

# **MEG Lepton Flavor Violation Search: Challenges and Solutions**

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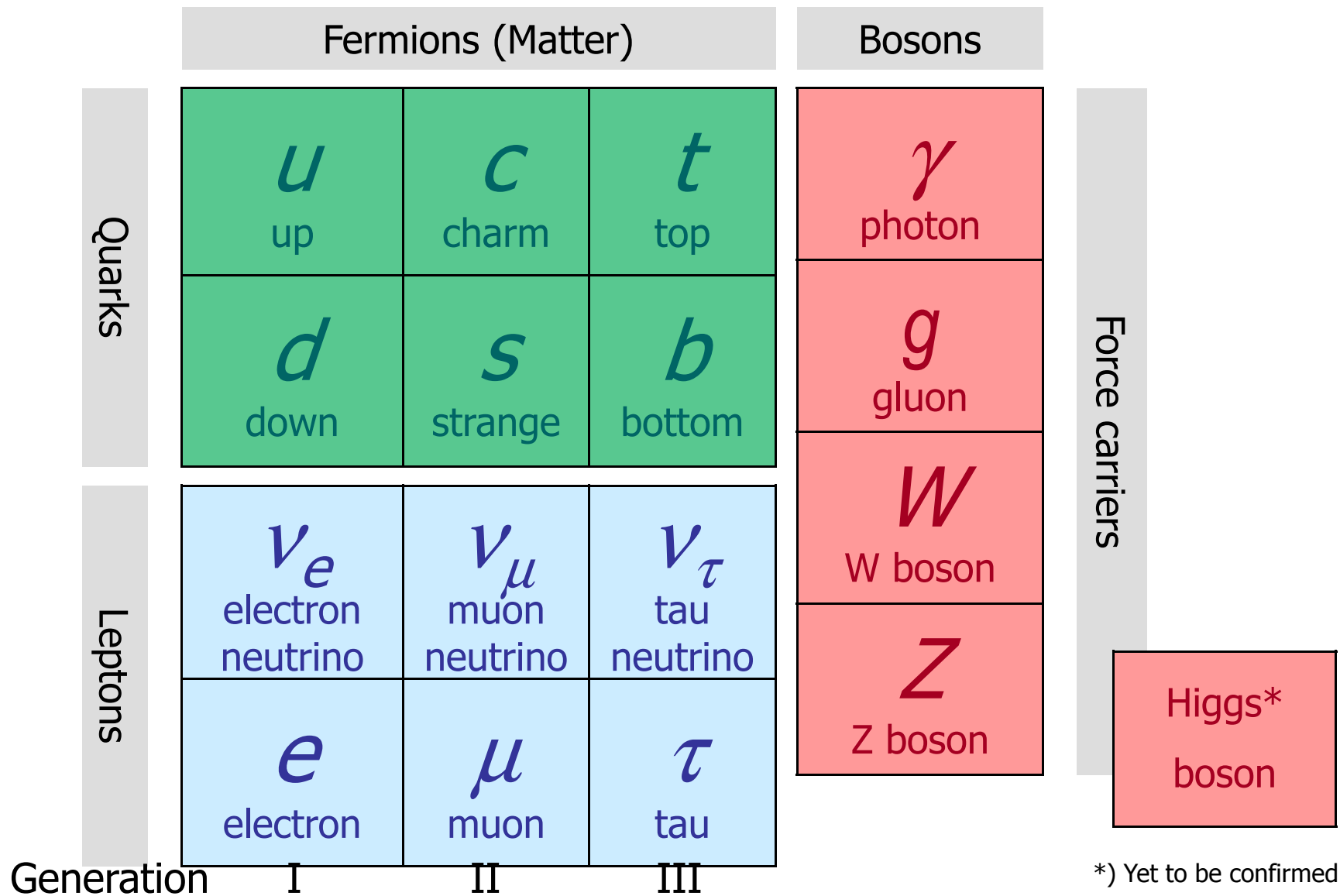
- Motivation to search for  $\mu \rightarrow e \gamma$
- Challenges:
  - Beam, Detectors, Electronics
- Status and Outlook



# Motivation

Why should we search for  $\mu \rightarrow e \gamma$  ?

# The Standard Model

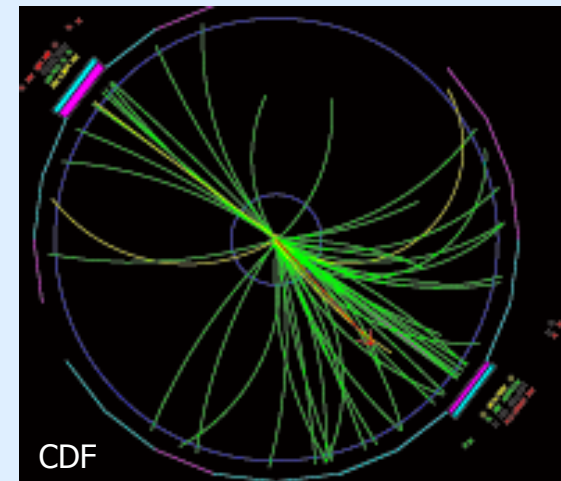


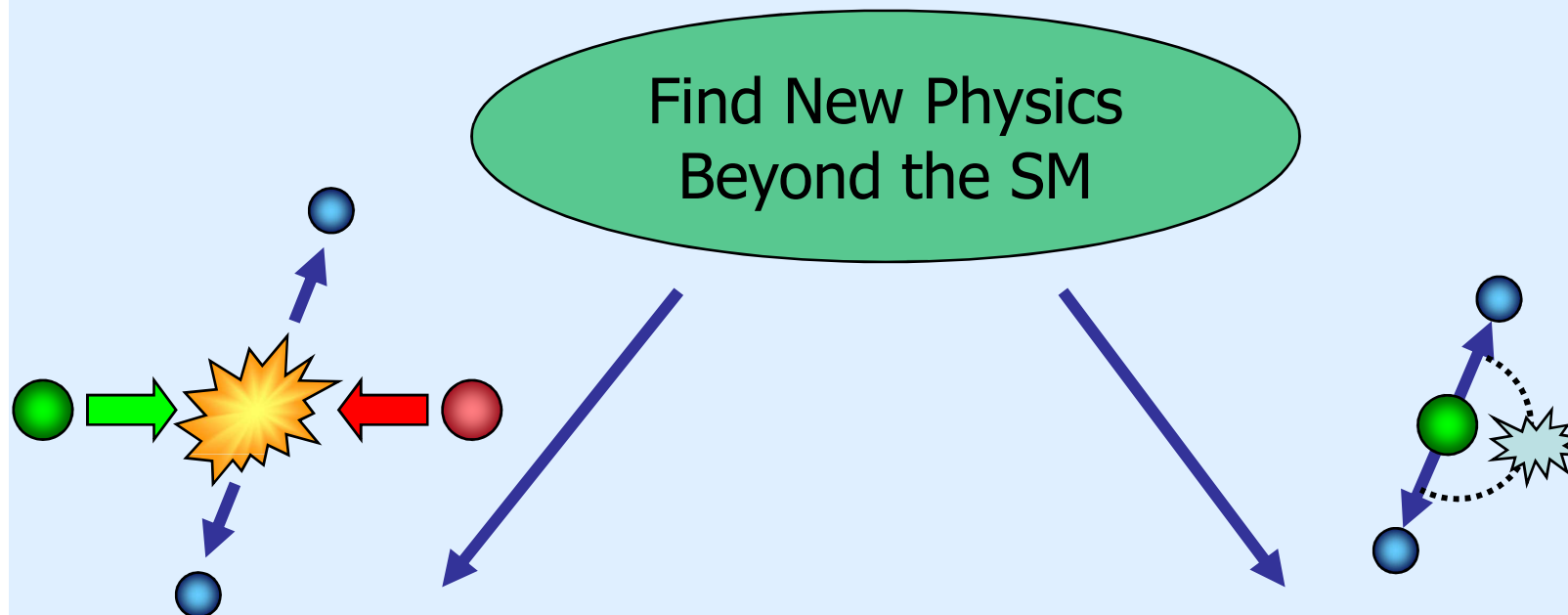
# The success of the SM



- The SM has been proven to be extremely successful since 1970's
  - Simplicity (6 quarks explain >40 mesons and baryons)
  - Explains all interactions in current accelerator particle physics
  - Predicted many particles (most prominent  $W$ ,  $Z$ )
- Limitations of the SM
  - Currently contains 19 (+10) free parameters such as particle (neutrino) masses
  - Does not explain cosmological observation such as Dark Matter and Matter/Antimatter Asymmetry

Today's goal is to look for  
physics beyond the standard  
model





## High Energy Frontier

- Produce heavy new particles directly
- Heavy particles need large collider
- LHC / ILC

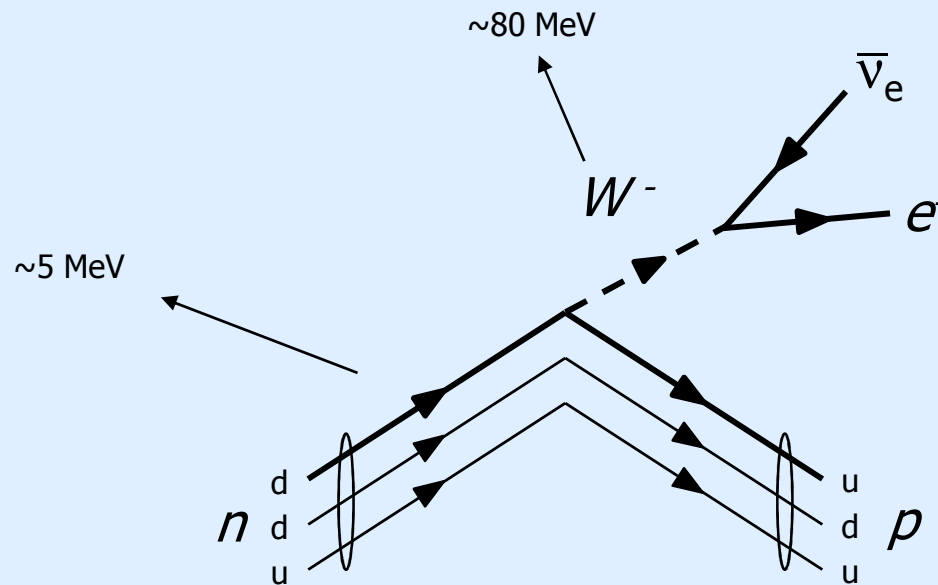
## High Precision Frontier

- Look for small deviations from SM ( $g-2$ )
- Look for forbidden decays
- Requires high precision at low energy

# Neutron beta decay



Neutron  $\beta$  decay via intermediate heavy  $W^-$  boson



Neutron mean  
life time:

886 s

$\beta$  decay discovery:

~1934

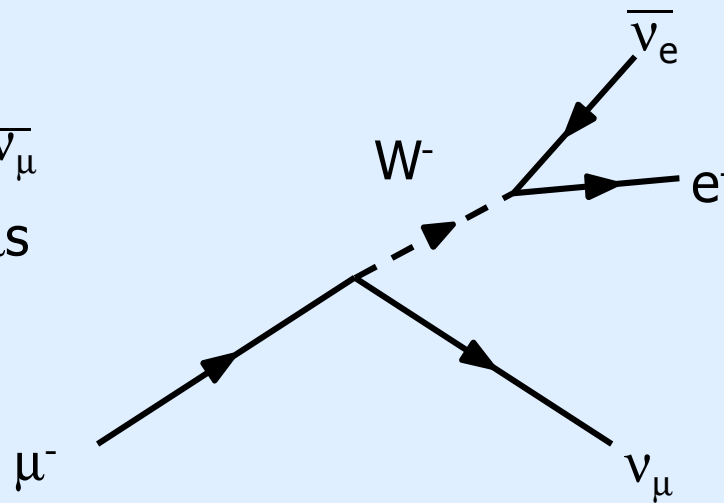
$W^-$  discovery:

1983

# The Muon



- Discovery: 1936 in cosmic radiation
- Mass:  $105 \text{ MeV}/c^2$
- Decay:  $\mu^+ \rightarrow e^+ \nu_e \bar{\nu}_\mu$
- Mean lifetime:  $2.2 \mu\text{s}$



Seth Neddermeyer



MAY 15, 1937

PHYSICAL REVIEW

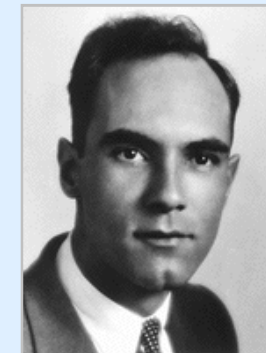
VOLUME 51

## Note on the Nature of Cosmic-Ray Particles

SETH H. NEDDERMEYER AND CARL D. ANDERSON  
*California Institute of Technology, Pasadena, California*  
(Received March 30, 1937)

**M**EASUREMENTS<sup>1</sup> of the energy loss of particles occurring in the cosmic-ray showers have shown that this loss is proportional to the energy of the particles. In contrast to the massive than protons but more penetrating than electrons obeying the Bethe-Heitler theory, we have taken about 6000 counter-tripped photo-

Carl Anderson



# Lepton number conservation



Fermions (Matter)			
Quarks	$u$ up	$c$ charm	$t$ top
	$d$ down	$s$ strange	$b$ bottom
Leptons	$\nu_e$ electron neutrino	$\nu_\mu$ muon neutrino	$\nu_\tau$ tau neutrino
	$e$ electron	$\mu$ muon	$\tau$ tau
Generation	I	II	III
	$\uparrow$ $L_e$	$\uparrow$ $L_\mu$	$\uparrow$ $L_\tau$

$$\mu^+ \rightarrow e^+ \nu_e \bar{\nu}_\mu \quad \approx 100\%$$

$$\begin{array}{lcl} L_e: & 0 & = -1 + 1 + 0 \\ L_\mu: & -1 & = 0 + 0 - 1 \end{array}$$

$$\mu^+ \rightarrow e^+ \nu_e \bar{\nu}_\mu \gamma \quad 0.014$$

$$\begin{array}{lcl} L_e: & 0 & = -1 + 1 + 0 + 0 \\ L_\mu: & -1 & = 0 + 0 - 1 + 0 \end{array}$$

$$\mu^+ \rightarrow e^+ \gamma \quad < 10^{-11}$$

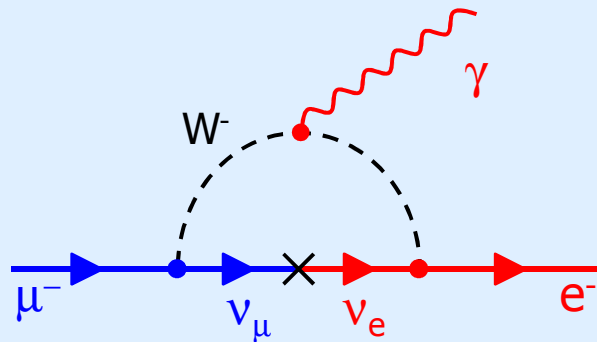
$$\begin{array}{lcl} L_e: & 0 & \neq -1 + 0 \\ L_\mu: & -1 & \neq 0 + 0 \end{array}$$



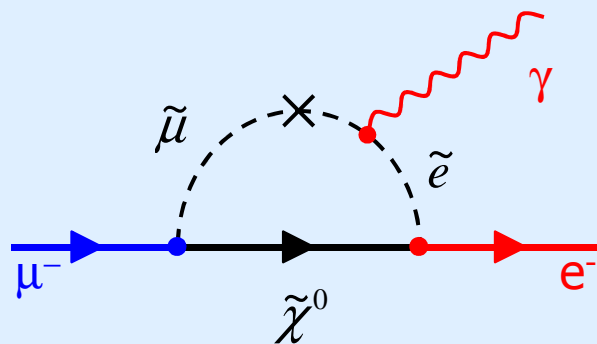
**Violates Lepton Number Conservation!**



- While LFV is forbidden in SM, it is possible in SUSY



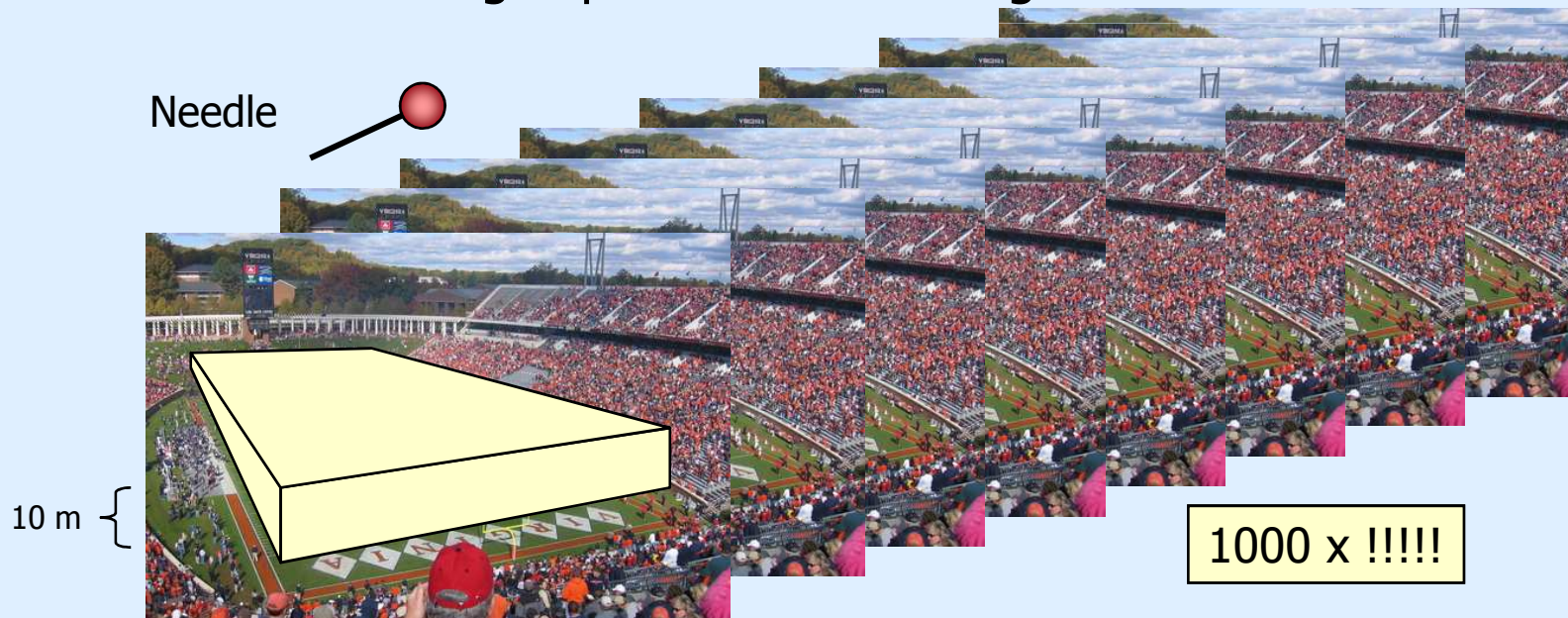
$$\text{BR}(\mu^- \rightarrow e^- \gamma)|_{\text{SM}} \propto \frac{m_\nu^4}{m_W^4} \approx 10^{-60}$$



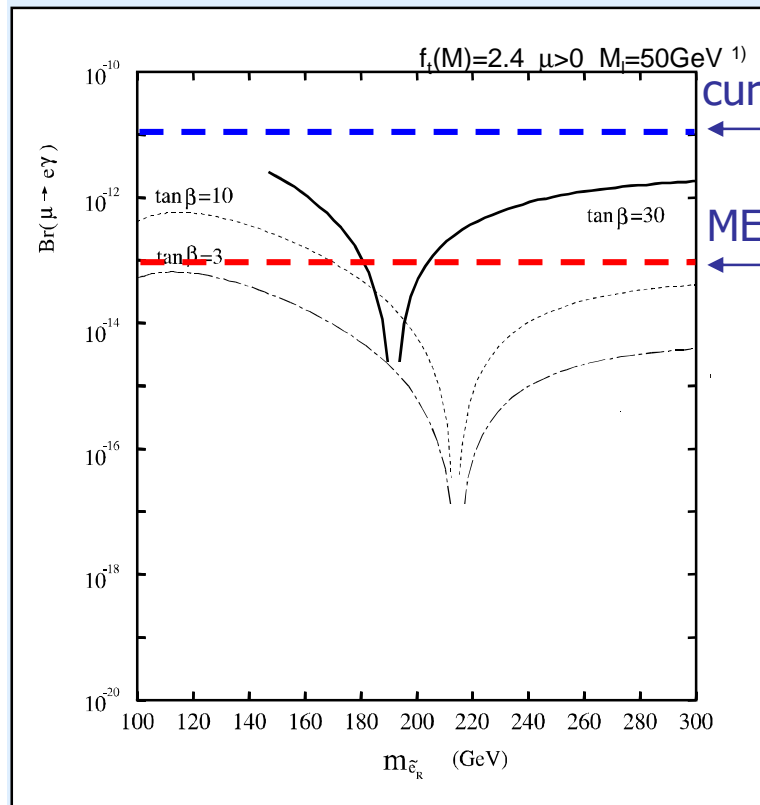
$$\text{BR}(\mu^- \rightarrow e^- \gamma)|_{\text{SUSY}} \approx 10^{-5} \frac{\Delta m_{\tilde{e}\tilde{\mu}}^2}{\bar{m}_\ell^2} \left( \frac{100 \text{ GeV}}{m_{\text{SUSY}}} \right)^4 \tan^2 \beta \approx 10^{-12}$$

Current experimental limit:  $\text{BR}(\mu \rightarrow e \gamma) < 10^{-11}$

- LFV is forbidden in the SM, but possible in SUSY (and many other extensions to the SM) though loop diagrams ( $\rightarrow$  heavy virtual SUSY particles)
- If  $\mu \rightarrow e \gamma$  is found, new physics beyond the SM is found
- Current exp. limit is  $10^{-11}$ , predictions are around  $10^{-12} \dots 10^{-14}$
- Goal of  $10^{-13}$  is a big experimental challenge!

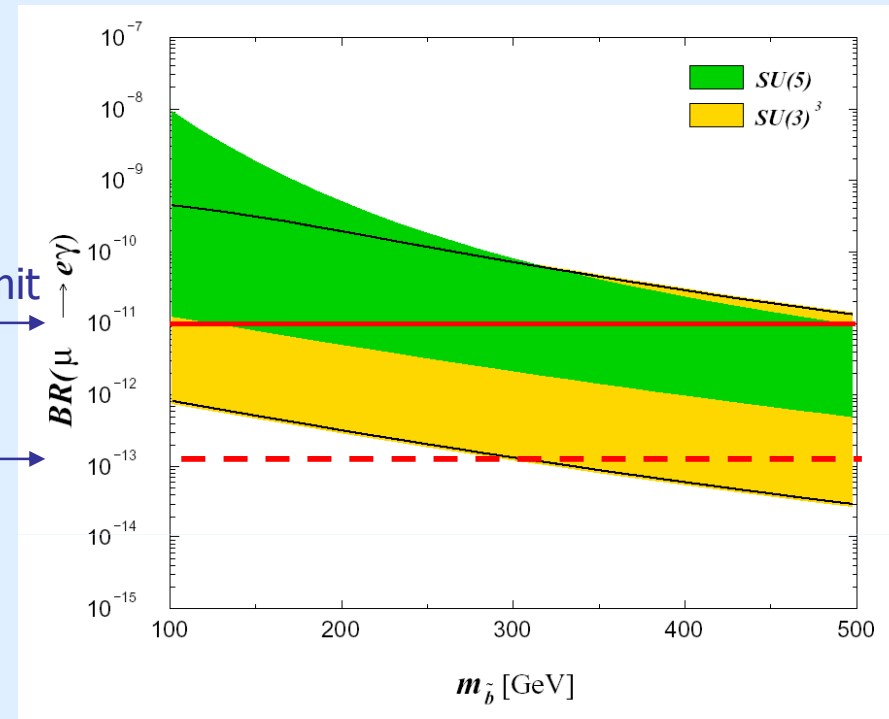


# Current SUSY predictions



current limit

MEG goal



"Supersymmetric parameterspace  
accessible by LHC"

- 1) J. Hisano et al., Phys. Lett. B391 (1997) 341
- 2) MEGA collaboration, hep-ex/9905013

W. Buchmueller, DESY, priv. comm.

# History of LFV searches



- Long history dating back to 1947!
- Best present limits:
  - $1.2 \times 10^{-11}$  (MEGA)
  - $\mu Ti \rightarrow e Ti < 7 \times 10^{-13}$  (SINDRUM II)
  - $\mu \rightarrow eee < 1 \times 10^{-12}$  (SINDRUM II)
- MEG Experiment aims at  $10^{-13}$
- Improvements linked to **advance in technology**

**SUSY SU(5)**

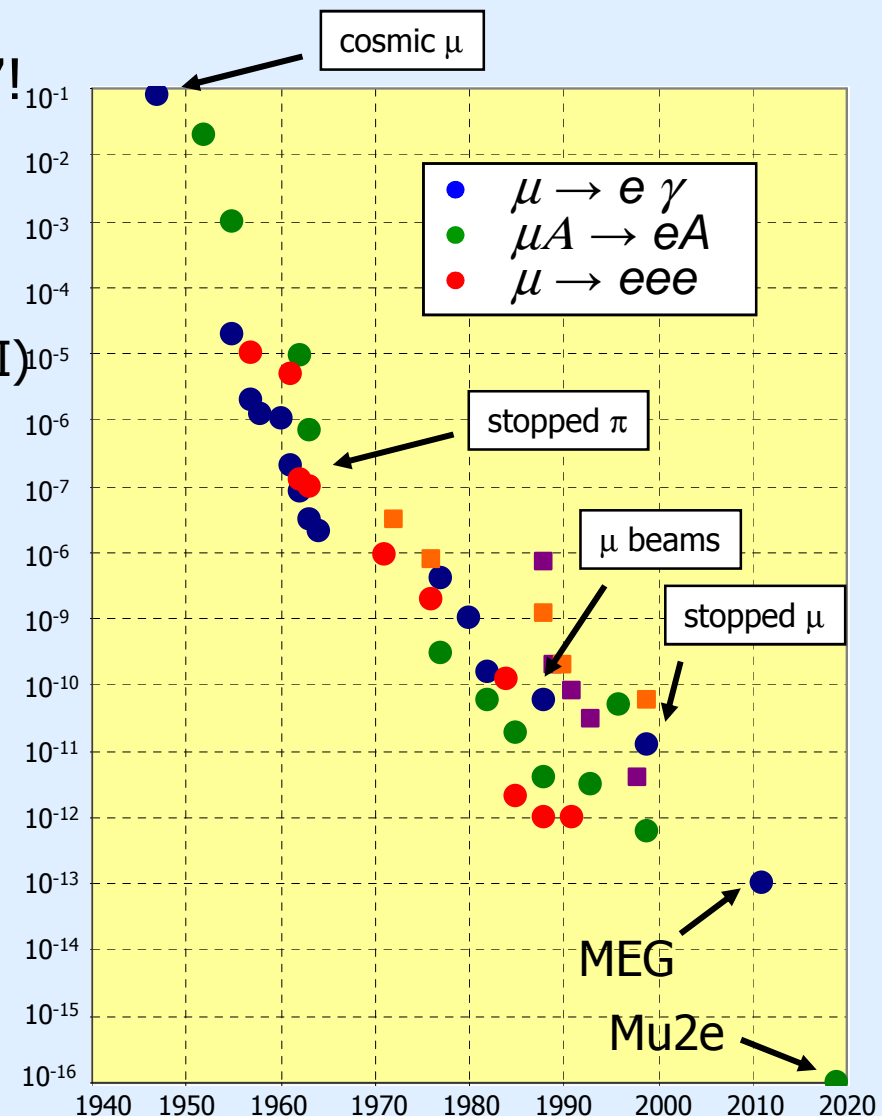
$$\text{BR}(\mu \rightarrow e \gamma) = 10^{-13}$$

$$\Leftrightarrow$$

$$\mu Ti \rightarrow e Ti = 4 \times 10^{-16}$$

$$\Leftrightarrow$$

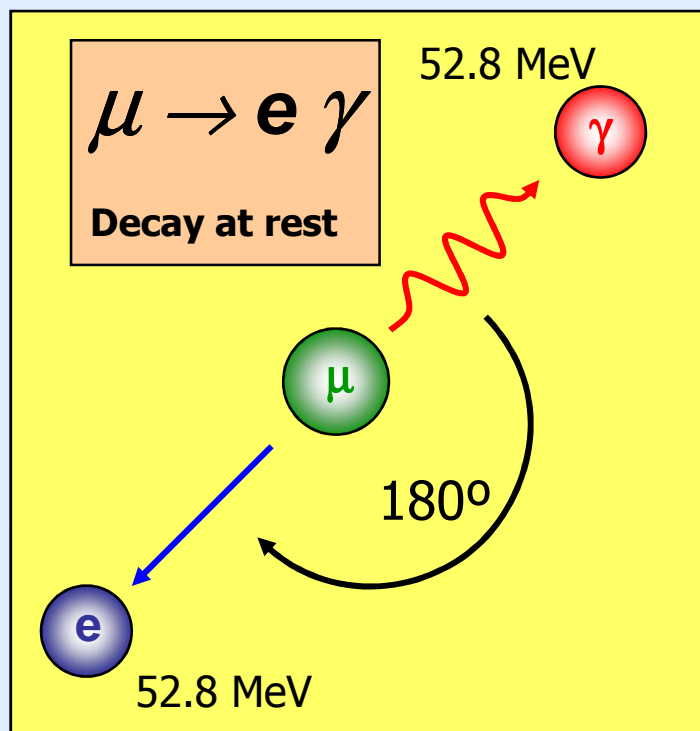
$$\text{BR}(\mu \rightarrow eee) = 6 \times 10^{-16}$$



# Experimental Method

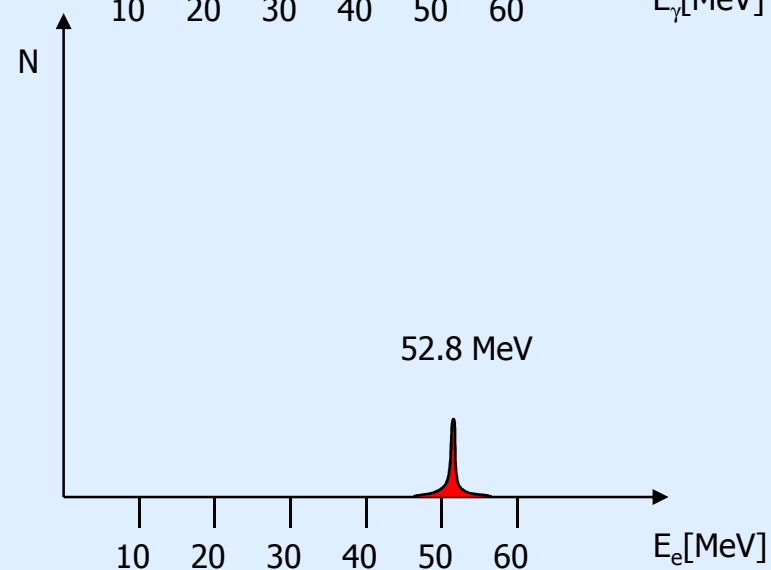
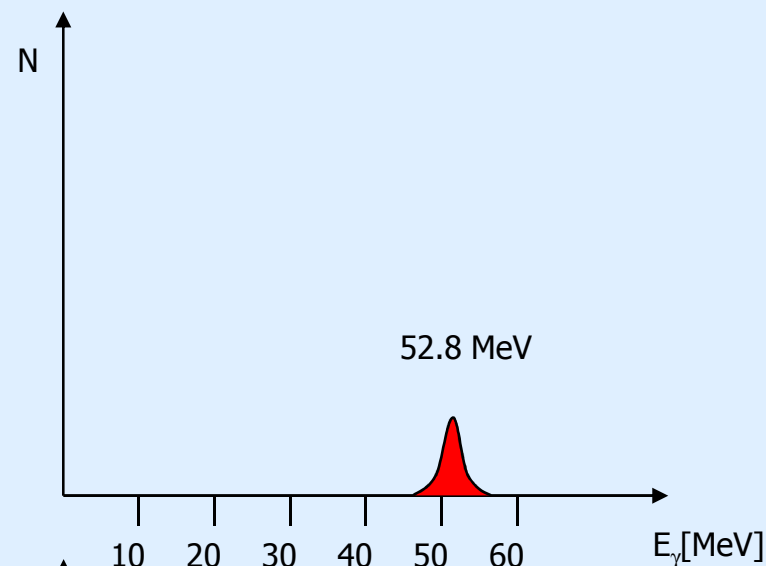
How to detect  $\mu \rightarrow e \gamma$ ?

# Decay Topology $\mu \rightarrow e \gamma$



$\mu \rightarrow e \gamma$  signal very clean

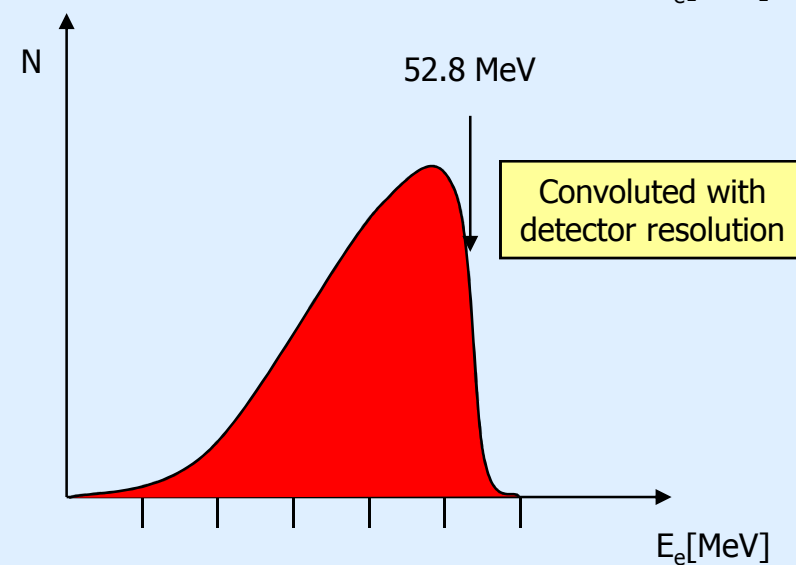
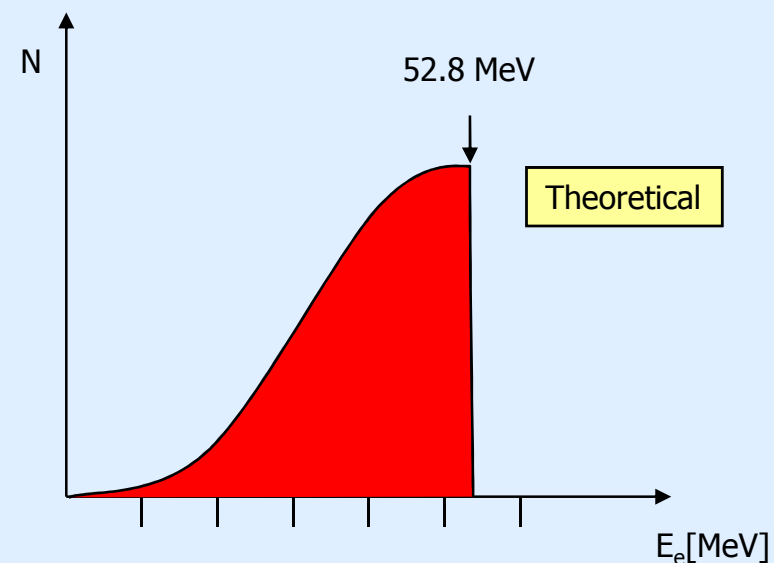
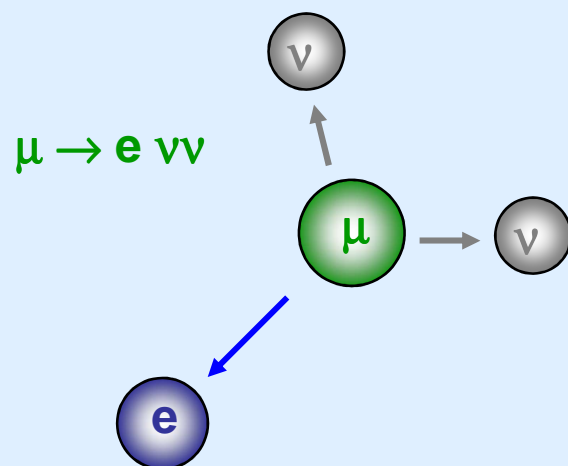
- $E_g = E_e = 52.8 \text{ MeV}$
- $\theta_{\gamma e} = 180^\circ$
- $e$  and  $\gamma$  in time



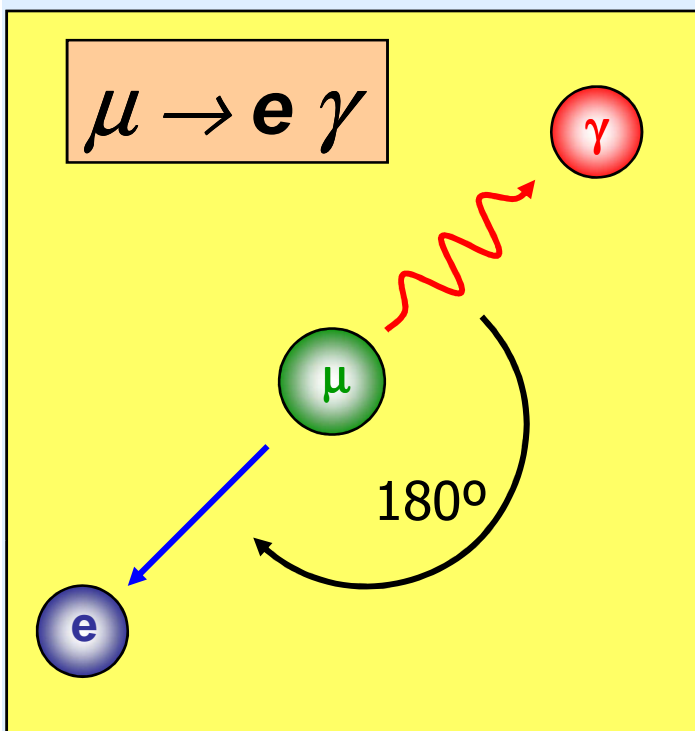
# Michel Decay ( $\sim 100\%$ )



Three body decay: wide energy spectrum

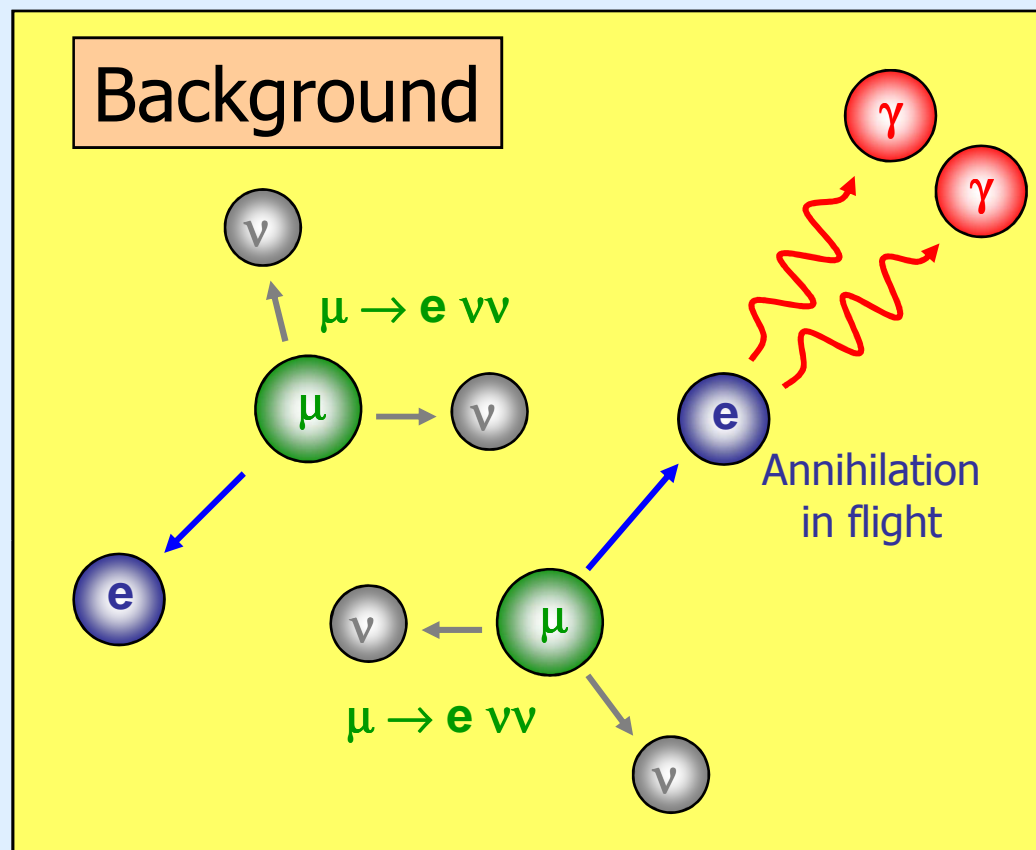


# "Accidental" Background



$\mu \rightarrow e \gamma$  signal very clean

- $E_g = E_e = 52.8 \text{ MeV}$
- $\theta_{\gamma e} = 180^\circ$
- $e$  and  $\gamma$  in time



- ➡ Good energy resolution
- ➡ Good spatial resolution
- ➡ Excellent timing resolution
- ➡ Good pile-up rejection



Aimed experiment parameters:

$N_\mu$	$3 \times 10^7 / \text{s}$
$T$	$2 \times 10^7 \text{ s}$ (~50 weeks)
$\Omega/4\pi$	0.09
$\epsilon_e$	0.90
$\epsilon_\gamma$	0.60
$\epsilon_{\text{sel}}$	0.70

Aimed resolutions:

	FWHM
$\Delta E_e$	0.8%
$\Delta E_\gamma$	4.3%
$\Delta \theta_{e\gamma}$	18 mrad
$\Delta t_{e\gamma}$	180 ps



Single event sensitivity  $(N_\mu \cdot T \cdot \Omega/4\pi \cdot \epsilon_e \cdot \epsilon_\gamma \cdot \epsilon_{\text{sel}})^{-1} = 3.6 \times 10^{-14}$

Prompt Background  $B_{\text{pr}} \cong 10^{-17}$

Accidental Background  $B_{\text{acc}} \propto \Delta E_e \cdot \Delta t_{e\gamma} \cdot (\Delta E_\gamma)^2 \cdot (\Delta \theta_{e\gamma})^2 \rightarrow 4 \times 10^{-14}$

90% C.L. Sensitivity  $\rightarrow 1.3 \times 10^{-13}$





Tokyo U.  
Waseda U.  
KEK

INFN & Uni  
Pisa  
Roma  
Genova  
Pavia  
Lecce

PSI

UC Irvine

JINR Dubna  
BINP Novosibirsk

March 9th, 2009

UVa Seminar

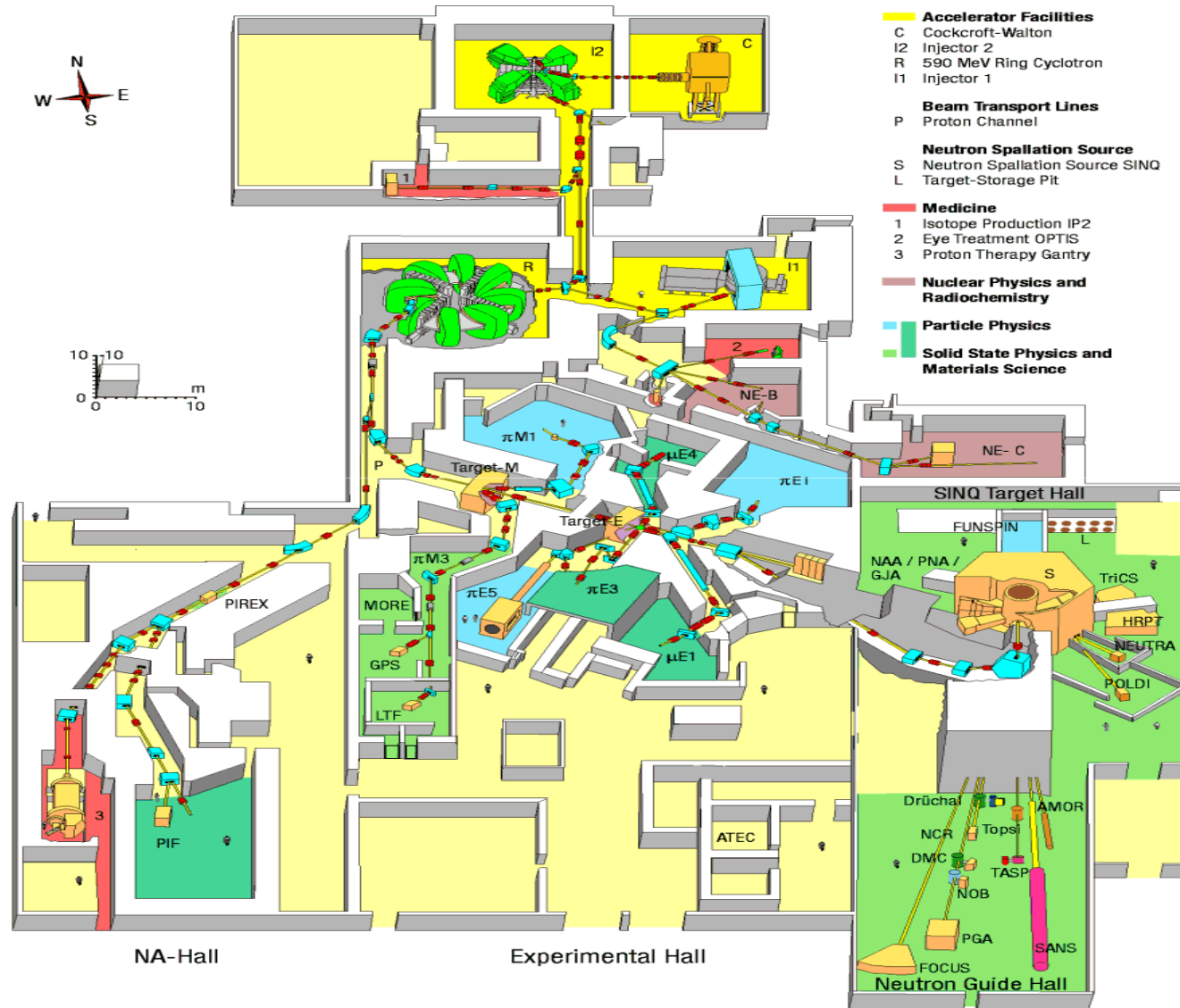
19



Proton Accelerator

Swiss Light Source

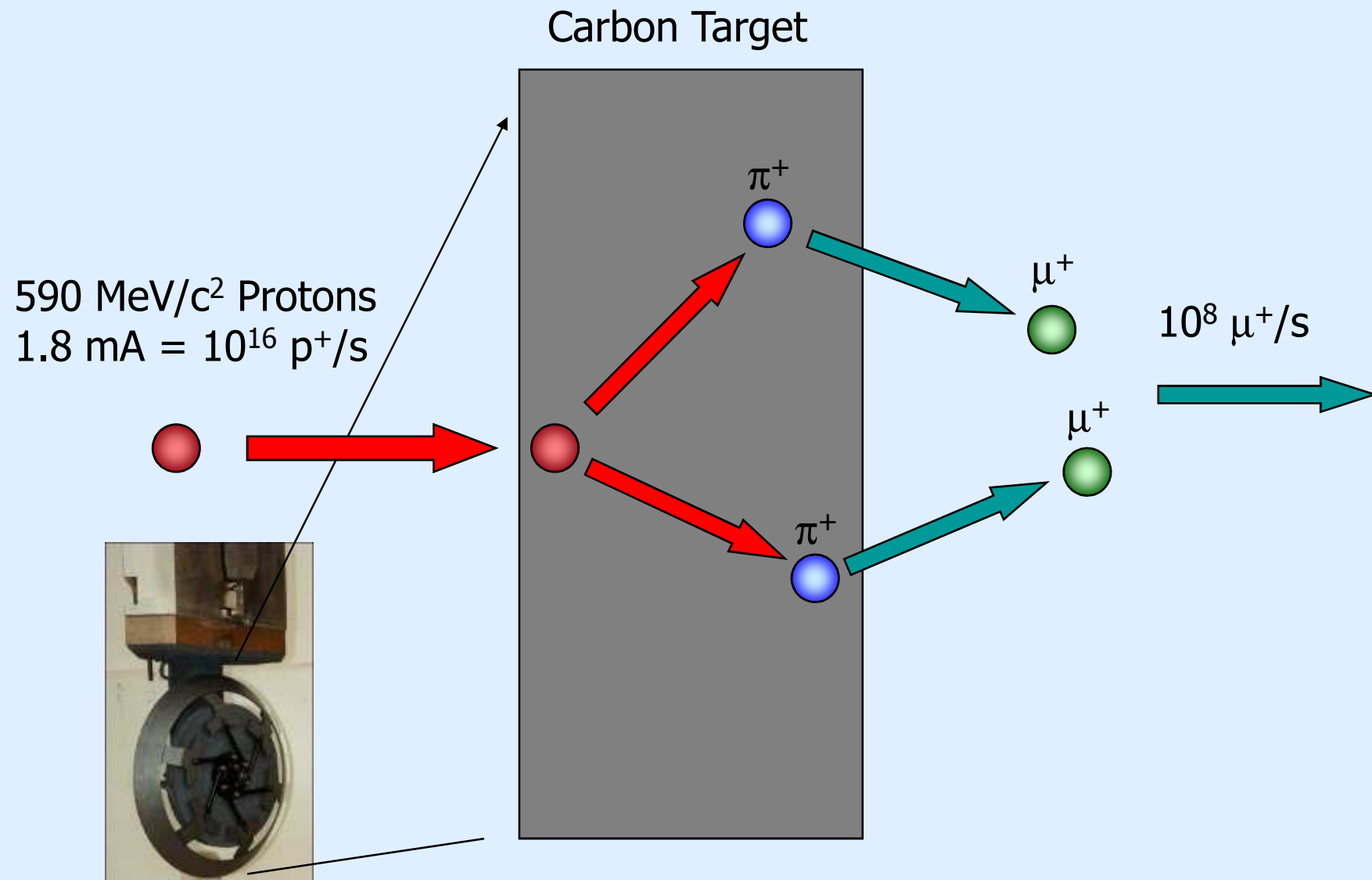


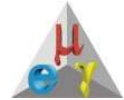


# Challenge 1: Muon Beam

How to get  $10^8 \mu/\text{sec}$  on a small stopping target?

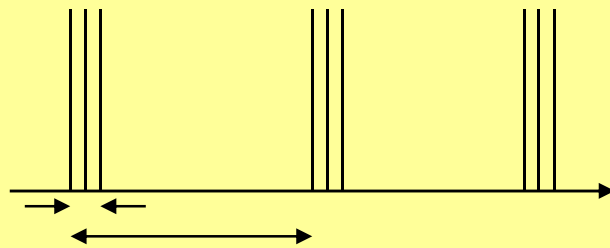
# Generating muons





Muon beam structure differs for different accelerators

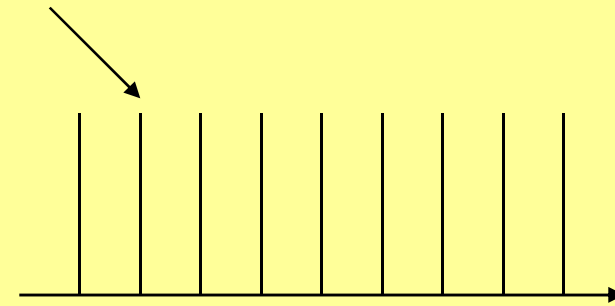
Pulsed muon beam, LANL



Duty cycle: Ratio of pulse width over period

Duty cycle: 6 %

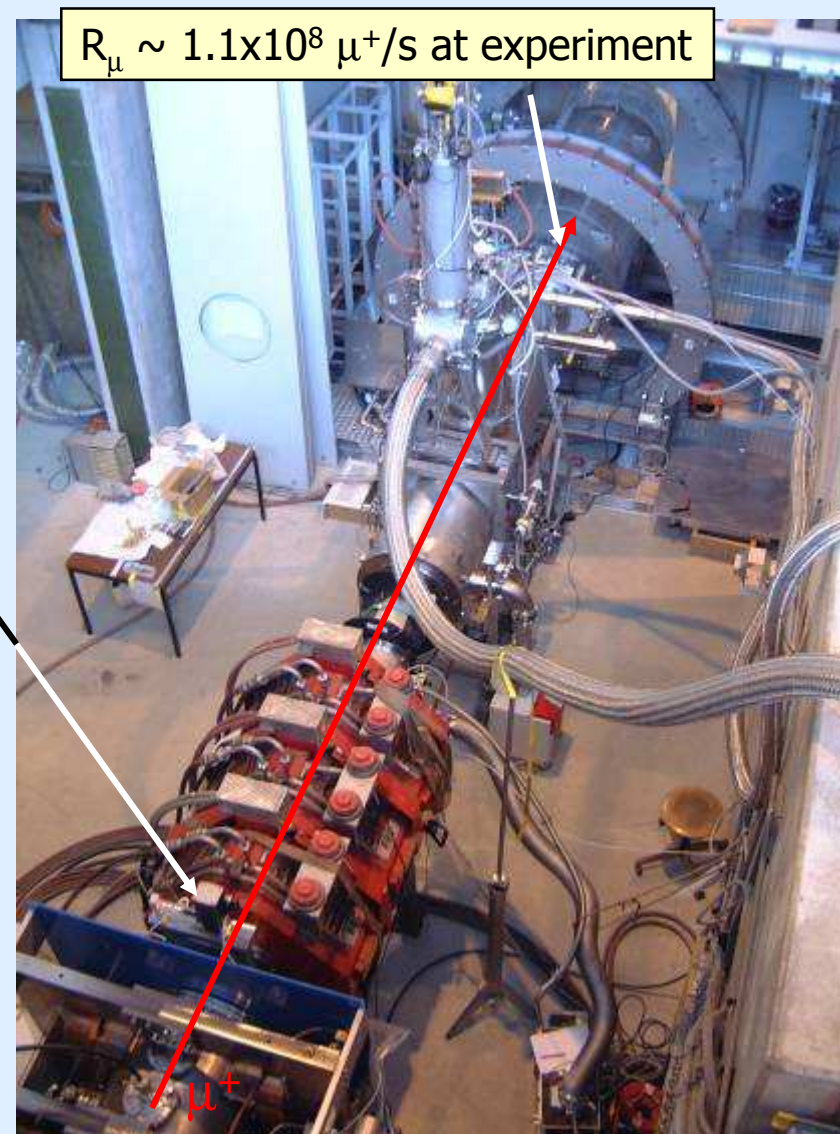
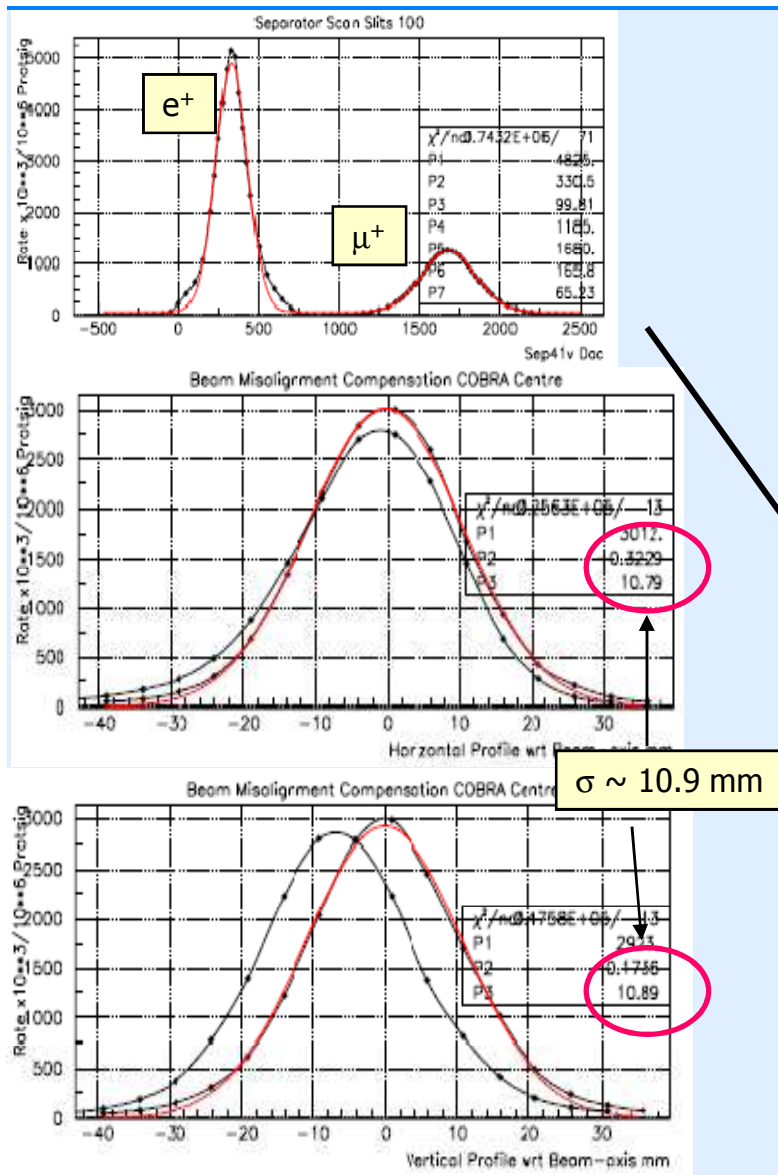
DC muon beam, PSI



Duty cycle: Ratio of pulse width over period

Duty cycle: 100 %

# Results of beam line optimization



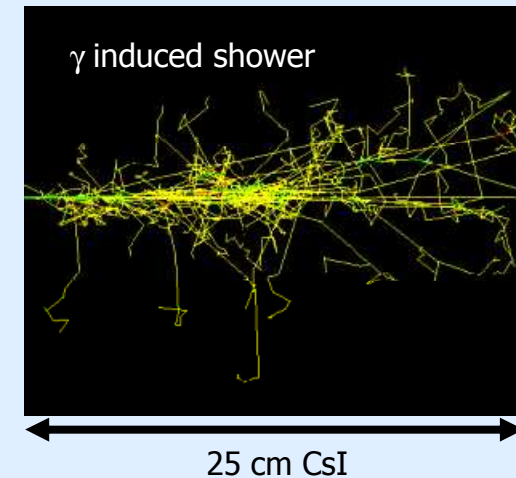
# Challenge 2: Calorimeter

$$\mu \rightarrow e \gamma$$

Energy  
Position  
Time

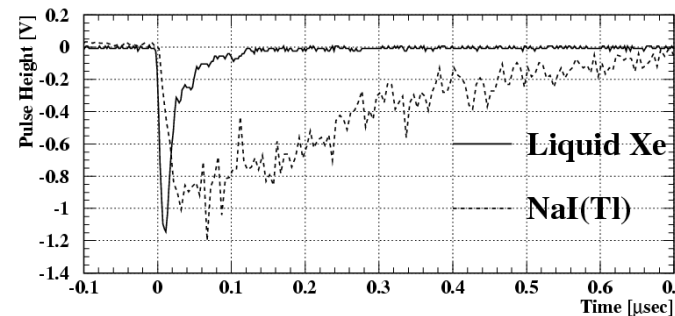
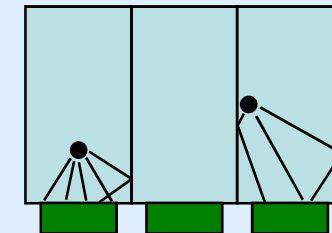


- Anorganic crystals (NaI, CsI):
  - Good efficiency, good energy resolution, poor position resolution, poor homogeneity
- Liquid Noble Gases:
  - No crystal boundaries
  - Good efficiency, resolutions



## Liquid Xenon:

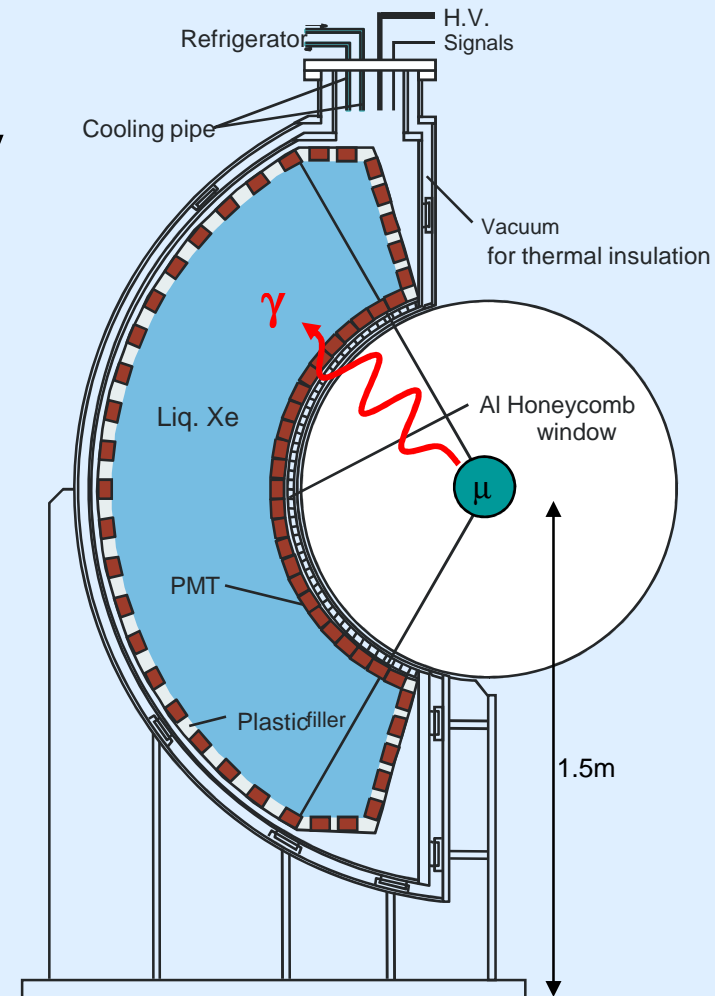
Density	3 g/cm <sup>3</sup>
Melting/boiling point	161 K / 165 k
Radiation length	2.77 cm
Decay time	45 ns
Absorption length	> 100 cm
Refractive index	1.57



# Liquid Xenon Calorimeter

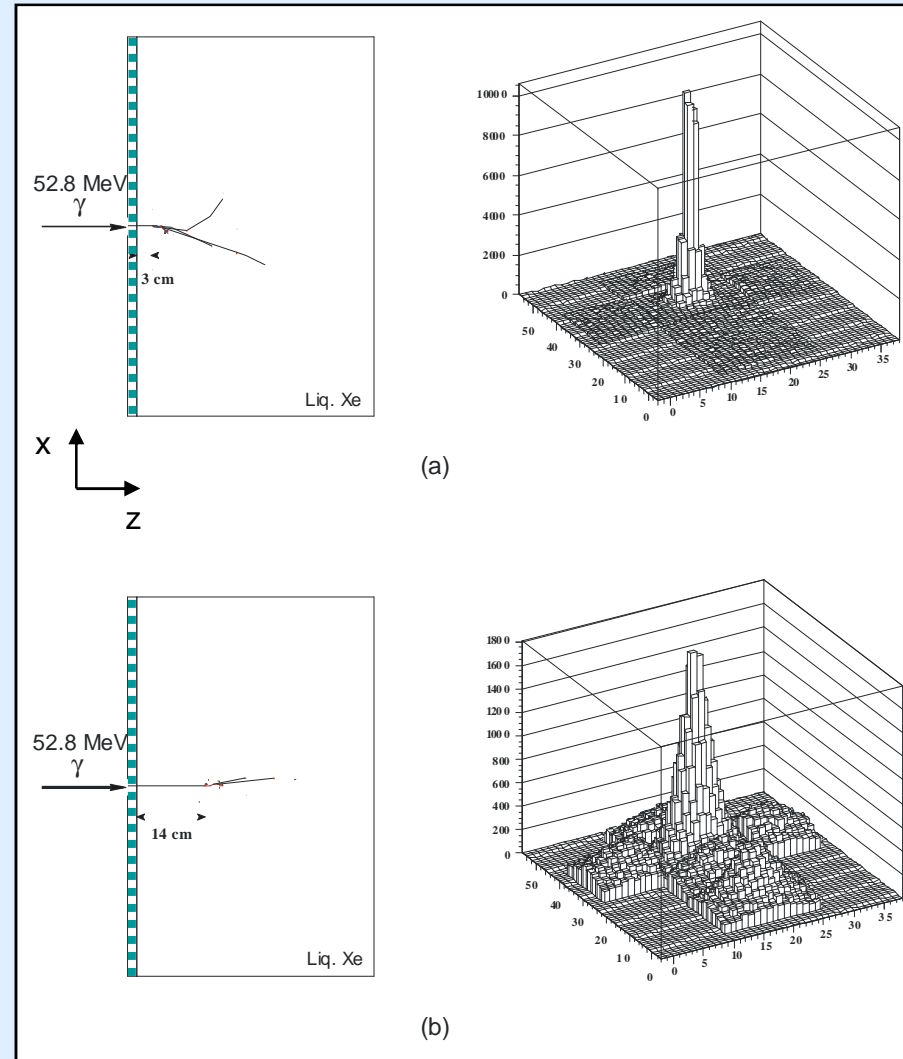


- Calorimeter: Measure  $\gamma$  Energy, Position and Time
- Liquid Xenon has high Z and homogeneity
- $\sim 900$  l (3t) Xenon with  $\sim 850$  PMTs
- Cryogenics required:  $-120^{\circ}\text{C} \dots -108^{\circ}$
- Extremely high purity necessary:  
1 ppm  $\text{H}_2\text{O}$  absorbs 90% of light
- Currently largest LXe detector in the world: Lots of pioneering work necessary





- Light is distributed over many PMTs
- Weighted mean of PMTs on front face  
→  $dx \sim 10$  mm FWHM
- Broadness of distribution  
→  $dz \sim 16$  mm FWHM
- Timing resolution  
→  $dt \sim 100$  ps FWHM
- Energy resolution  
~ 4.3% FWHM  
depends on light attenuation in LXe





- Use “Monte Carlo” simulation (GEANT) to carefully study detector
- Placement of PMTs were optimized according to MC results

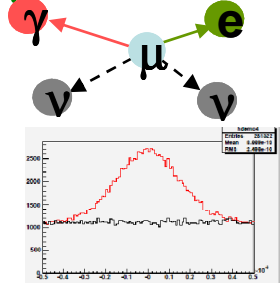
Currently being assembled, will go into operation summer '07



# Calorimeter Calibrations



## $\mu$ radiative decay

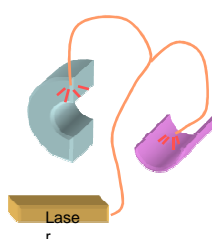


Lower beam intensity  $< 10^7$   
to reduce pile-ups  
A few days ~ 1 week to get  
enough statistics

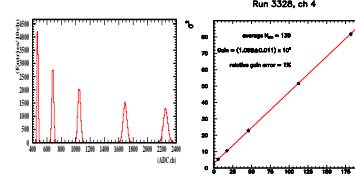
## Laser

(rough)  
relative  
timing calib.

$< 2\sim 3$  nsec



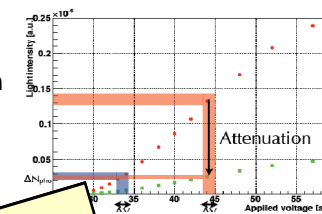
## LED



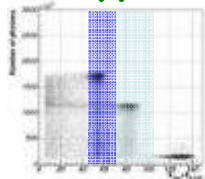
## PMT Gain

Higher V with  
light att.

Can be  
repeated  
frequently



## $\pi^0 \rightarrow \gamma\gamma$



$\pi + p \rightarrow \pi^0 + n$

$\pi^0 \rightarrow \gamma\gamma$  (55MeV, 83MeV)

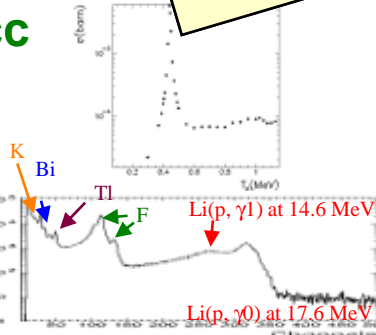
$\pi + p \rightarrow \gamma + n$  (129MeV)

10 days to scan all  
volume precisely



LH<sub>2</sub> target

## Proton Acc



## $\text{Li}(p,\gamma)\text{Be}$

LiF target at  
COBRA center

17.6MeV  $\gamma$

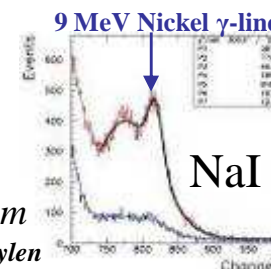
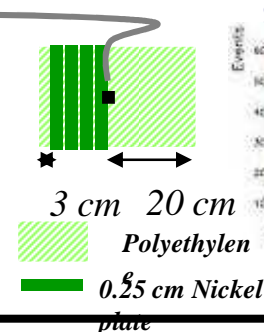
~daily calib.

Can be used  
also for initial  
setup

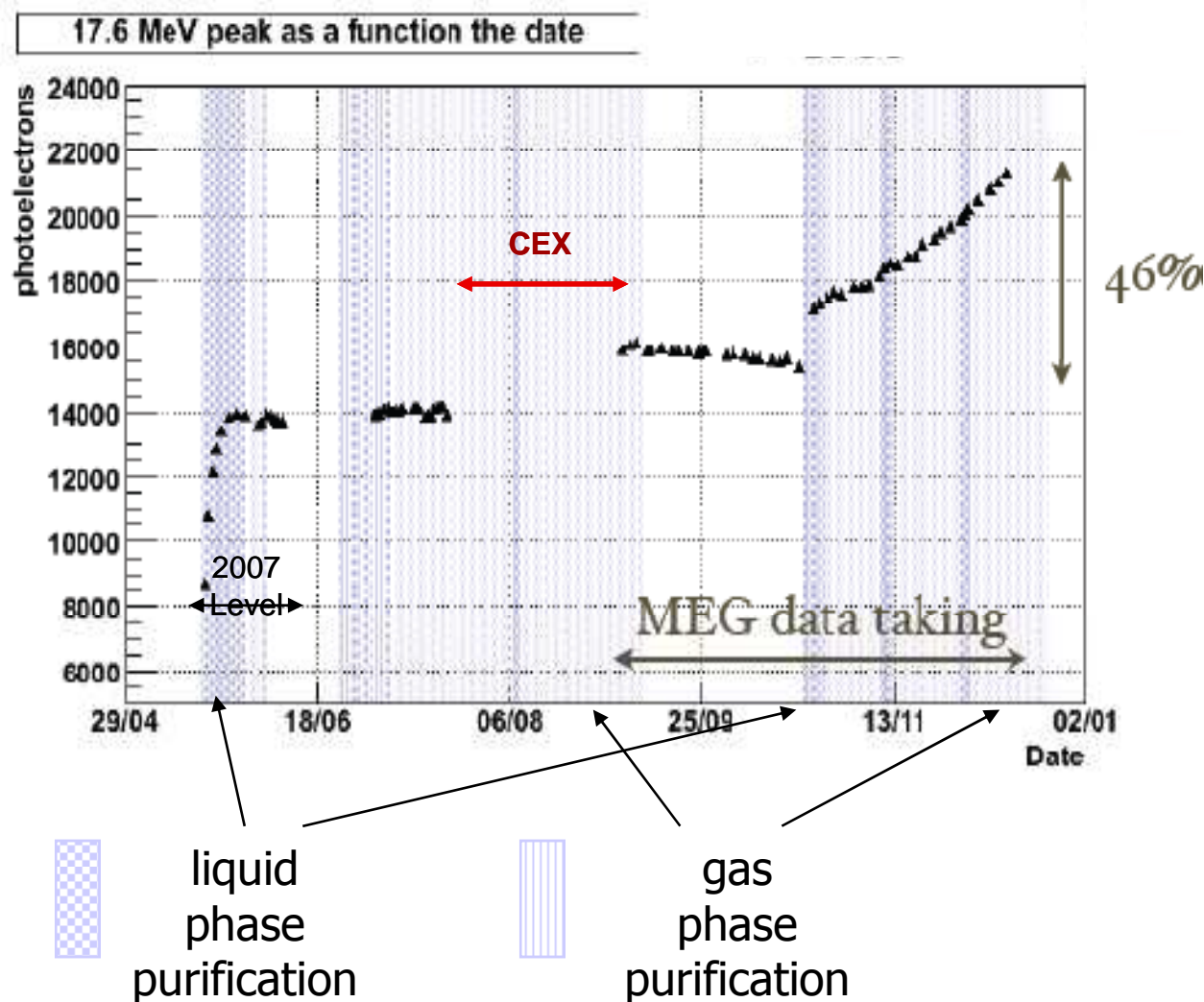
## Nickel $\gamma$ Generator

off on  
Illuminate Xe  
from the back

Source (Cf)  
transferred by  
comp air  $\rightarrow$   
on/off



**Dedicated Talk!**



# Challenge 2: Spectrometer

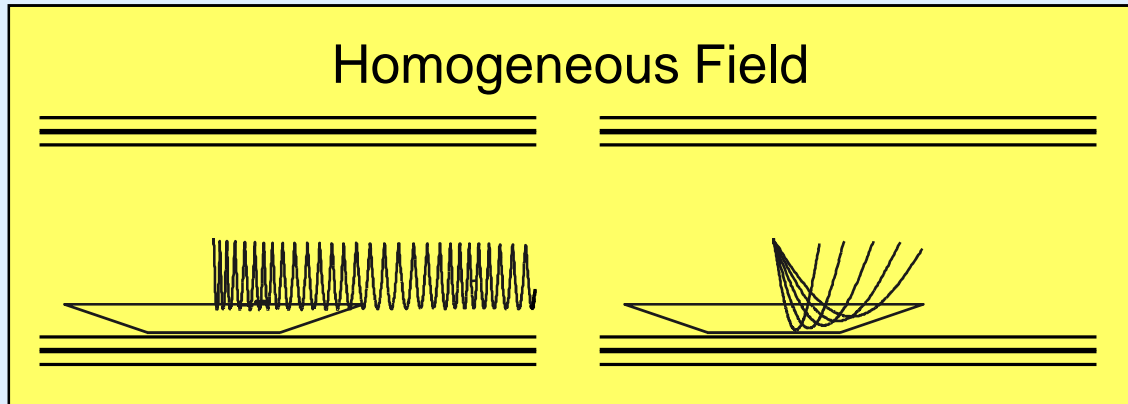
$$\mu \rightarrow e \gamma$$

Energy  
Position

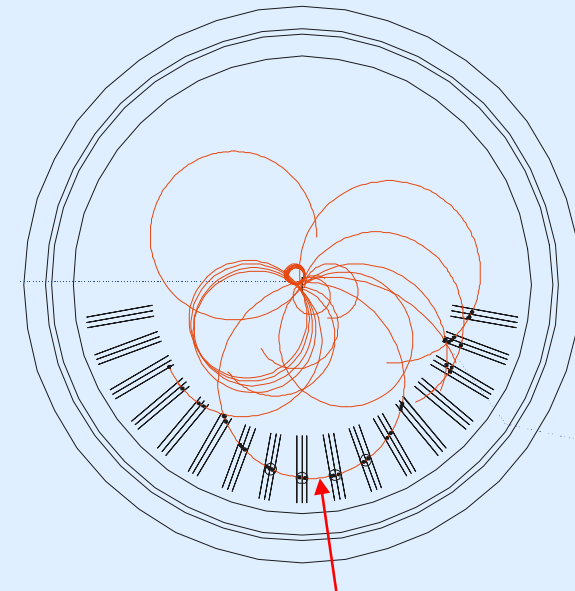
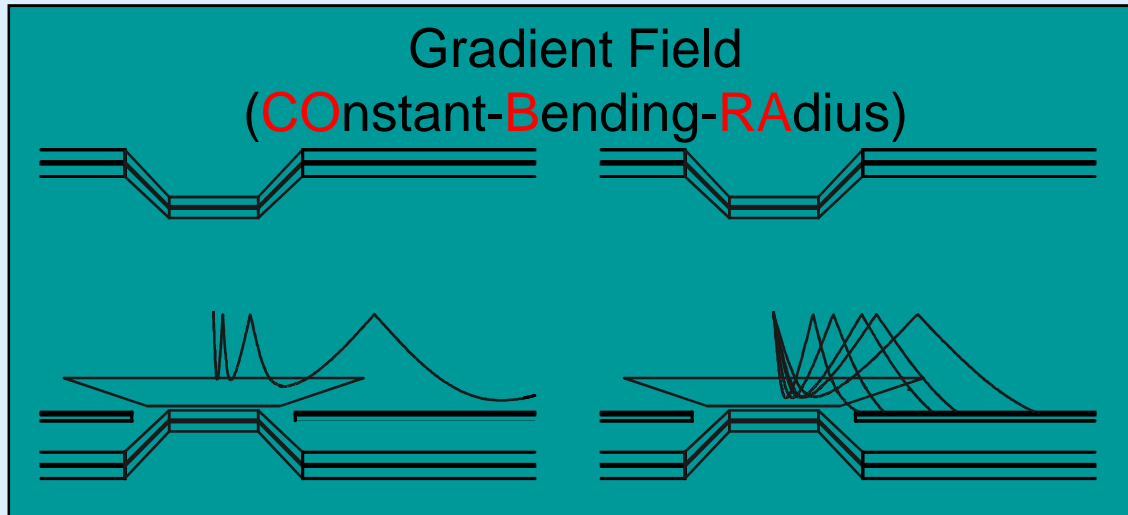


Ultra-thin ( $\sim 3\text{g/cm}^2$ ) superconduction solenoid with 1.2 T magnetic field

## Homogeneous Field

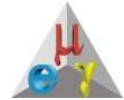


## Gradient Field (Constant-Bending-Radius)

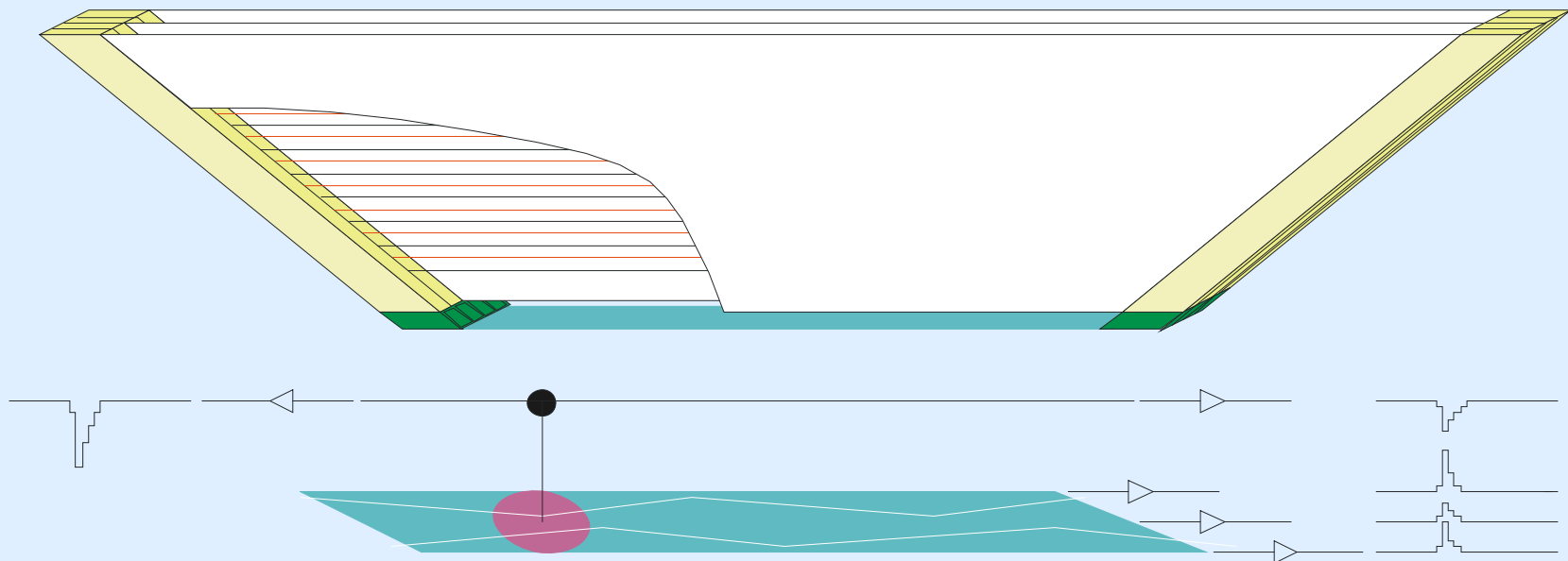


$e^+$  from  $\mu^+ \rightarrow e^+ \gamma$

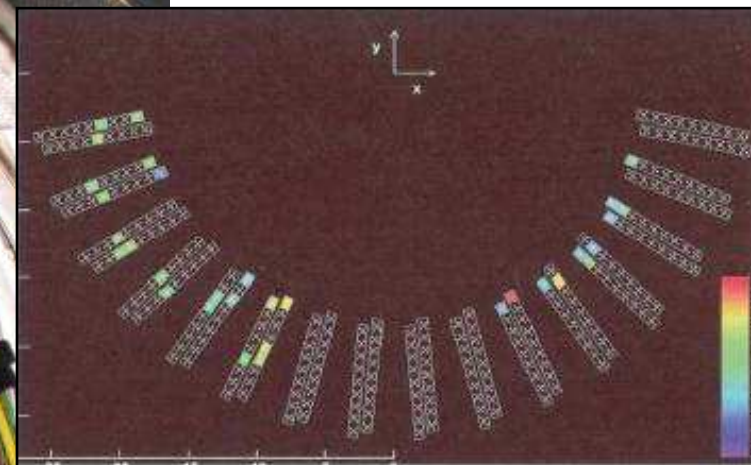
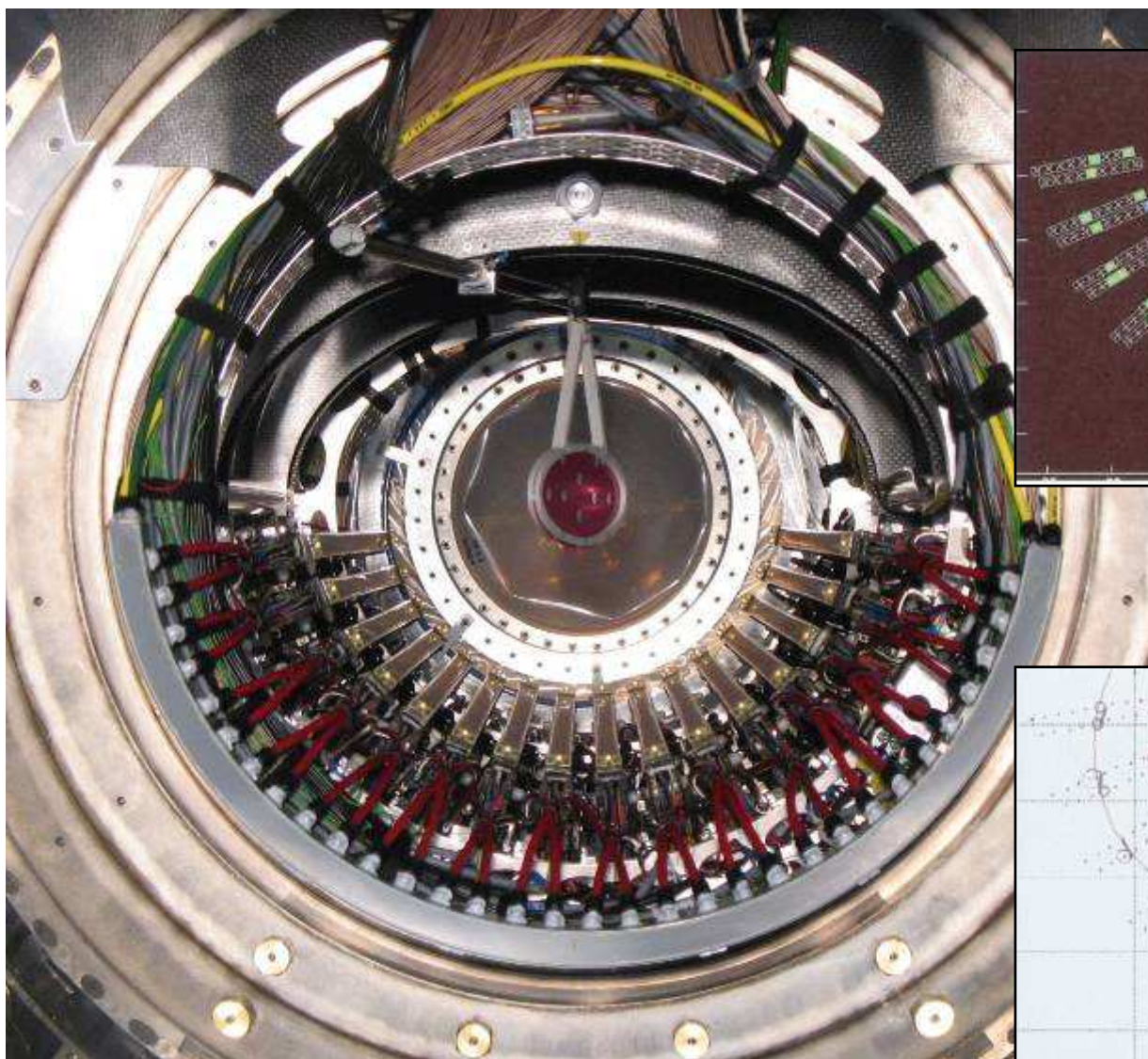
# Drift Chamber



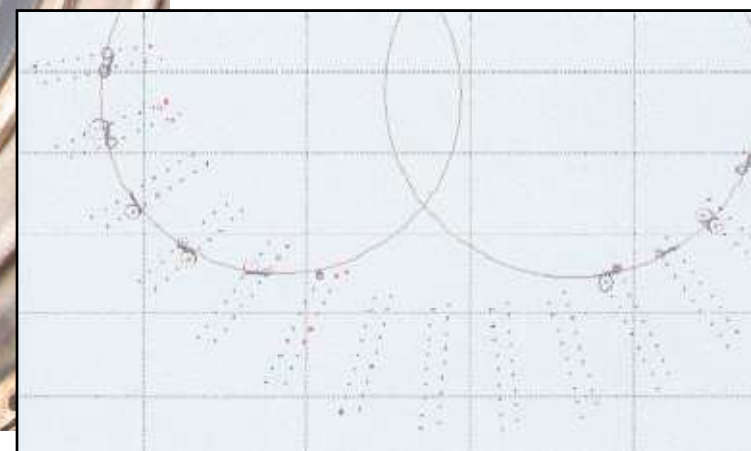
- Measures position, time and curvature of positron tracks
- Cathode foil has three segments in a vernier pattern → Signal ratio on vernier strips to determine coordinate along wire



# Final Spectrometer



$\sigma_R = 470 \text{ } \mu\text{m}$  (FWHM)  
 $\sigma_Z = 760 \text{ } \mu\text{m}$  (FWHM)

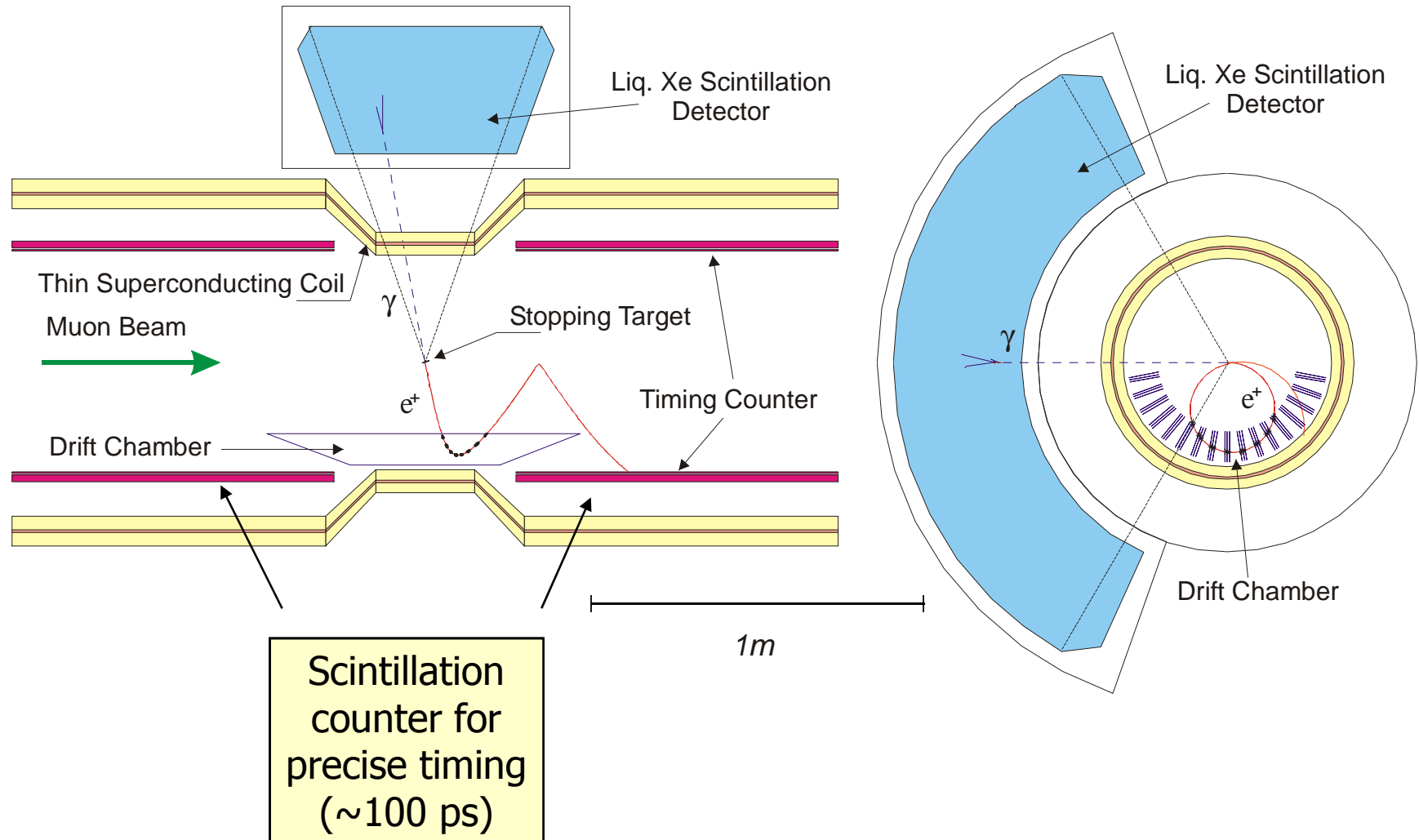


# Challenge 3: Timing Counter

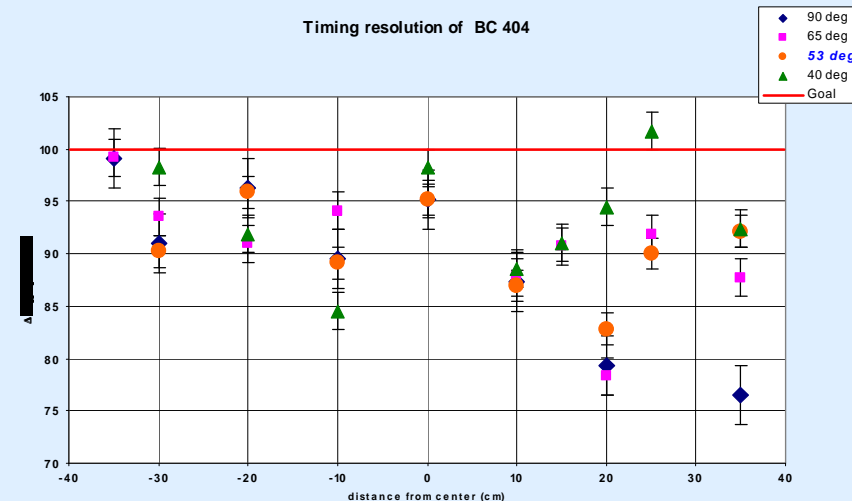
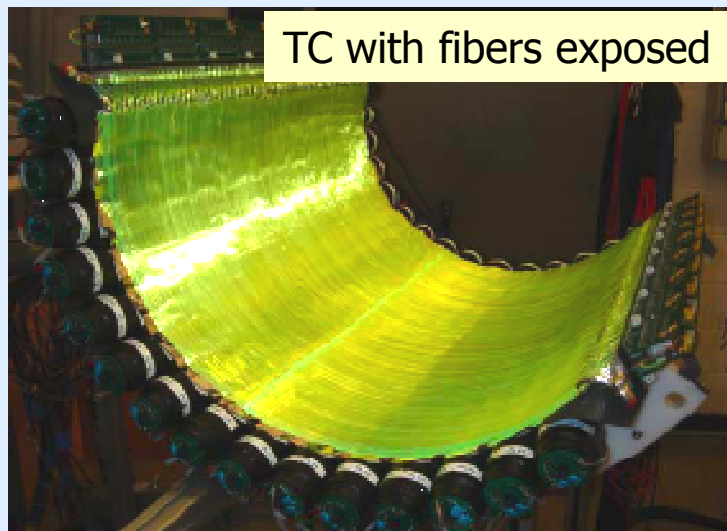
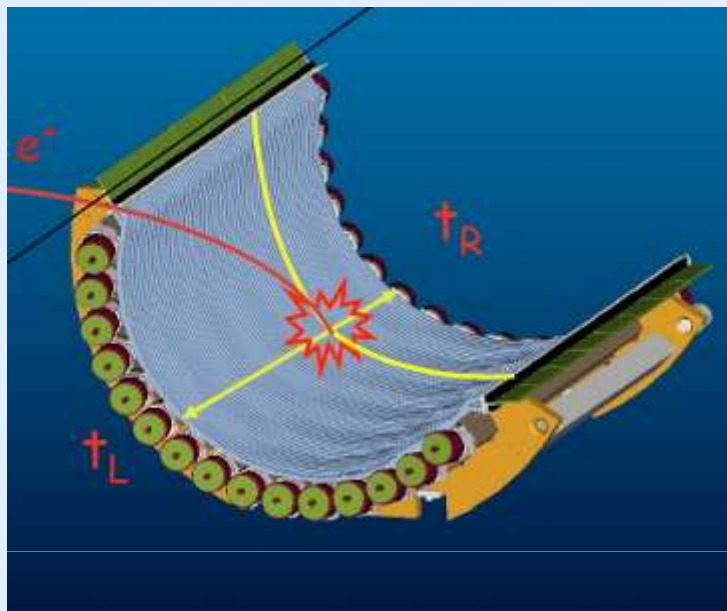
$$\mu \rightarrow e \gamma$$

Time

# Timing Counter Location



# Timing Counter



- Staves along beam axis for timing measurement
- Curved fibers with APD readout for z-position
- Resolution 90 ps FWHM measured at  $e^-$  - beam
- Resolution in experiment: 140-200 ps FWHM preliminary

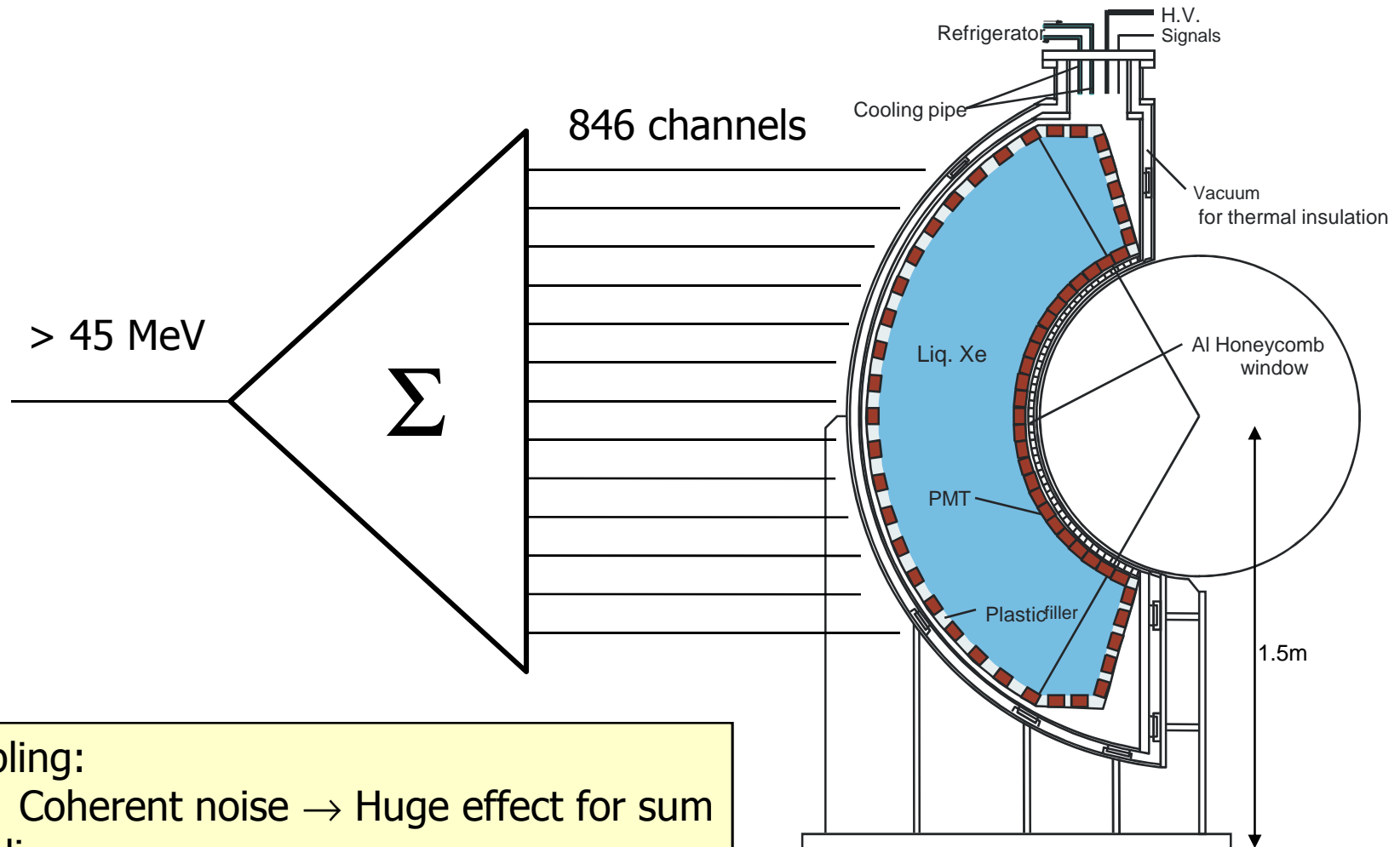
Exp./ Lab	Author	Year	$\Delta E_e/E_e$ %FWHM	$\Delta E_\gamma/E_\gamma$ %FWHM	$\Delta t_{e\gamma}$ (ns)	$\Delta\theta_{e\gamma}$ (mrad)	Inst. Stop rate (s <sup>-1</sup> )	Duty cycle (%)	Result
SIN (PSI)	A. Van der Schaaf	1977	8.7	9.3	1.4	-	$(4..6) \times 10^5$	100	$< 1.0 \times 10^{-9}$
TRIUMF	P. Depommier	1977	10	8.7	6.7	-	$2 \times 10^5$	100	$< 3.6 \times 10^{-9}$
LANL	W.W. Kinnison	1979	8.8	8	1.9	37	$2.4 \times 10^5$	6.4	$< 1.7 \times 10^{-10}$
Crystal Box	R.D. Bolton	1986	8	8	1.3	87	$4 \times 10^5$	(6..9)	$< 4.9 \times 10^{-11}$
MEGA	M.L. Brooks	1999	1.2	4.5	1.6	17	$2.5 \times 10^8$	(6..7)	$< 1.2 \times 10^{-11}$
MEG		<b>2008</b>	<b>0.8</b>	<b>4.3</b>	<b>0.18</b>	<b>18</b>	<b><math>3 \times 10^7</math></b>	<b>100</b>	<b><math>\sim 10^{-13}</math></b>

# Challenge 4: Electronics

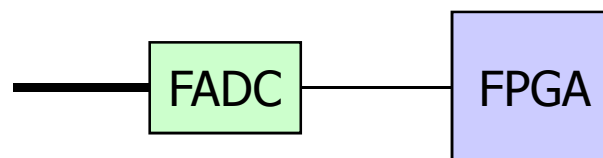
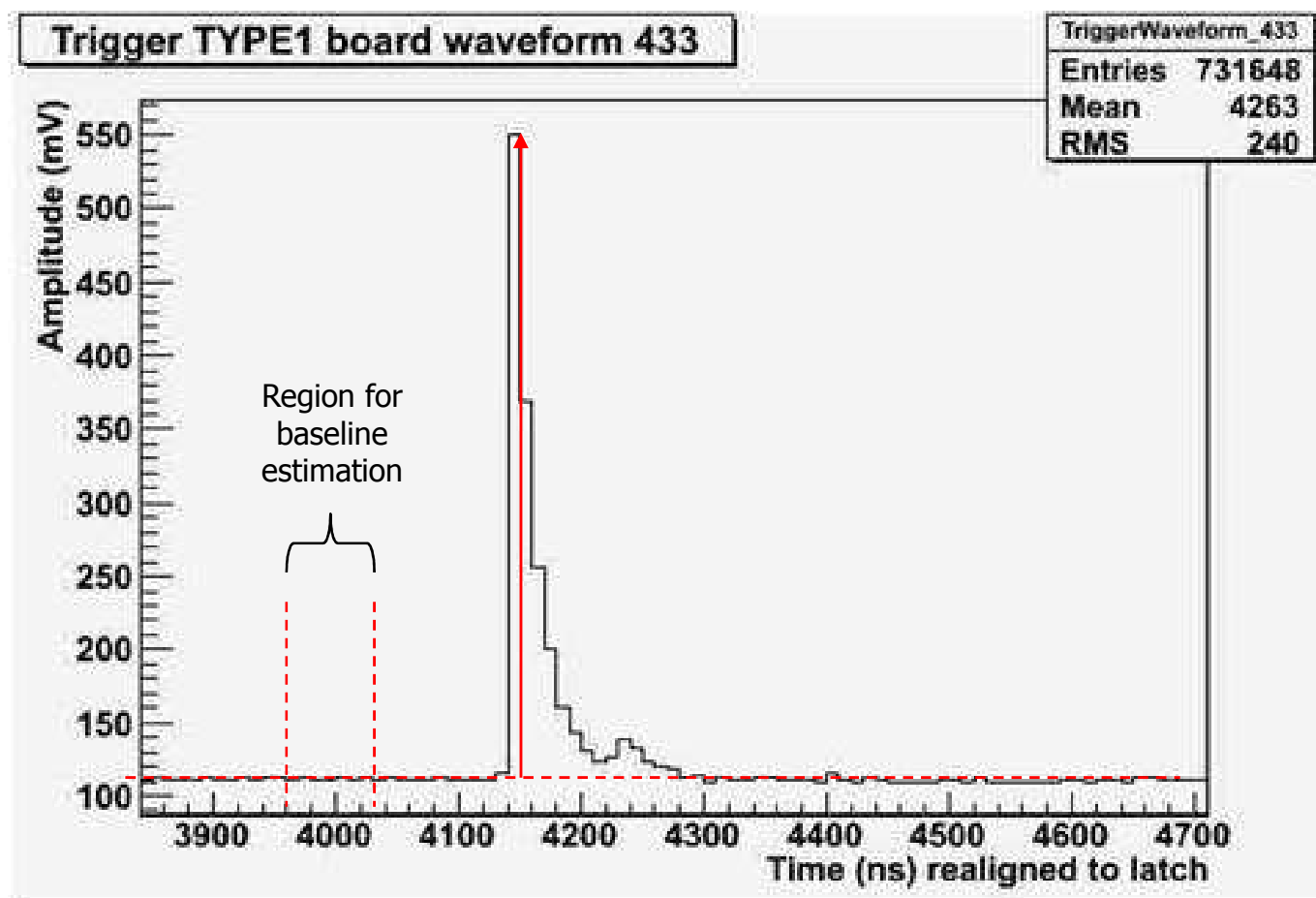
How to do effective triggering?

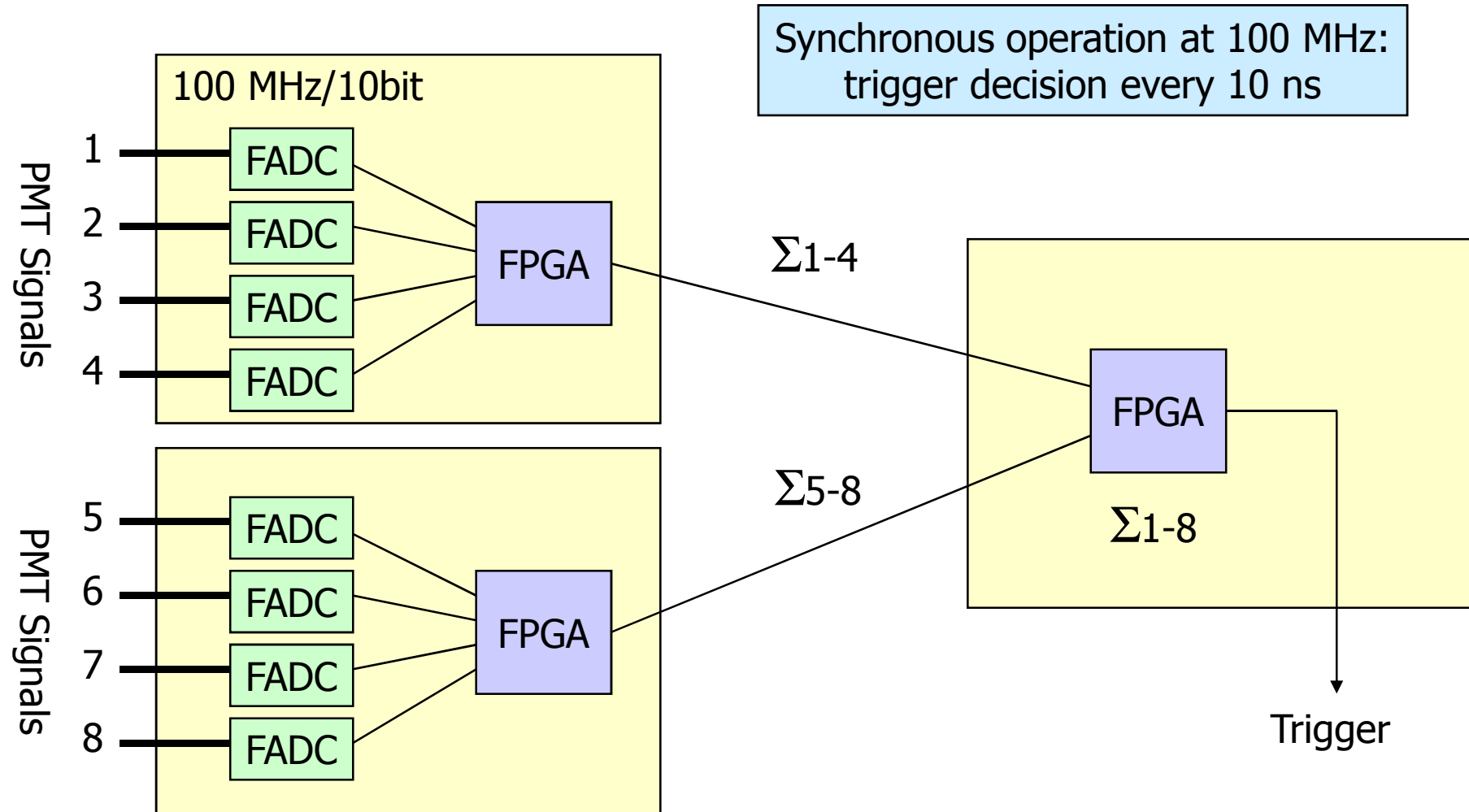
How to deal with pile-up?

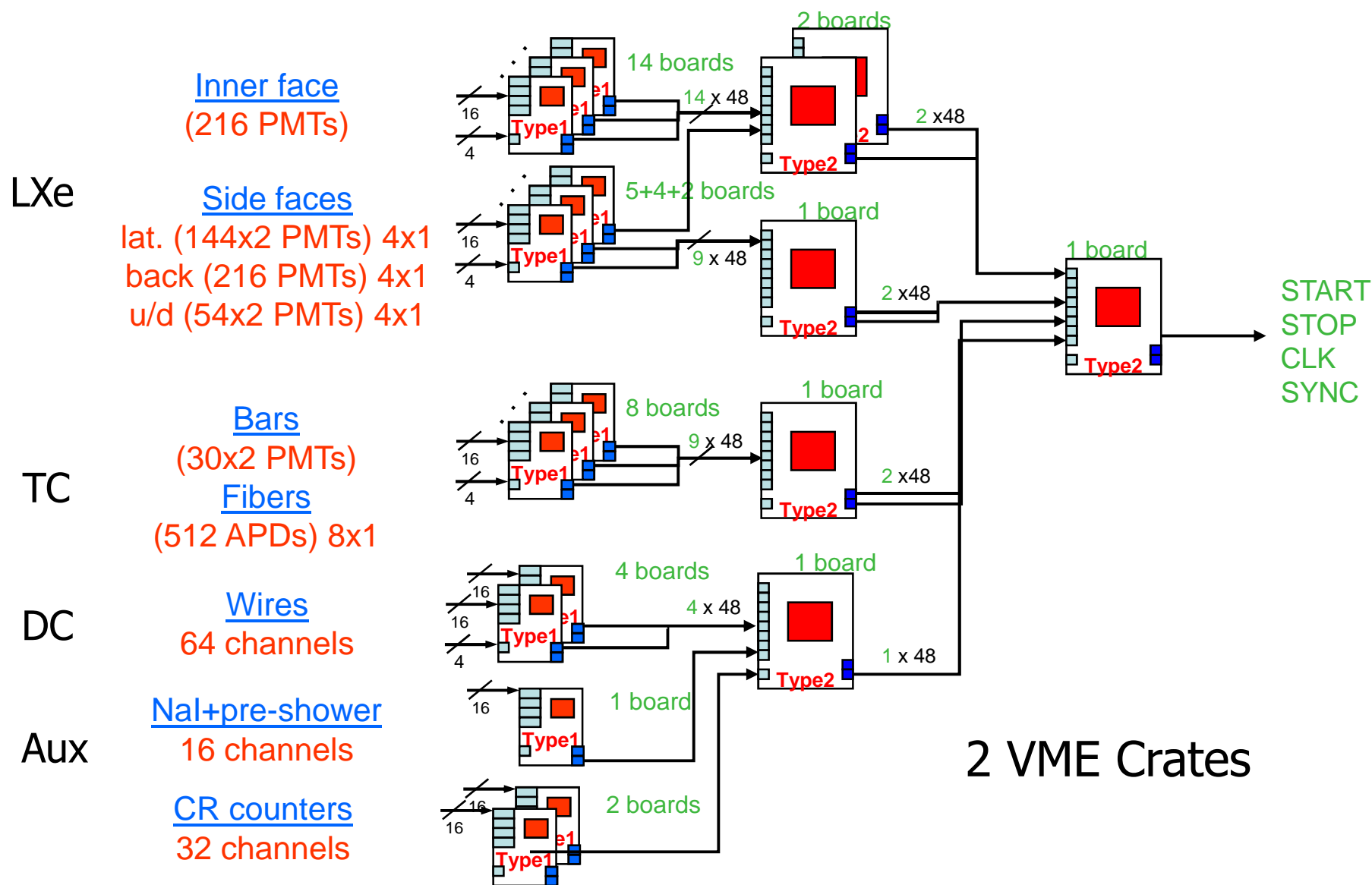
How to measure timing for  $\sim 1000$  channels with  $< 100$  ps accuracy?



DC-coupling:  
Coherent noise → Huge effect for sum  
AC-coupling:  
Varying baseline with intensity







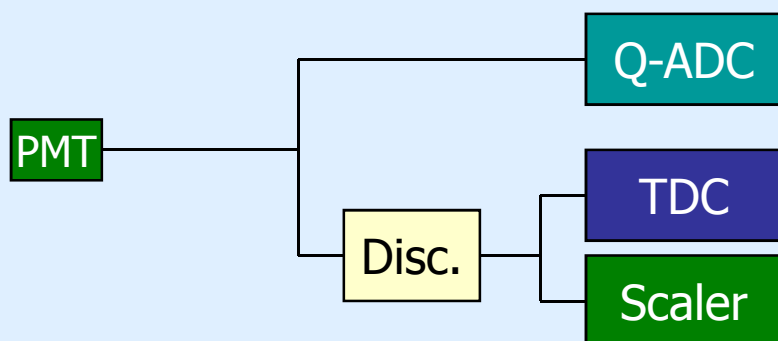


TRGDAQRateMeter					
Proton Current	Total trigger rate		Live Time		Live Time (%)
1988.4 $\mu$ Amp	6.096 Hz		3.738 sec		81.382
	#Ev(#DAQ)	EvRate(DAQ Rate,%)		#Ev(#DAQ)	EvRate(DAQ Rate,%)
Id0 MuEGamma	32 (26)	7.0Hz(5.7Hz,92.9)	Id16 Michel	1.3e+06 (0)	2.8e+05Hz(0.0Hz,0.0)
Id1 MEG LowQ	106 (0)	23.1Hz(0.0Hz,0.0)	Id17 DC Trackout	3.6e+06 (0)	7.8e+05Hz(0.0Hz,0.0)
Id2 MEG WidAng	170 (0)	37.0Hz(0.0Hz,0.0)	Id18 DC Track	3.7e+06 (0)	8.1e+05Hz(0.0Hz,0.0)
Id3 MEG WidTime	58 (0)	12.6Hz(0.0Hz,0.0)	Id19 DC Cosm	0 (0)	0.0Hz(0.0Hz,0.0)
Id4 Rad NarTime	326 (0)	71.0Hz(0.0Hz,0.0)	Id20 DC single	6.4e+06 (0)	1.4e+06Hz(0.0Hz,0.0)
Id5 Rad WidTime	614 (0)	133.7Hz(0.0Hz,0.0)	Id21 Cosm Alone	0 (0)	0.0Hz(0.0Hz,0.0)
Id6 Pi0	0 (0)	0.0Hz(0.0Hz,0.0)	Id22 TC Alone	5.0e+06 (0)	1.1e+06Hz(0.0Hz,0.0)
Id7 Pi0 NPrSh	0 (0)	0.0Hz(0.0Hz,0.0)	Id23 CR Coinc	0 (0)	0.0Hz(0.0Hz,0.0)
Id8 NaI	2 (0)	0.4Hz(0.0Hz,0.0)	Id24 TC Pair	4.6e+05 (0)	1.0e+05Hz(0.0Hz,0.0)
Id9 LXe HighQ	1.2e+04 (1)	2.6e+03Hz(0.2Hz,3.6)	Id25 NaI Cosmic	0 (0)	0.0Hz(0.0Hz,0.0)
Id10 LXe LowQ	4.9e+04 (0)	1.1e+04Hz(0.0Hz,0.0)	Id26 APD Single	7.2e+06 (0)	1.6e+06Hz(0.0Hz,0.0)
Id11 CW Bo	1.1e+04 (0)	2.4e+03Hz(0.0Hz,0.0)	Id27 LXe Cosmic	856 (1)	186.4Hz(0.2Hz,3.6)
Id12 Alpha	3.9e+04 (0)	8.5e+03Hz(0.0Hz,0.0)	UNUSED	32 (0)	7.0Hz(0.0Hz,0.0)
Id13 Laser	0 (0)	0.0Hz(0.0Hz,0.0)	UNUSED	106 (0)	23.1Hz(0.0Hz,0.0)
Id14 LED	4 (0)	0.9Hz(0.0Hz,0.0)	UNUSED	0 (0)	0.0Hz(0.0Hz,0.0)
Id15 NeutronNi	0 (0)	0.0Hz(0.0Hz,0.0)	Id31 Pedestal	4.6e+03 (0)	1.0e+03Hz(0.0Hz,0.0)

# How to digitize signals?



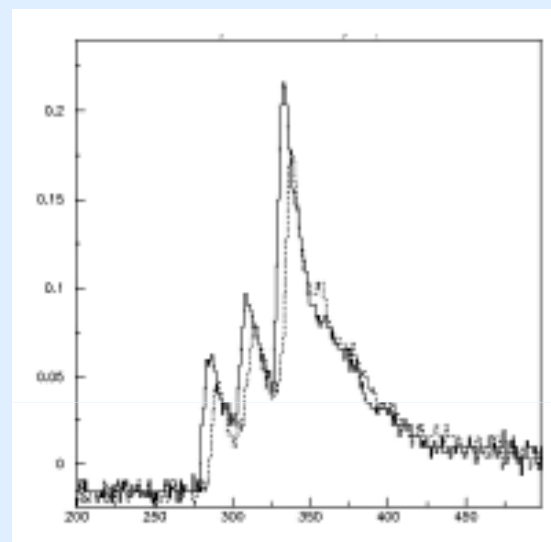
“Old fashioned”



“Modern”

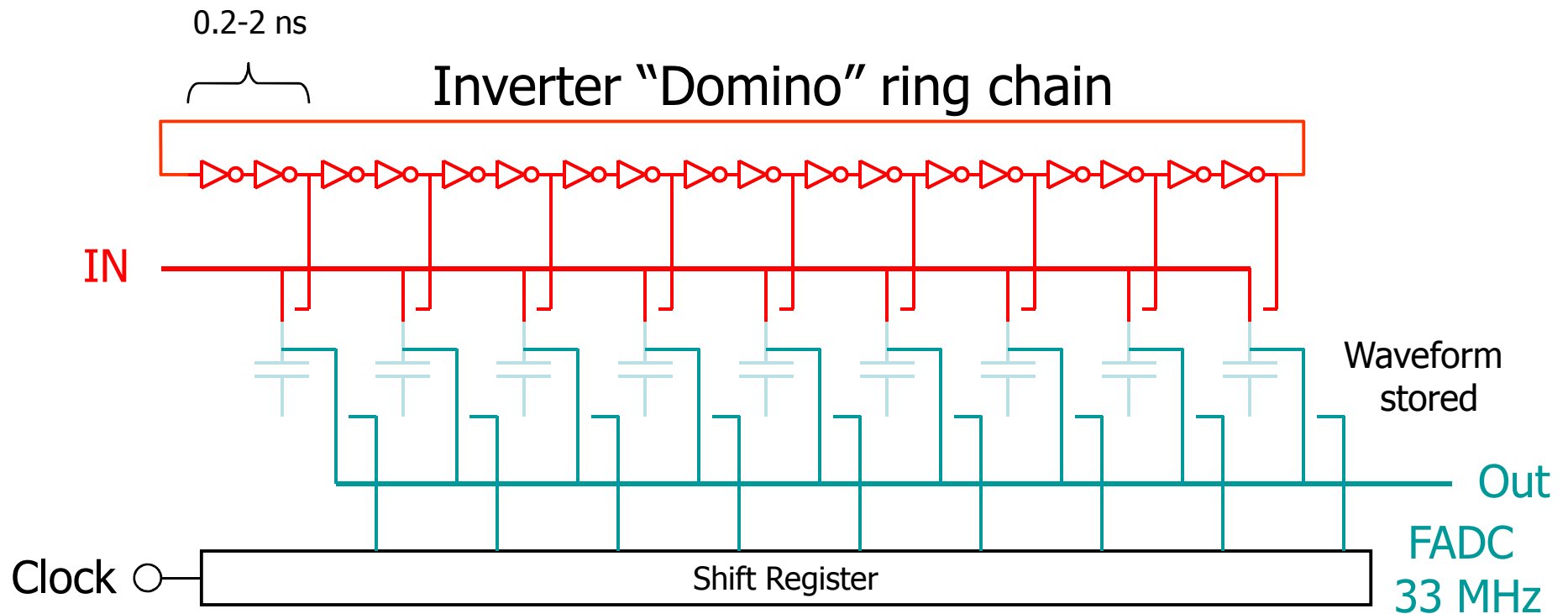
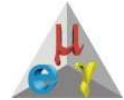


Problem: Pile-up



Problems:

- Power consumption
- Price
- Limited timing resolution



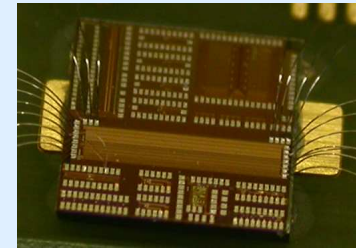
"Time stretcher" GHz → MHz

# The DRS Chip



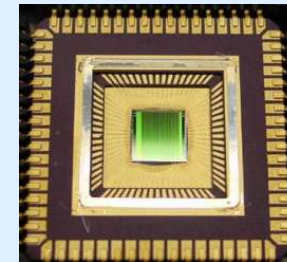
- Development of SCA chip based on experience in  $\pi\beta$  experiment
- Took four iterations to produce a flexible and powerful chip
- Goal was to design a chip which can be used in many experiments

2001



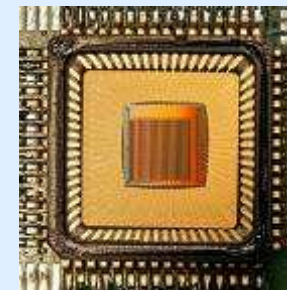
**DRS1**

2004



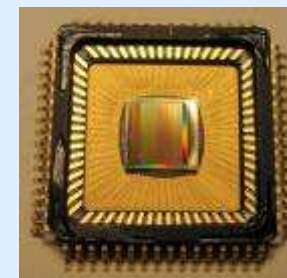
**DRS2**

2006



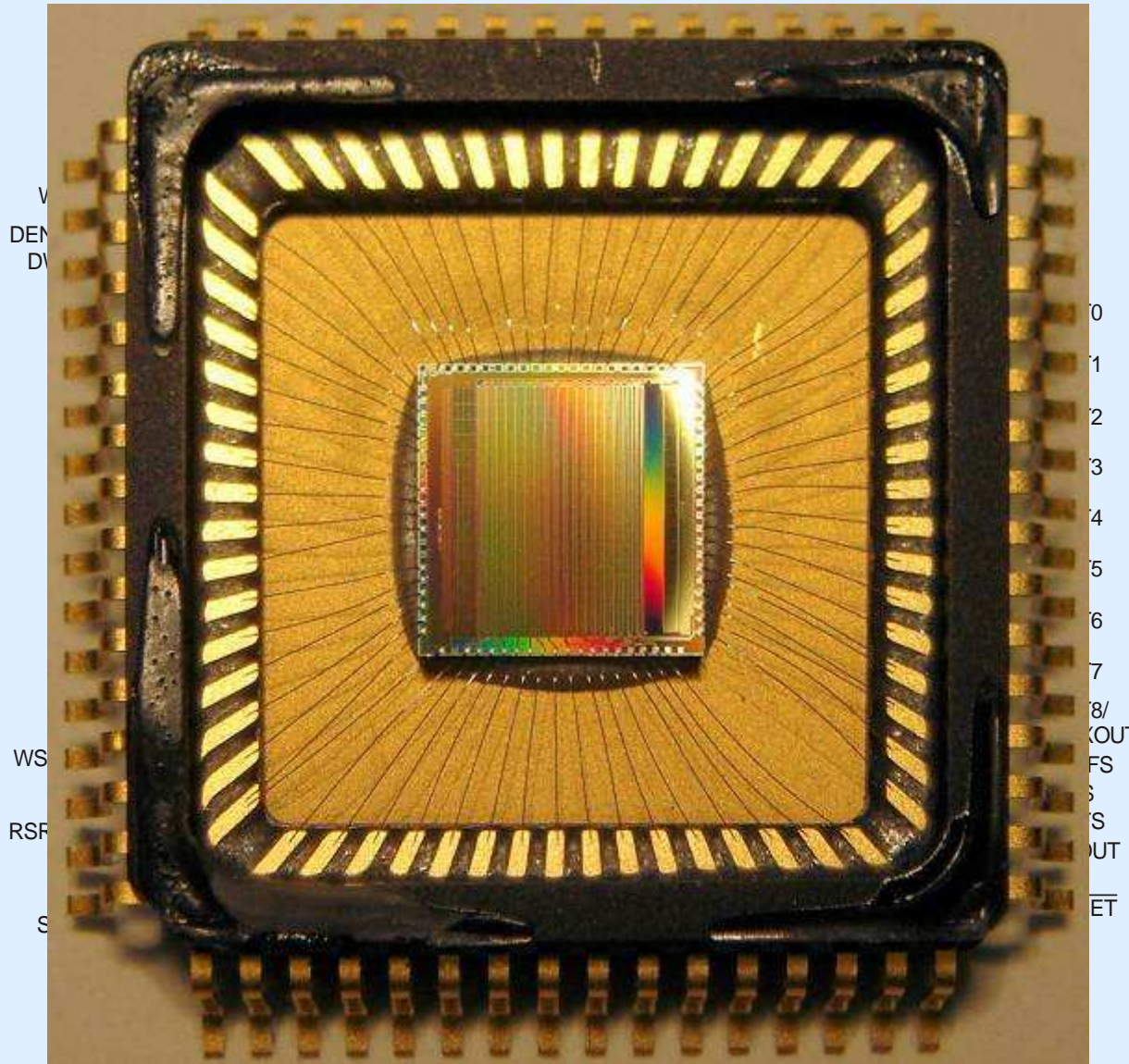
**DRS3**

2008

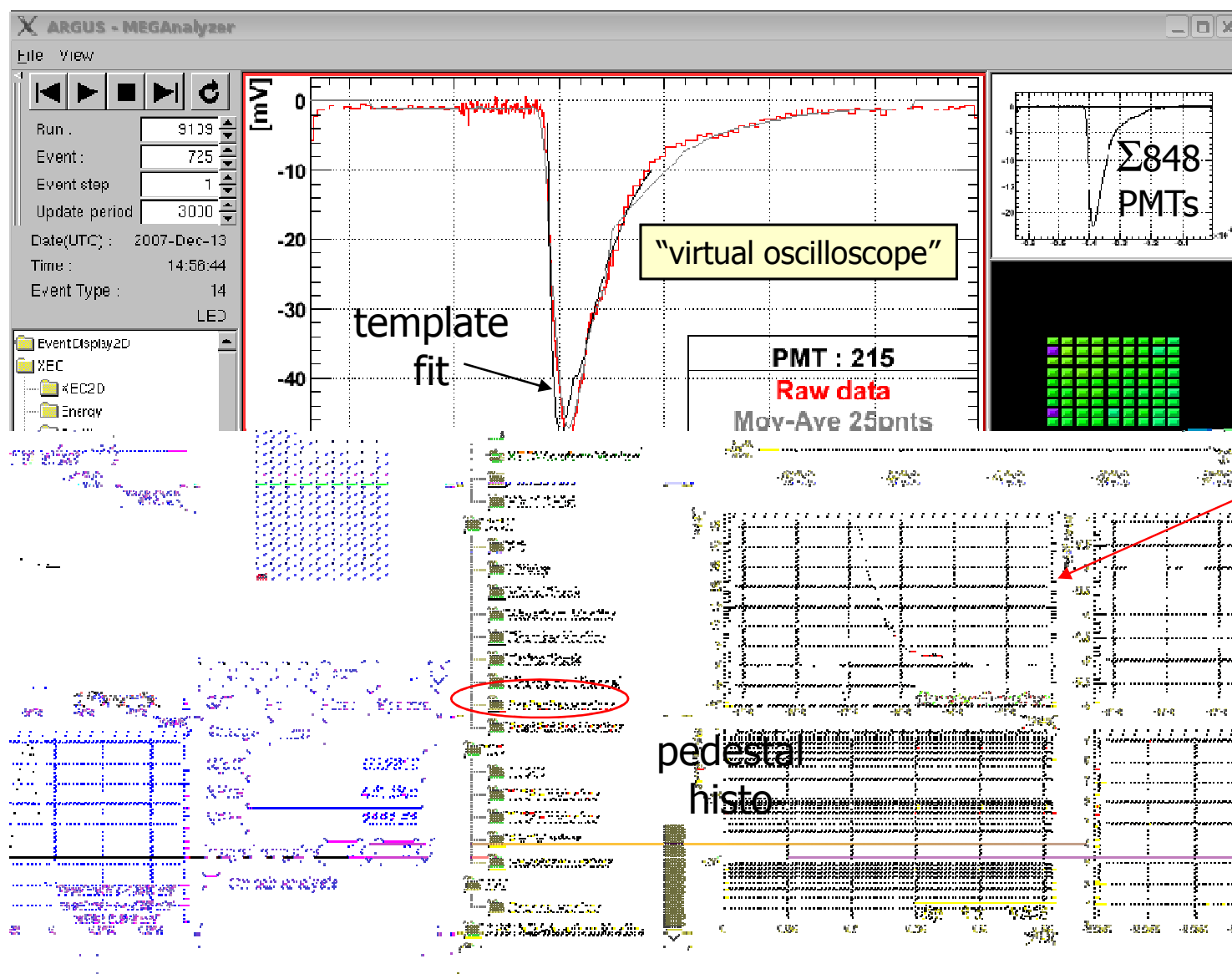


**DRS4**

- Fabricated in 0.25  $\mu\text{m}$  1P5M MMC process (UMC), 5 x 5 mm<sup>2</sup>, radiation hard
- 8+1 ch. each 1024 bins, 4 ch. 2048, ..., 1 ch. 8192
- Differential inputs/outputs
- Sampling speed 500 MHz ... 6 GHz
- On-chip PLL stabilization
- Readout speed 30 MHz, multiplexed or in parallel

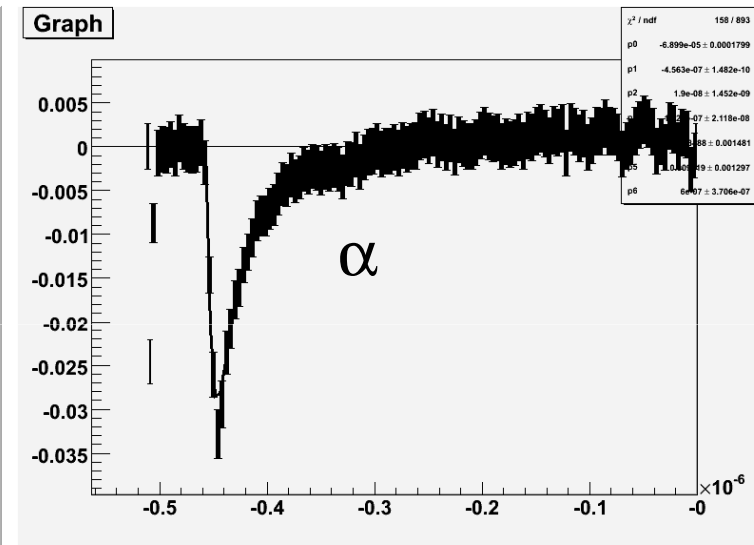
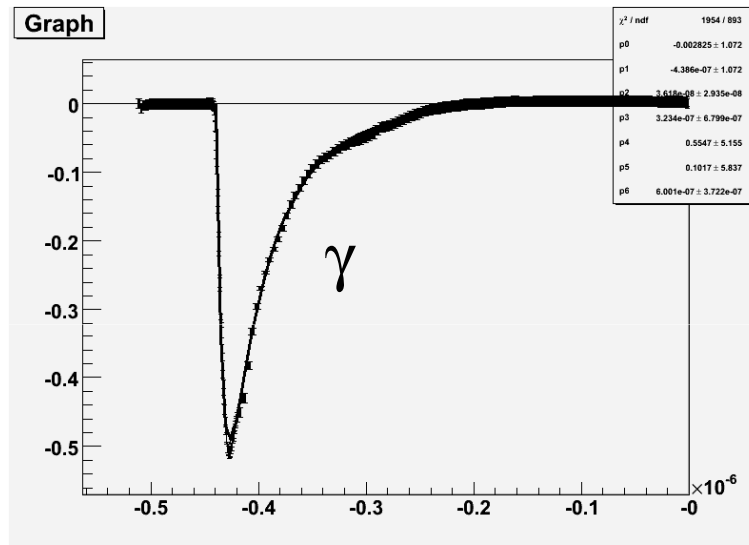


# On-line waveform display





Example:  $\alpha/\gamma$  source in liquid xenon detector (or:  $\gamma/p$  in air shower)



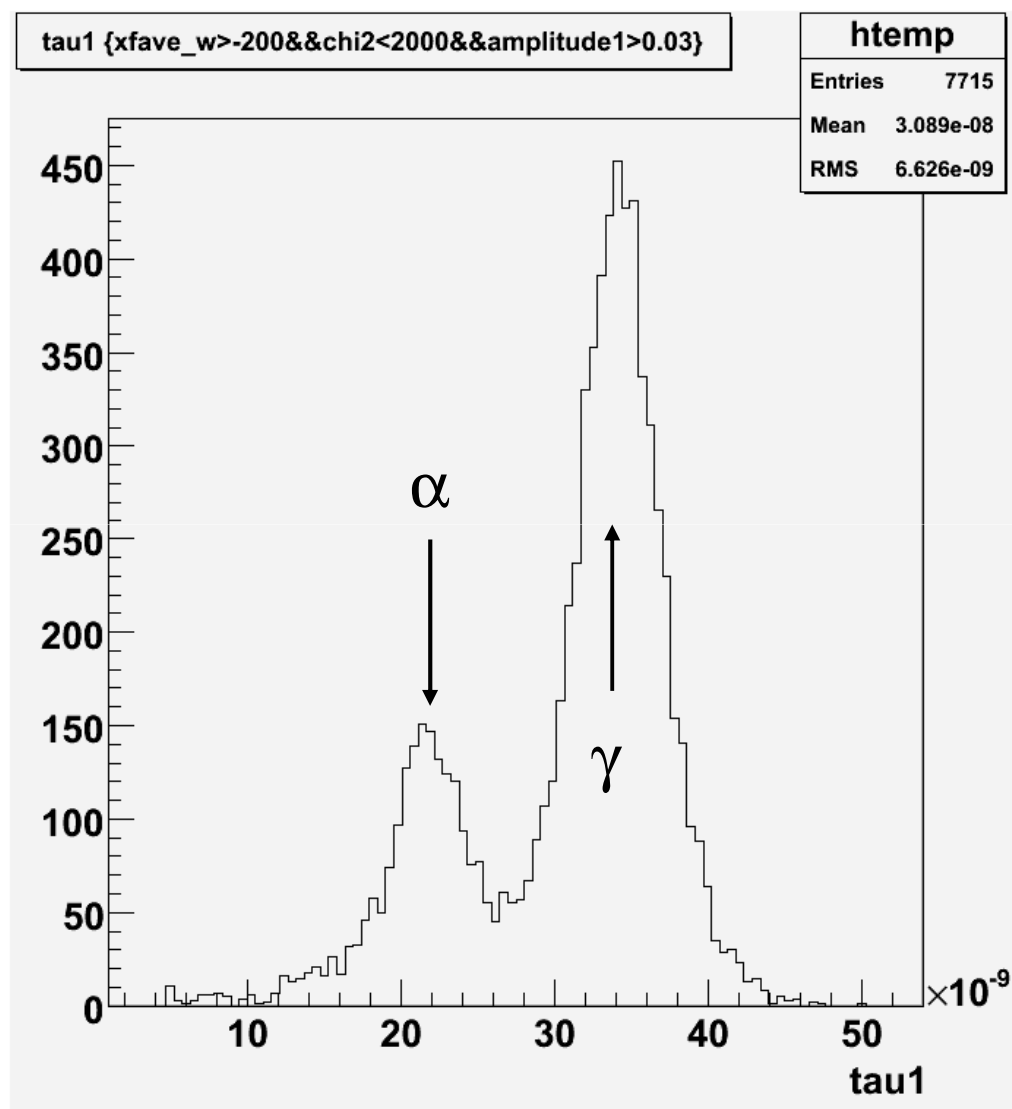
$$V(t) = \left[ A e^{-(t-t_0)/\tau_i} + B e^{-(t-t_0)/\tau_s} + C e^{-(t-t_0)/\tau_d} \right] \theta(t - t_0) + [\dots] \theta(t - t_r)$$

$\uparrow$   
Leading edge

$\uparrow$   
Decay time

$\uparrow$   
AC-coupling

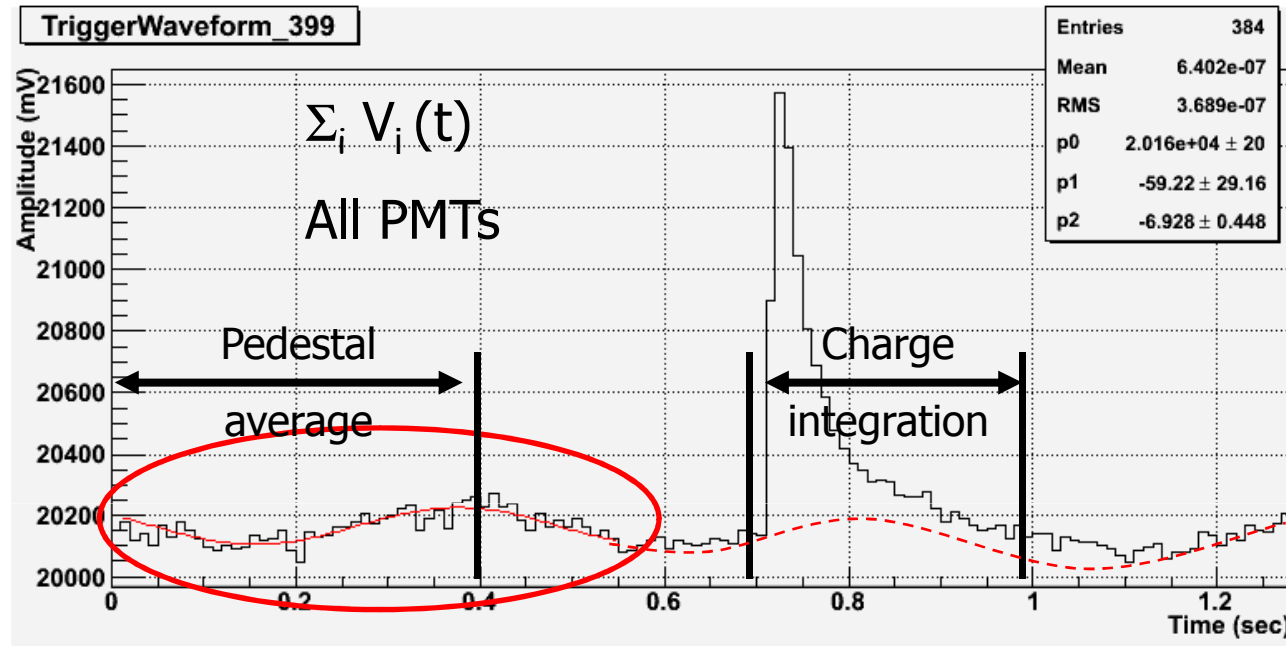
$\uparrow$   
Reflections



$$\tau_{\alpha} = 21 \text{ ns}$$

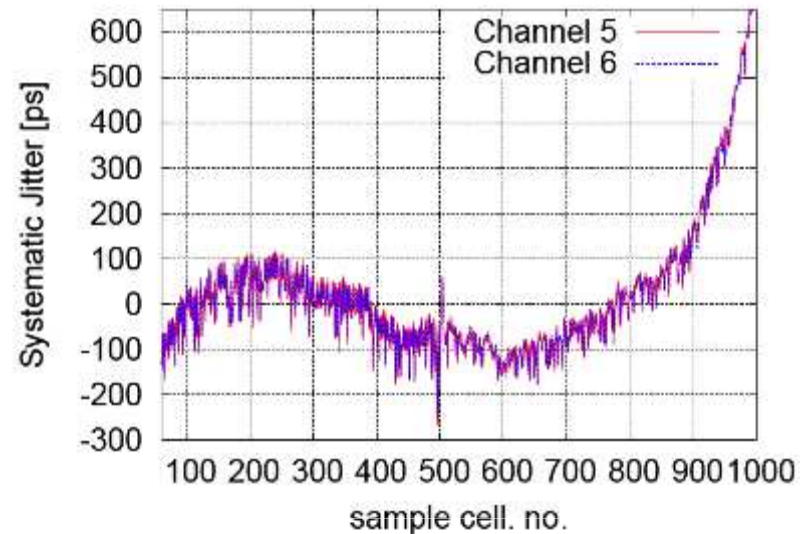
$$\tau_{\gamma} = 34 \text{ ns}$$

Waveforms can  
be clearly  
distinguished

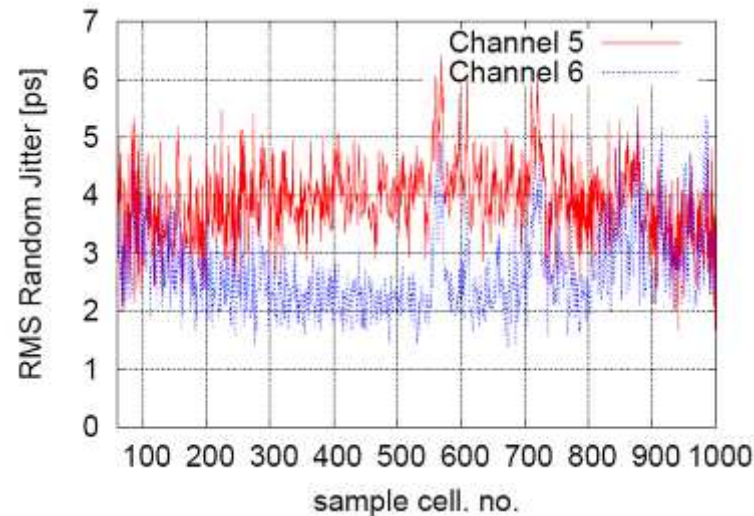


- Found some coherent low frequency ( $\sim$ MHz) noise
- Energy resolution dramatically improved by properly subtracting the sinusoidal background
- Usage of "dead" channels for baseline estimation

# Fixed Pattern Jitter Results



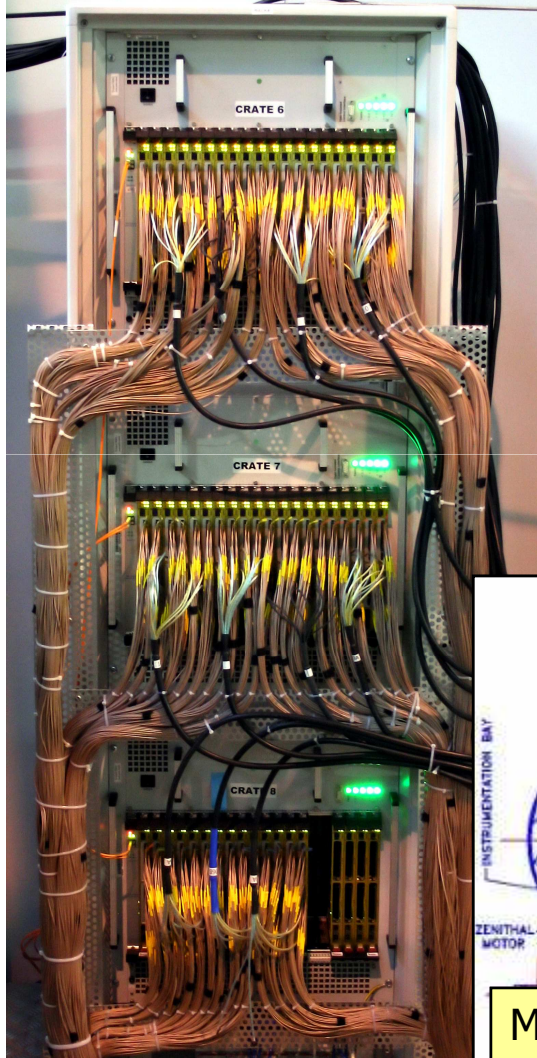
- $TD_i$  typically  $\sim 50$  ps RMS @ 5 GHz
- $TI_i$  goes up to  $\sim 600$  ps
- Jitter is mostly constant over time,  
→ measured and corrected
- Residual random jitter 3-4 ps RMS



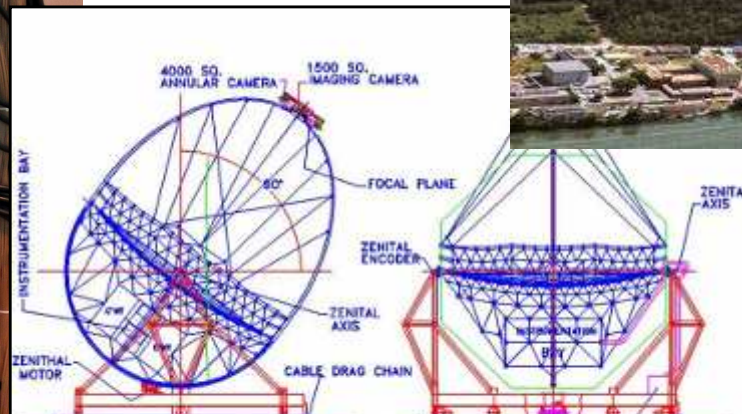
# Experiments using DRS chip



MEG 3000 channels DRS2



MAGIC-II 400 channels DRS2



BPM for XFEL@PSI  
1000 channels DRS4 (planned)

MACE (India) 400 channels DRS4 (planned)



- DRS4 can be obtained from PSI on a “non-profit” basis
  - Delivery “as-is”
  - Costs ~ 10-15 USD/channel
- USB Evaluation board as reference design
- VME boards from industry in 2009



32-channel  
65 MHz/12bit digitizer  
“boosted” by  
DRS4 chip to 5 GHz

## Challenge 5: Monitoring

How to keep the experiment stable for

## Challenge 6: Control

How to control 50 (IV, Temperatures, Pressures)?

## Challenge 7: Data Analysis

How to deal with 130 TB of data per year?

## Challenge 8: ...

For my next visit to UVA ...



# Status and Outlook

Where are we, where do we go?

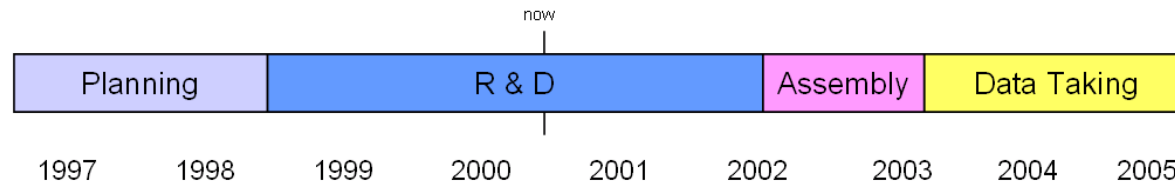
# Current Schedule



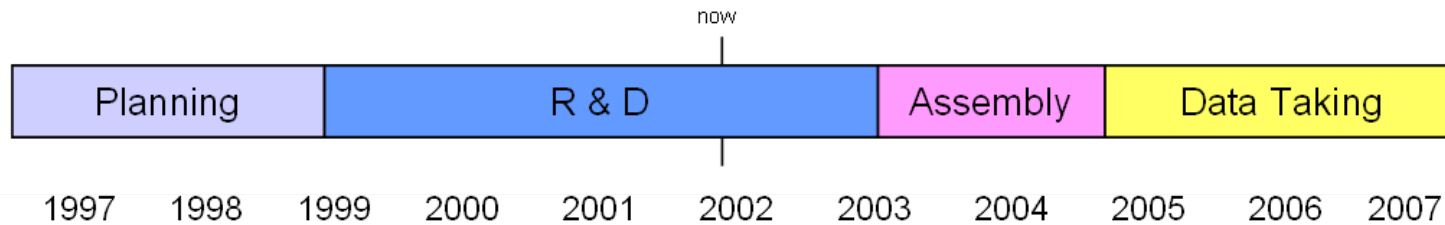
1999

We request a total of 6,000 hours of beam time starting in year 2003 with a proton intensity above 1 mA. Tests and engineering runs with  $\pi$ E5 will be requested in years 2000 to 2002. A

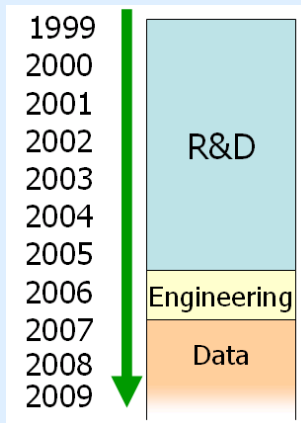
2000



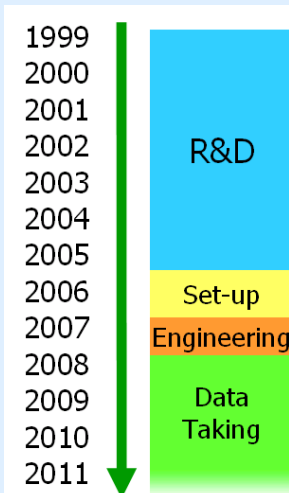
2002



2006



2008





CAUTION: All 2008 numbers are provisional

Still lots of things to learn from the data

- Blue numbers likely to change

- Grey numbers may vanish

## Efficiencies

(%)	"Goal"	2008 Provisional Lower Limits	2009 Provisional Prospects
Gamma	> 40	> <sup>depth</sup> 50 x ( <sup>pileup</sup> 65 x 85)	> 50 x 90
e+	65	<sup>DC</sup> 30 x <sup>DC-TC</sup> 40	85 x 50
Trigger	100	<sup>energy</sup> 100 x <sup>time</sup> 99 x <sup>direction</sup> 80	> 99
Selection	$90^4 = 66$	$90^3 \times 95 = 69$	69
DAQ	( > 90 )	<sup>live</sup> > 80 x <sup>run transition</sup> 93	> 90 x 99
Calibration Run etc	( > 95 )	~70	90
Running Time (week)	100*	11.5**	11.5
Single Event Sensitivity ( $10^{-13}$ )	0.5	< 30 - 50	< 3 - 5

\* 1 week =  $4 \times 10^5$  sec (66%)

\*\* CEX runs not included



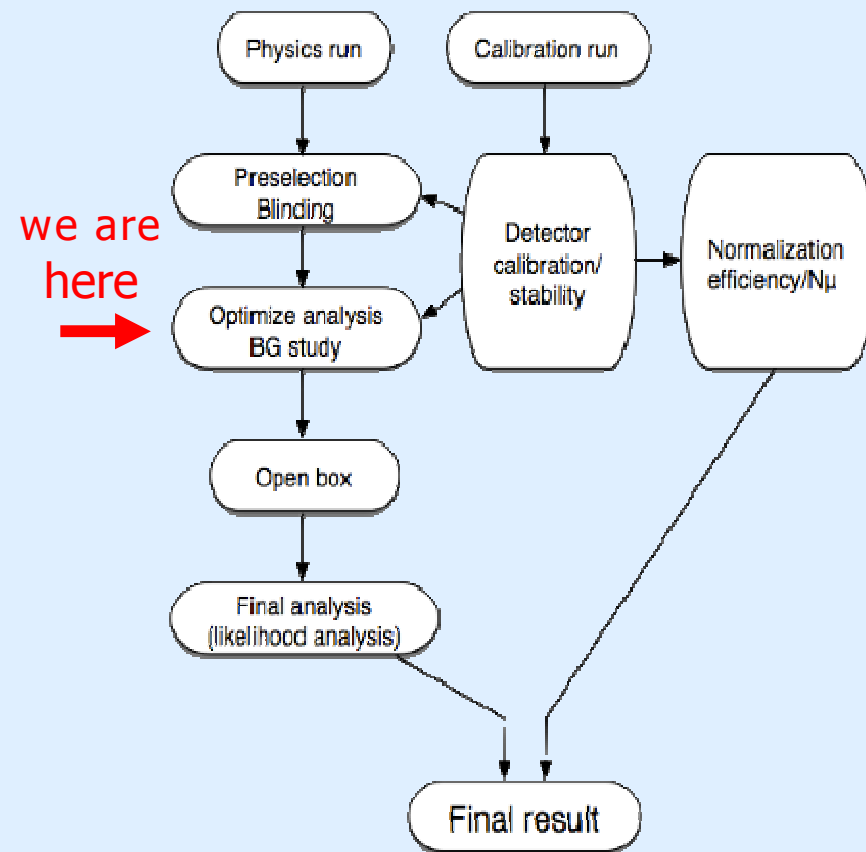
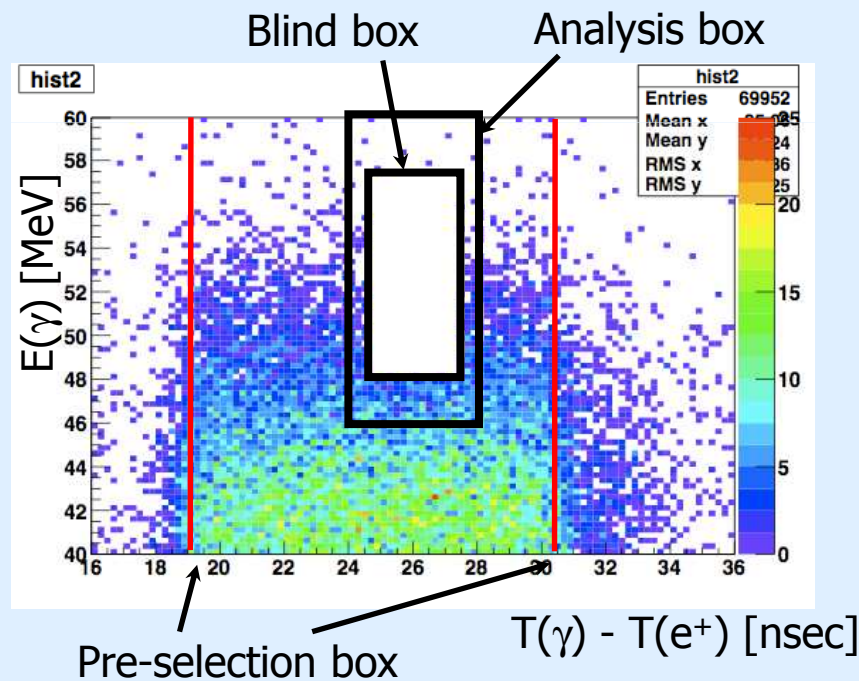
CAUTION: All 2008 numbers are provisional

## Resolutions

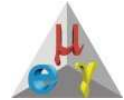
Resolutions are improving as we understand the detectors better.

(in sigma)	"Goal"	2008 Provisional	2009 Provisional Prospects
Gamma Energy (%)	1.2 - 1.5	< 2.3	< 1.7
Gamma Timing (ps)	65	< 100*	< 80
Gamma Position (mm)	2 - 4	5 - 6.5	5
e+ Momentum (%)	0.35	1.5 - 2.0	0.7 - 0.8
e+ Timing (ps)	45	< 60 - 90	60
e+ Angle (mrad)	4.5	9 - 18	11
mu Decay Point (mm)	0.9	3 - 4	2
Gamma - e+ Timing (ps)	80	150	100
Background ( $10^{-13}$ )	0.1 - 0.3	-	< 0.6 - 3

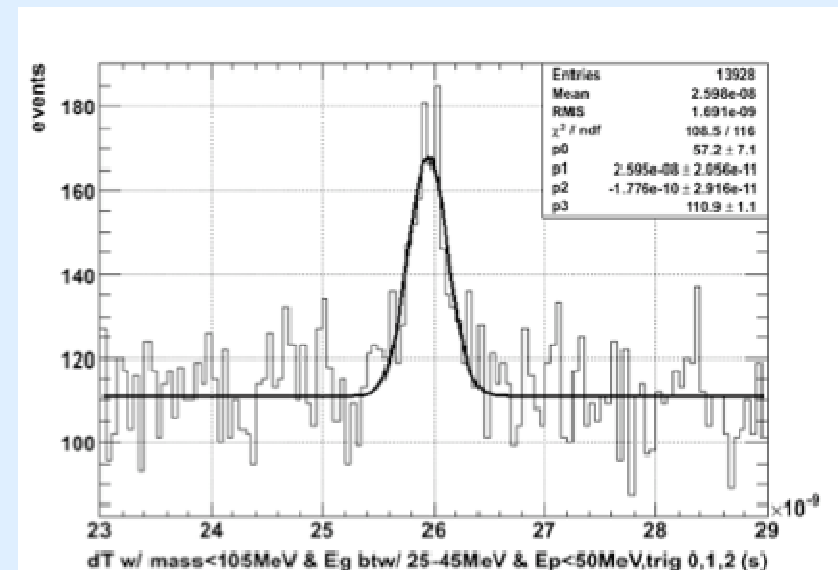
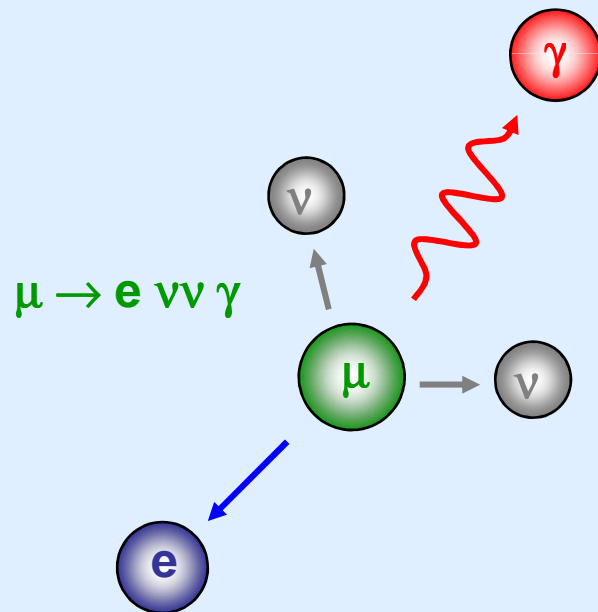
- 11.5 weeks of data taking in 2008 (130 TB)
- Currently doing blind analysis



# Radiative Muon Decay



- This decay is a benchmark for the whole detector
- Branching ratio 1.4%
- Decays clearly visible in high rate environment



$T(\gamma) - T(e^+)$

# "Polarized" MEG

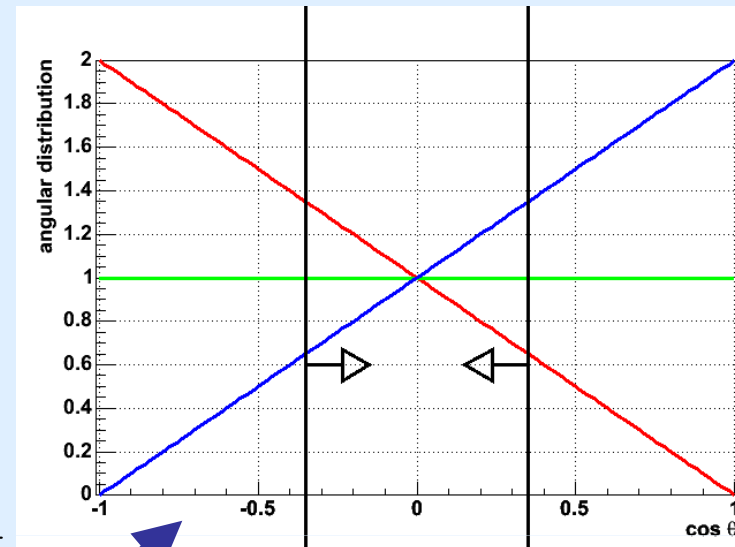


- $\mu$  are produced already polarized
- Different target to keep  $\mu$  polarization
- Angular distribution of decays predicted differently by different theories (compare Wu experiment for Parity Violation)

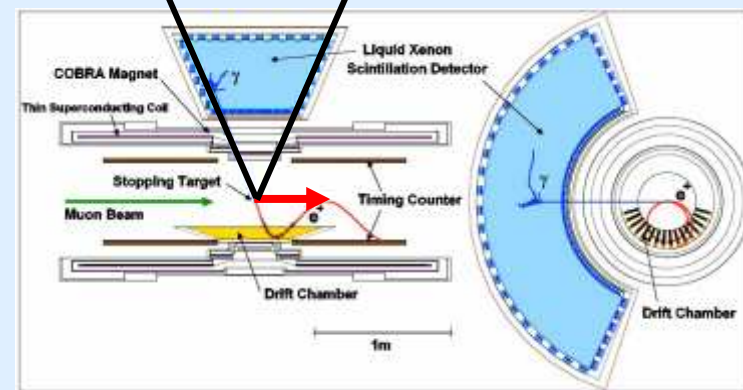
$$\frac{dN(\mu^+ \rightarrow e^+ \gamma)}{d \cos \theta_e} \propto BR(\mu^+ \rightarrow e^+ \gamma) \cdot \frac{1 + AP_\mu \cos \theta_e}{2}$$

SU(5) SUSY-GUT  
A = +1  
SO(10) SUSY-GUT  
A ≈ 0  
MSSM with  $\nu_R$   
A = -1

**Y.Kuno *et al.*,  
Phys.Rev.Lett. 77 (1996) 434**



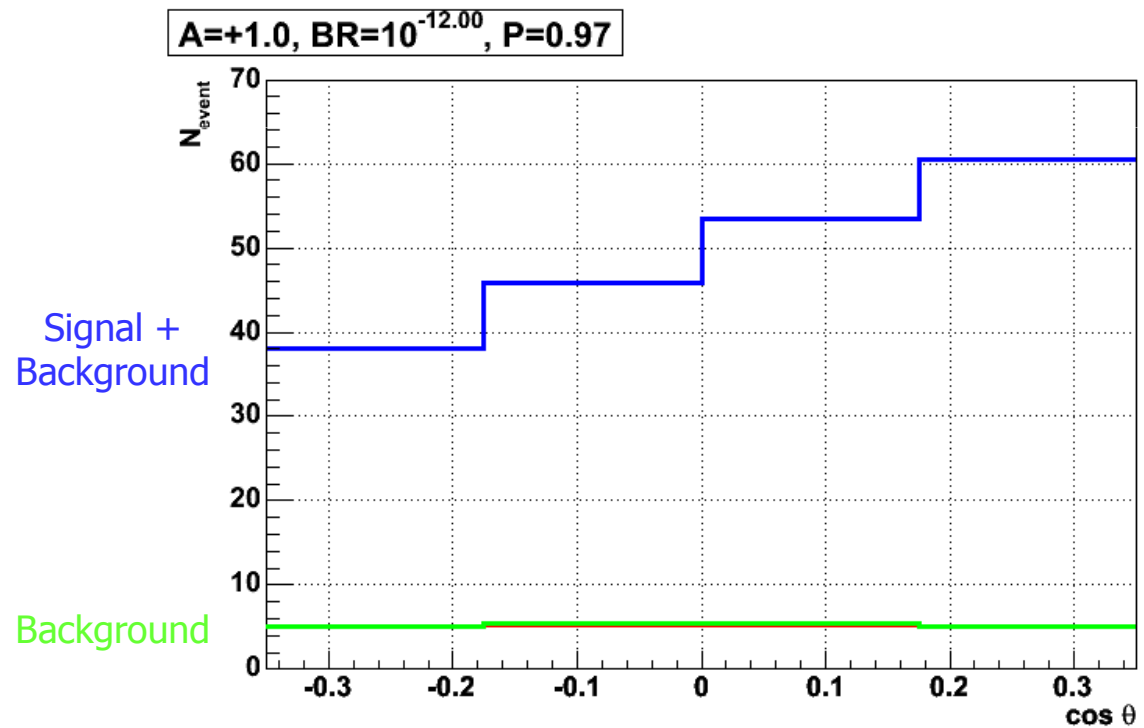
Detector acceptance



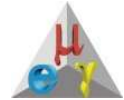
# Expected Distribution



- $A = +1$
- $B(\mu^+ \rightarrow e^+ \gamma) = 1 \times 10^{-12}$
- $1 \times 10^8 \mu^+/\text{s}$
- $5 \times 10^7 \text{ s}$  beam time (2 years)
- $P_\mu = 0.97$



**S. Yamada**  
@ SUSY 2004, Tsukuba



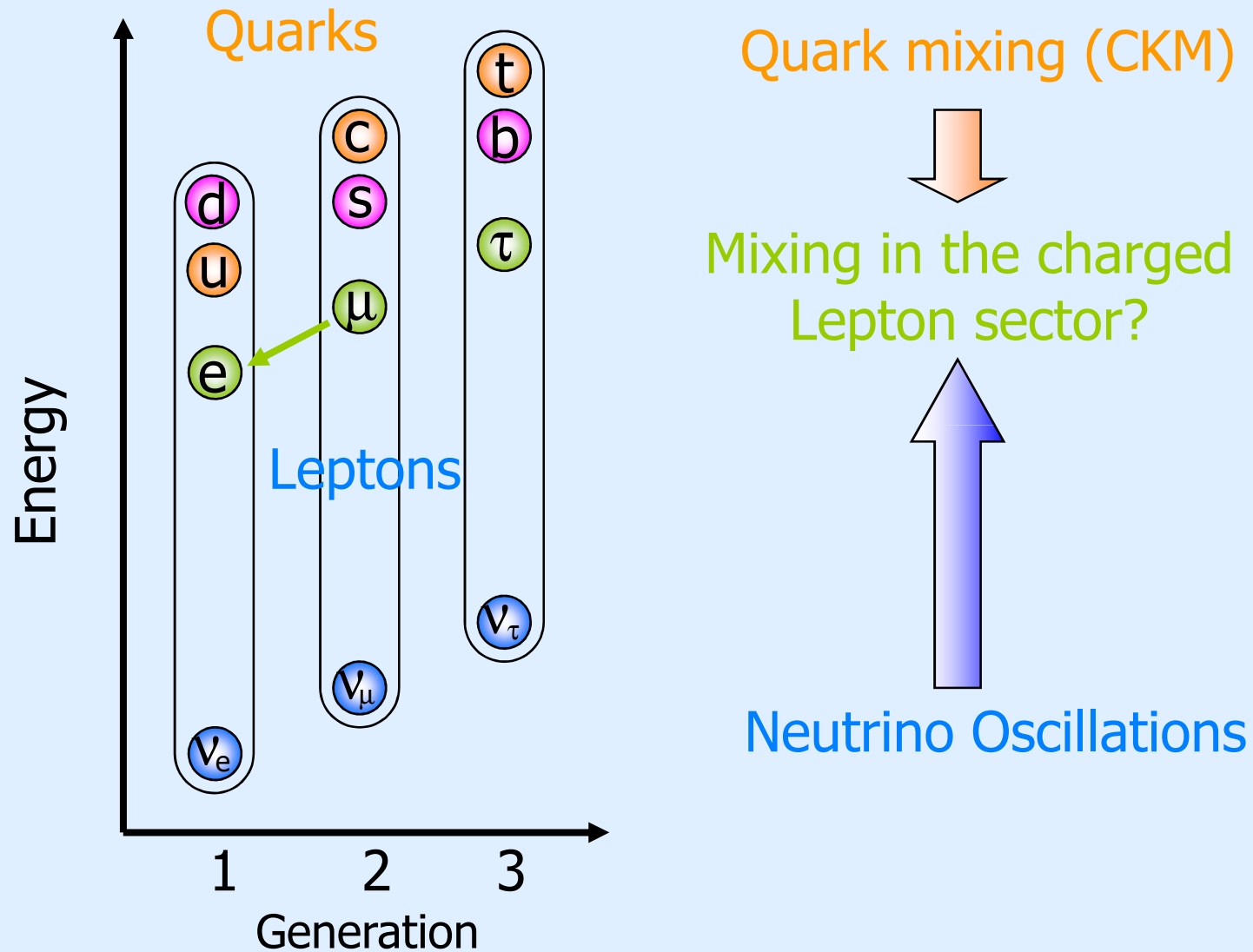
- Many challenges faced in the MEG Experiment, solutions have been worked out
- Some technologies might be interesting for other experiment
  - Liquid Xenon Calorimetry
  - Fast Waveform Digitizing using the DRS chip
- MEG just started taking data, so expect exciting results in the upcoming years



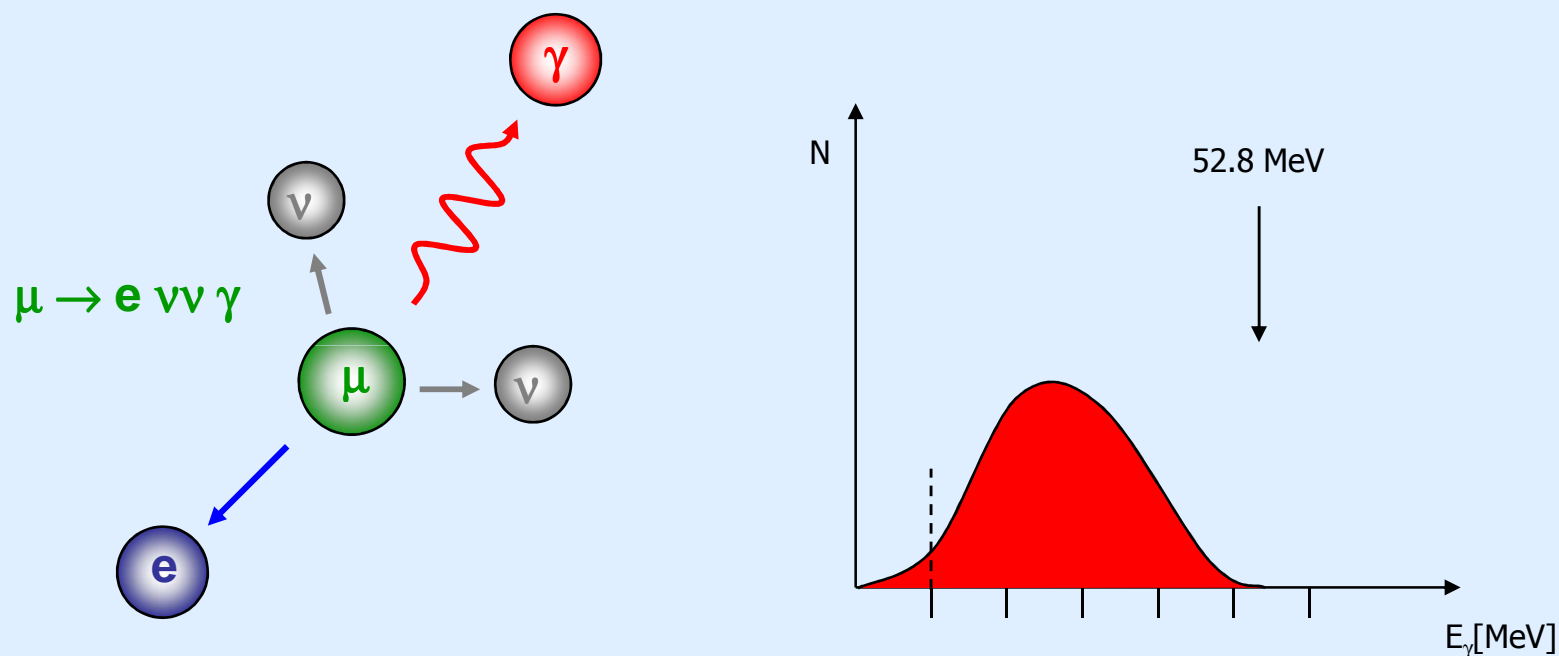
<http://meg.psi.ch>

Backup Slides

# Mixing of Generations

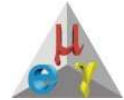


# Radiative Muon Decay (1.4%)

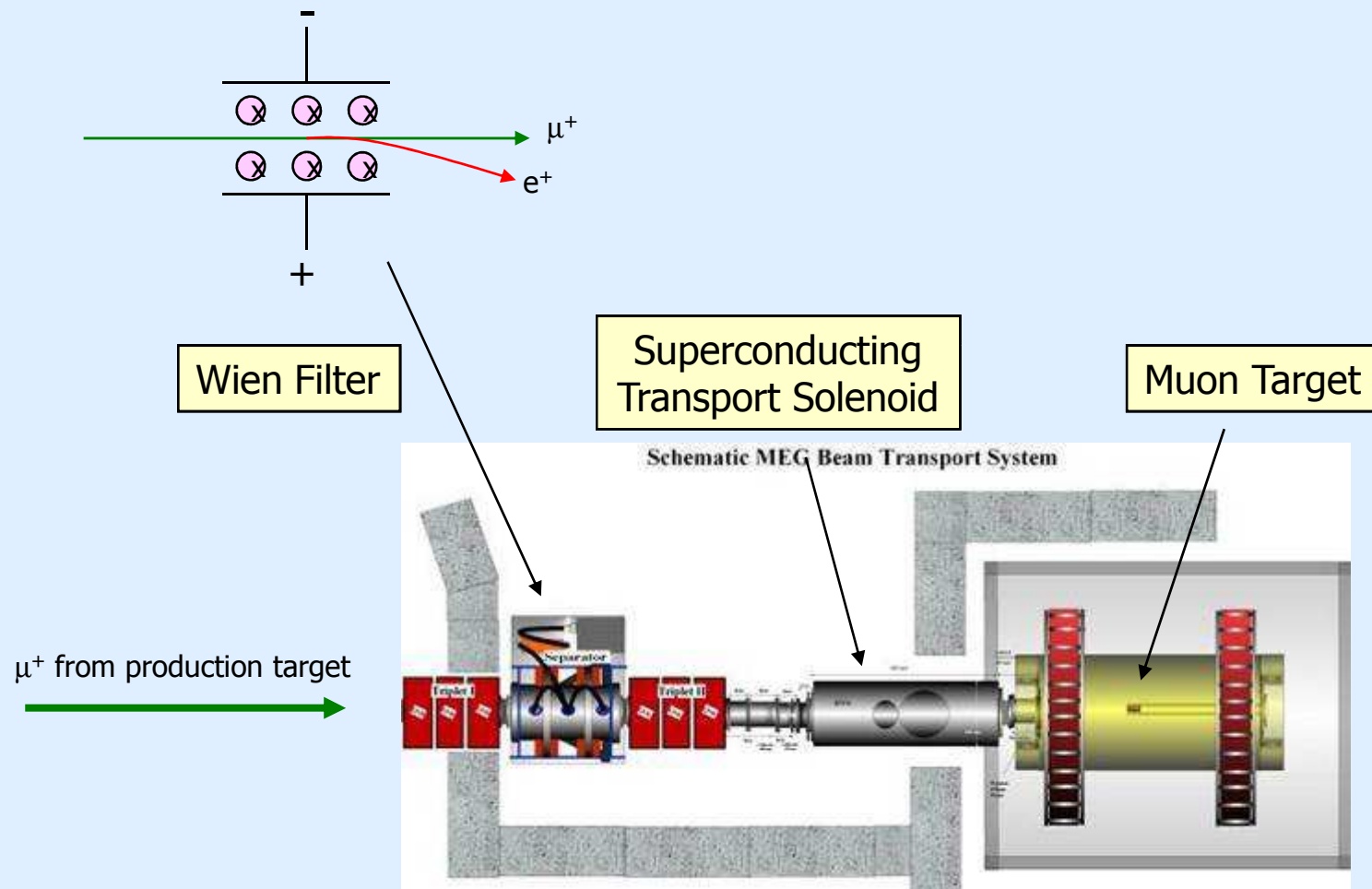


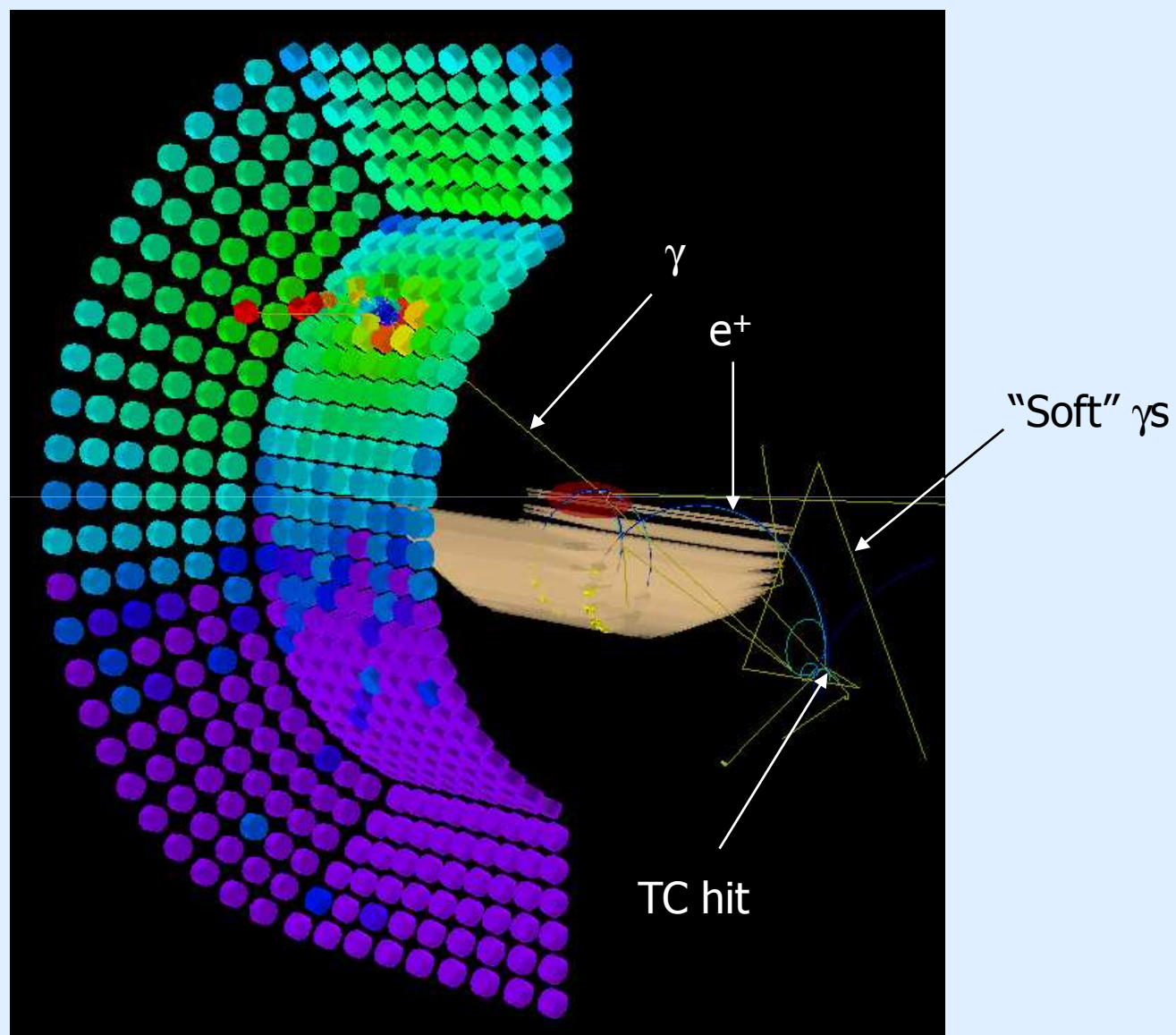
"Prompt" Background:  $< 10^{-17}$

# Muon Beam Line



Transport  $10^8 \mu^+$ /s to stopping target inside detector with minimal background

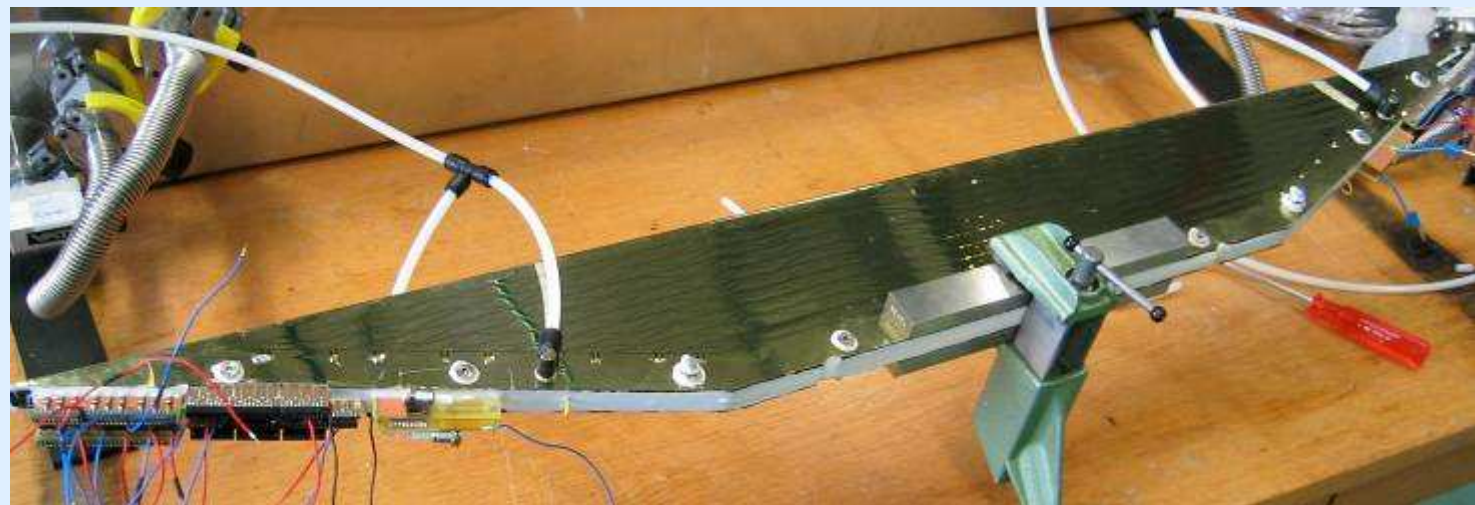
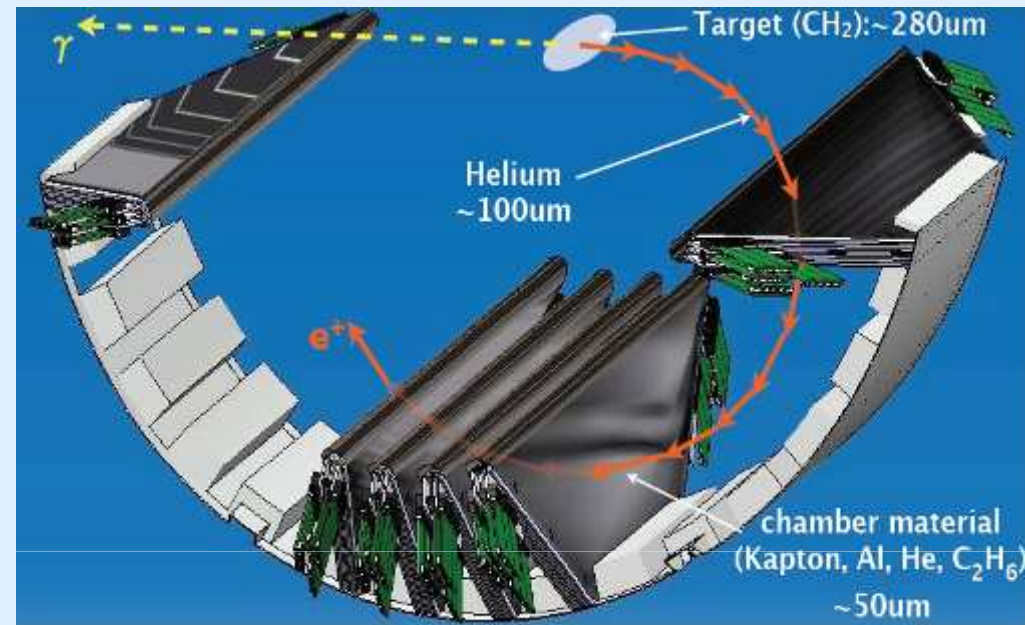




# Positron Detection System



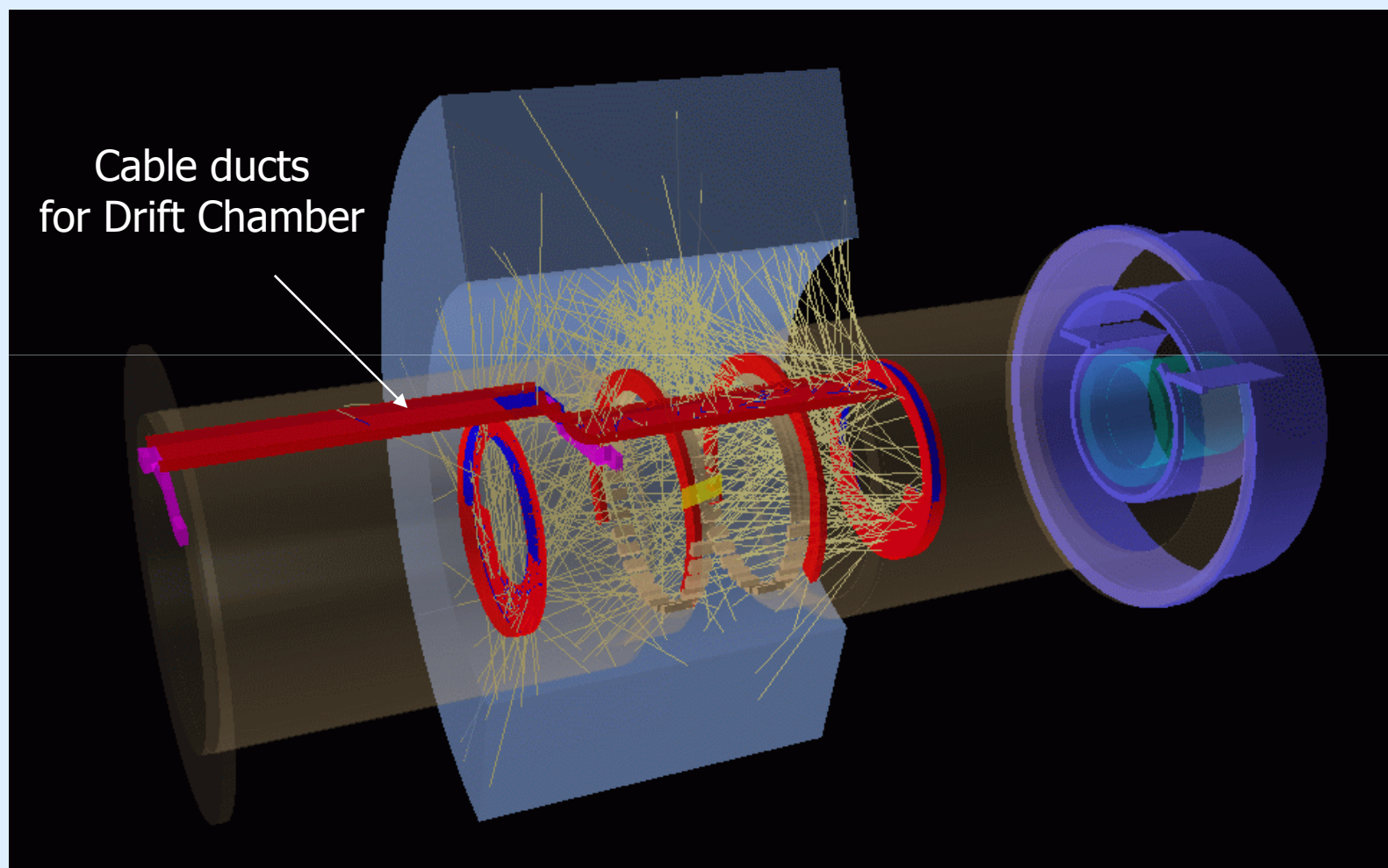
- 16 radial DCs to measure positron tracks
- Extremely low mass
- He:C<sub>2</sub>H<sub>6</sub> gas mixture
- Scintillation counter for precise timing

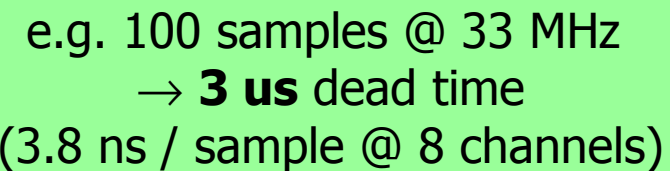


# Beam induced background

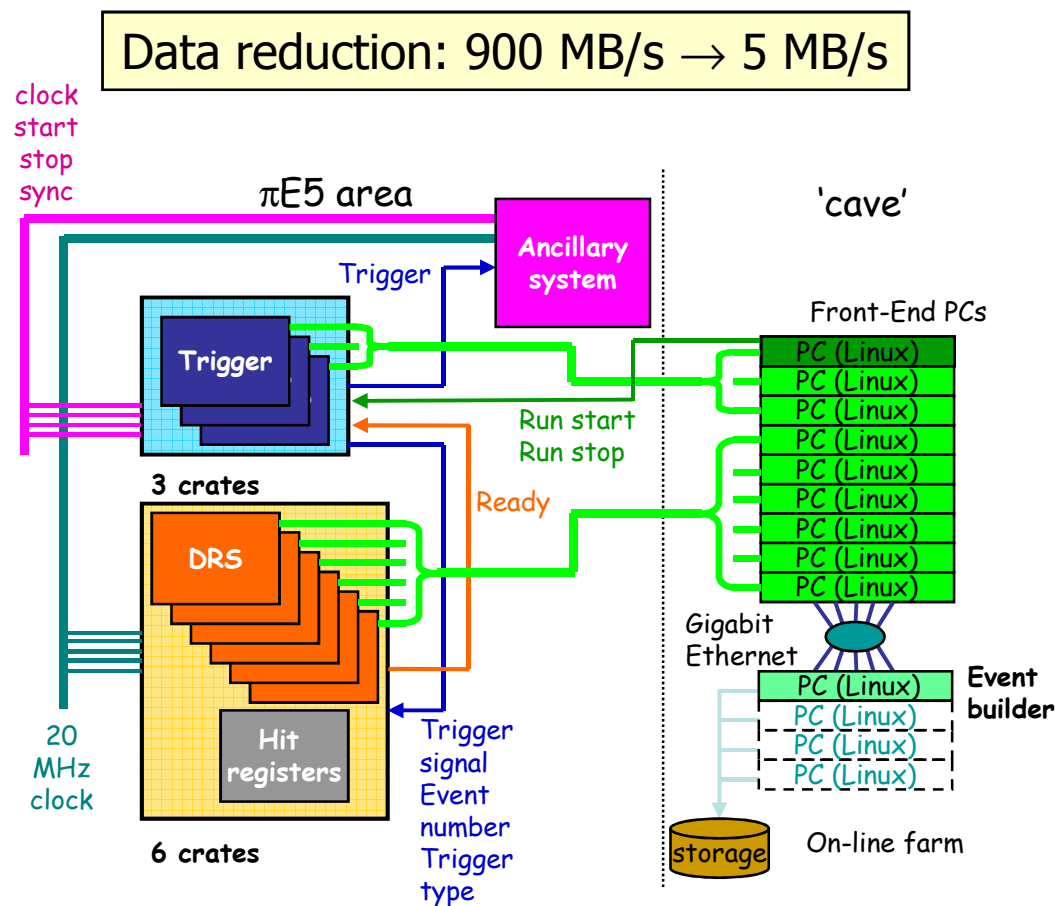
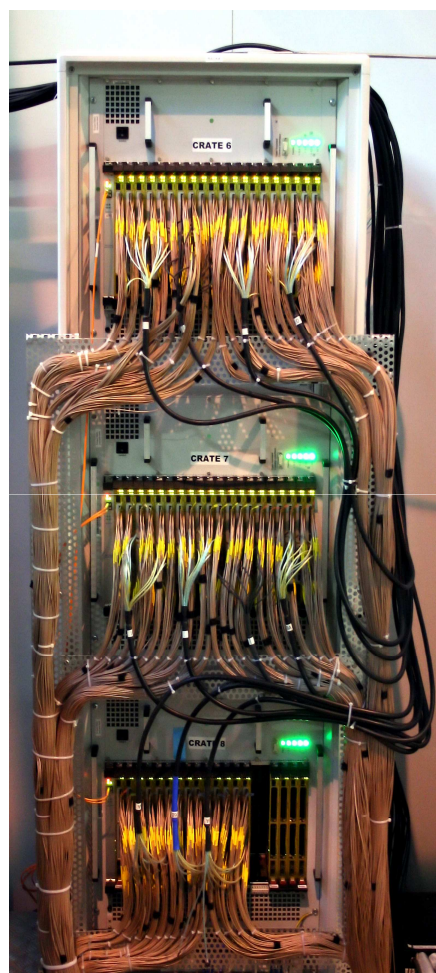


$10^8 \mu/s$  produce  $10^8 e^+/s$  produce  $10^8 \gamma/s$

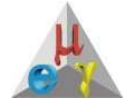




# Complete DAQ System



# What next?



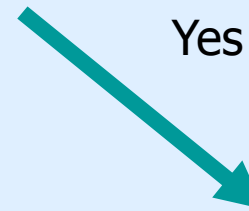
Will we find  
 $\mu \rightarrow e \gamma$ ?

No



- Improve experiment from  $10^{-13}$  to  $10^{-14}$ :
  - Denser PMTs
  - Second Calorimeter

Yes



- Carefully check results
- Be happy ☺
- Result must be combined with other experiments:
  - $\mu \rightarrow e$  conversion
  - $\mu \rightarrow eee$