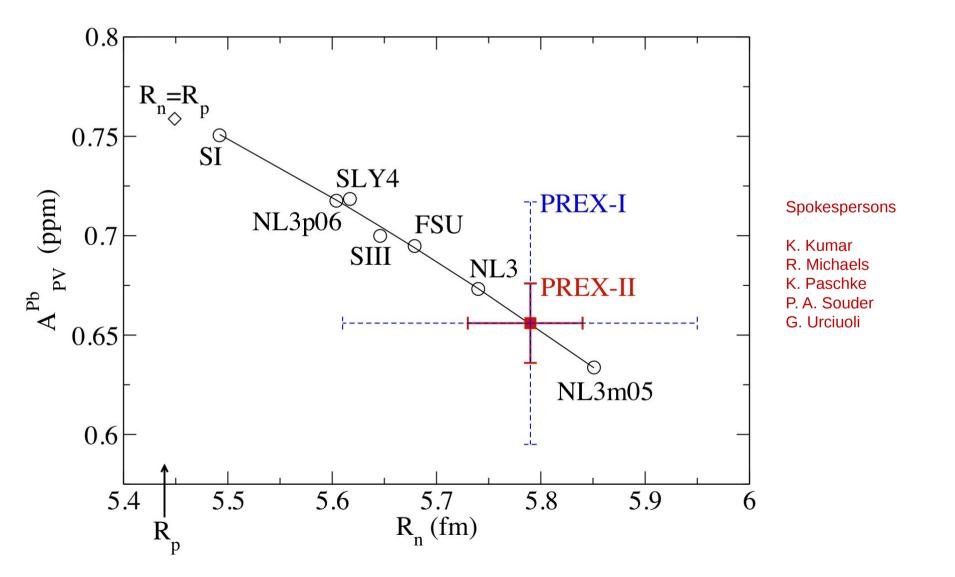
Courtesy of C.J. Horowitz and J. Piekarewicz

- R_N calibrates equation of state (pressure vs density) of Neutron rich matter
- Combine PREX $R_{\mbox{\tiny N}}$ with observed neutron star radii
 - Phase transition to "Exotic" Core?
 - Strange star? Quark star?
- Some neutron stars seem too cold
 - Explained by cooling by neutrino emission (URCA process)?
 - Only if ($R_N R_p$) \rightarrow 0.2 fm : URCA is probable

Crab Pulsar

PREX : Earlier Results

Neutron Skin = $R_N - R_P = 0.33 + 0.16 - 0.18$ fm



PREX/CREX : Next Run

- PREX-II is on its way to make many improvements over several PREX-I radiation damage issues
 - Damaging neutron (0.1 < E < 10 MeV) dose is reduced by 78% compared to PREX-I
 - High energy (E>10 MeV) photon dose is reduced by 80%
 - Collimator design is almost ready
 - Neutron radiation shielding optimization is underway
 - Final design will further improve dose reduction
- Neutron density measurements for ²⁰⁸Pb and ⁴⁸Ca will provide necessary support for better nuclear structure theory models
 - For nuclei to neutron stars with implications on nuclear structure studies to astrophysics

PVES as a Probe of EMC effect

- PVDIS offers a picture into partonic distributions by probing new flavor combinations
- Expanding the a_1 term about the isoscalar limit

•
$$a_1 \simeq \frac{9}{5} - 4\sin^2\theta_W - \frac{12}{25}\frac{u_A^+ - d_A^+}{u_A^+ + d_A^+}$$

Where $u_A = u$ in p and d in n

- PVDIS asymmetry is sensitive to differences in the quark flavors
 - For isoscalar targets the asymmetry becomes a test for charge symmetry violation

For
$$Q^2 >> 1 \text{ GeV}^2$$
 and $W^2 > 4 \text{ GeV}^2$
 $A_{PV}^D = \frac{G_F Q^2}{4\sqrt{2}\pi\alpha} [a_1(x)Y_1(y) + a_3(x)Y_3(y)]$
Where, $Y_1 \simeq 1; Y_3 \simeq \frac{1 - (1 - y)^2}{1 + (1 - y)^2}$
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Isovector dependence of EMC effect

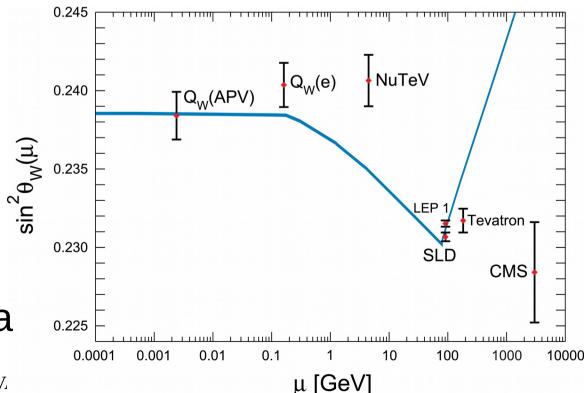
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NuTeV results from Fermilab

- Neutrino scattering is sensitive to different flavor combinations
- Asymmetry nuclei (iron target used in NuTeV) need corrections
- CSV or isovector EMC effects could play significant role and not well constrained by data

Pachos-Wolfenstein relation:

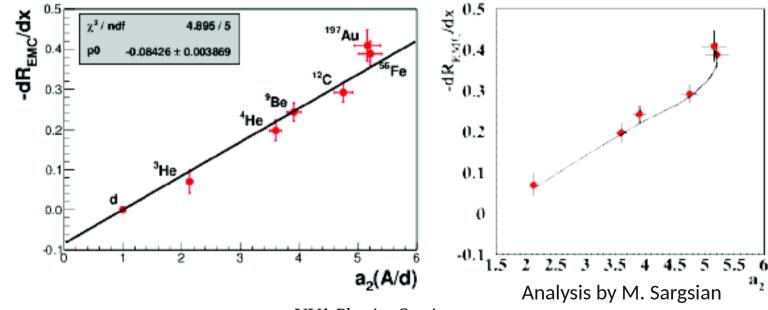
$$R_{\rm PW} \equiv \frac{\sigma(\nu_{\mu}N \to \nu_{\mu}X) - \sigma(\bar{\nu}_{\mu}N \to \bar{\nu}_{\mu}X)}{\sigma(\nu_{\mu}N \to \mu^{-}X) - \sigma(\bar{\nu}_{\mu}N \to \mu^{+}X)}$$
$$= \lim_{\to i.s.} \frac{1}{2} - \sin^{2}\theta_{W}$$



Isovector dependence of EMC effect

Short range correlations and EMC effect

- SRC show strong preference to n-p pairs over p-p/n-n
 - SRCs generated by interactions in short-distance (high density)
- EMC effect correlates with SRC
 - EMC effect driven by high-density nucleon configurations (pairs, clusters)
- Preliminary models make predictions for asymmetry nuclei
 - (Z N) boost by isovector enhancement?



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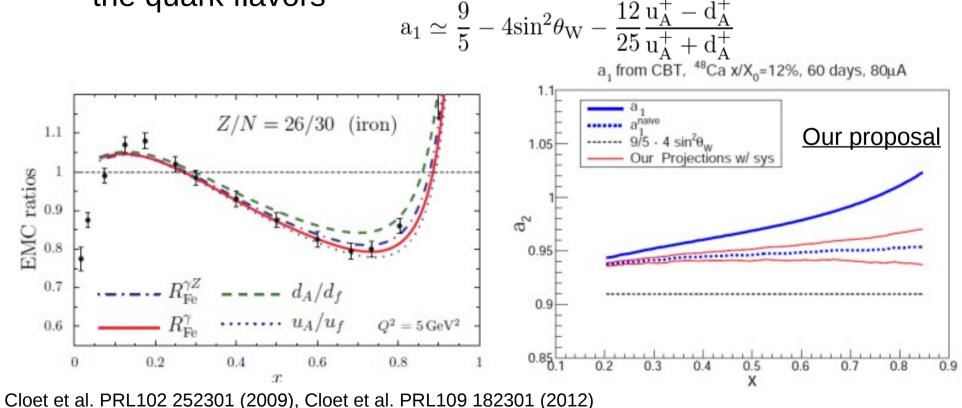
PVDIS Constraints on EMC Effect

- PVDIS on asymmetric target (48Ca or 9Be) will test isovector dependence,
 - Larger A \rightarrow larger EMC and larger (Z N) gives an boost to isovector enhancement
 - PV asymmetry is independent of overall size of EMC effect; only sensitive to difference in EMC effect for u and d quarks
- 48Ca DIS Rates and backgrounds are comparable for deuterium DIS
- Therefore isovector observables on an asymmetric target is doable with SoLID-PVDIS
- 60 days production will offer powerful constraints, help resolve the NuTeV anomaly, and test leading models to several sigma

Flavor Dependent Model EMC Predictions

PVDIS with neutron rich nuclei ⁴⁸Ca can constrain possible flavor-dependent nuclear medium modification effects on quarks

 PVDIS asymmetry is a direct measurement of differences in the quark flavors
 12 u⁺ d⁺



Conclusions

- Jlab is a great facility to do PVES
 - Polarized Continuous electron beam
- PVES provides unique information for nuclear physics
 - Nucleons : EMC effect, strangeness, weak form factors
 - Nuclei : PREX/CREX
- PVES is a precision frontier of testing the SM and physics beyond SM
 - Qweak, SoLID-PVDIS and MOLLER
 - Complementary to LHC

Random Beam Fluctuations and Beamline Instrumentation

Use Qweak experience (@ 1 kHz data rate) \rightarrow

Assess MOLLER specifications (@ 2 kHz data rate) for beam fluctuations/monitoring

Random beam fluctuations ("jitter") @2 kHz:

If 12 GeV machine is as "quiet" as 6 GeV machine, these will be easily satisfied!

Beam property	MOLLER spec.	Qweak observed
Intensity	< 1000 ppm	500 ppm
Energy	< 108 ppm	6.5 ppm
Position	< 47 μm	48 µm
Angle	< 4.7 µrad	1.4 μrad

Beamline monitor precision @2 kHz:

- Position nearly satisfied
- Charge monitoring will require further developments
- Start with BCM digital receiver studies

NEW: actually BPM spec is probably already achieved ^{8/12/2015} 11/10/15

1	Monitor type	MOLLER spec.	Qweak observed
<	Beam charge	10 ppm	65 ppm
	Beam position	3 μm	6 μm
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/er			

MOLLER Aug. 2015 Collab. Meeting UVA Physics Seminar

PV Deep Inelastic Scattering

Off the simplest isoscalar nucleus (deuterium) at high Bjorken x

$$A_{PV}^{DIS} = \frac{G_F Q^2}{4\sqrt{2}\pi\alpha} [2g_A^e Y_1(y) \frac{F_1^{\gamma Z}}{F_1^Z} + 2g_V^e Y_3(y) \frac{F_3^{\gamma Z}}{F_1^Z}] \quad \text{Where, } Y_1 \simeq 1; Y_3 \simeq \frac{1 - (1 - y)^2}{1 + (1 - y)^2} = \frac{G_F Q^2}{4\sqrt{2}\pi\alpha} [a_1(x)Y_1(y) + a_3(x)Y_3(y)]$$

At high x, deuterium PV asymmetry becomes independent of PDFs, x and W, with well defined SM predictions for given Q^2 and y = 1 - E'/E

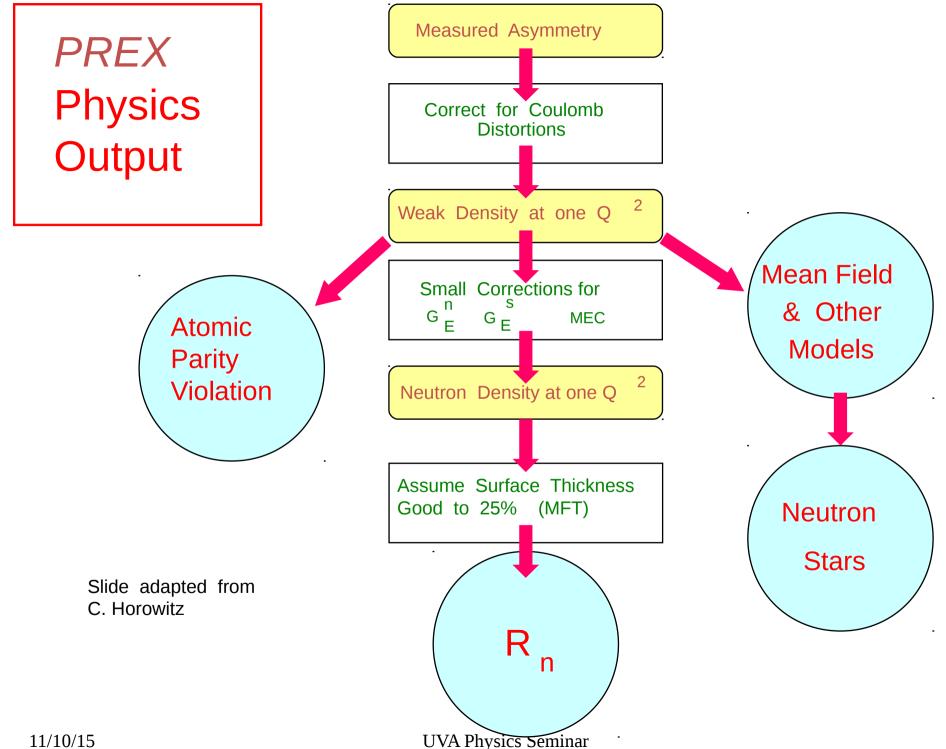
$$\begin{split} A_{PV}^{D} &= \frac{\sigma^{L} - \sigma^{R}}{\sigma^{L} + \sigma^{R}} \quad \text{For } Q^{2} >> 1 \text{ GeV}^{2} \text{ and } W^{2} > 4 \text{ GeV}^{2} \\ &= -\left(\frac{3G_{F}Q^{2}}{2\pi\alpha\sqrt{2}}\right) \frac{(2C_{1u} - C_{1d})(1 + R_{s}) + Y(2C_{2u} - C_{2d})R_{\nu})}{5 + R_{s}} \quad Y = \frac{1 - (1 - y)^{2}}{1 + (1 - y^{2}) - y^{2}\frac{R}{R+1}} \\ R(x, Q^{2}) = \frac{\sigma^{L}}{\sigma^{R}} \simeq 0.2 \end{split}$$

Interplay with QCD,

- Flavor dependent quark distributions (u,d, and s)
- Charge symmetry violations (CSV)
- Higher twist effects (HT)
- Nuclear medium effects (EMC)

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 e^-



11/10/15

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Anticipated Errors

PREX-II at E = 1.1 GeV ; $A_{PV} = 0.6$ ppm

Systematic Error	Contribution
Charge normalization	0.1%
Beam asymmetries	1.1%
Detector non-linearity	1.0%
Transverse	0.2%
Polarization	1.1%
Inelastic contribution	< 0.1%
Effective Q2	0.4%
Total	2%

CREX at E = 2.2 GeV; A_{PV} = 2 ppm

Systematic Error	Contribution
Charge normalization	0.1%
Beam asymmetries	0.3%
Detector non-linearity	0.3%
Transverse	0.1%
Polarization	0.8%
Inelastic contribution	0.2%
Effective Q2	0.8%
Total	1.2%

SoLID-PVDIS Error Budget

Error budget for PVDIS asymmetry at x=0.4

Source	Error $(\%)$
Statistics	0.3
Polarimetry	0.4
Q^2	0.2
Radiative corrections	0.3
Total	0.6